<u>Xenon High-pressure Gas</u> TPCs for $0-\nu \beta\beta$ & WIMP Searches, ...

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Why bother with high-pressure?

- Neutrino-less double beta decay:
 - Energy resolution may be an order of magnitude better than in liquid xenon ...
 - Tracking in gas phase reveals event topology...
- Direct detection WIMP search:
 - S2/S1 ratio shows better discrimination than LXe
 - Gas phase may permit extremely low threshold
- Flexibility: molecular additives, nuclear safeguards

Goals for 0-v $\beta\beta$ search

- Near-term: demonstrate energy resolution and tracking at 20 bars
 - Reconfigure, operate small LLNL TPC
 - Design & construct medium-scale TPC
- Longer-term:
 - If performance goals are met, a design for a serious ton-scale experiment can be prepared

To search for $0-\nu \beta\beta$ decay: 1. Acquire 100 - 1000 kg of candidate nuclei 2. Measure the two electron energies, <1% FWHM

3. Reject backgrounds!



Summed electron energy in units of the kinematic endpoint (Q)

Energy Resolution: <u>CUORE</u> $\delta E/E = 3 \times 10^{-3} FWHM$!



Xenon: Strong dependence of energy resolution on density!

A. Bolotnikov, B. Ramsey / Nucl. Instr. and Meth. in Phys. Res. A 396 (1997) 360-370



Fig. 5. Density dependencies of the intrinsic energy resolution (%FWHM) measured for 662 keV gamma-rays.

For ρ <0.55 g/cm³, ionization energy resolution is "intrinsic"

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Energy Partitioning in Xenon

Anomalously large fluctuations exist in partitioning of energy in LXe Fluctuations cause "anti-correlation" in LXe WIMP search: very bad news for S2/S1 resolution!

Only a small fraction of scintillation signal is recoverable Energy resolution cannot be restored in LXe

Anomalous fluctuations do not exist in HPXe A measurement of ionization <u>alone</u> is sufficient to obtain **near-intrinsic energy** resolution...

Intrinsic energy resolution

• $\delta E/E = 2.35 \cdot (F \cdot W/Q)^{1/2}$

- $F \equiv$ Fano factor: F = 0.15 (HPXe) (LXe: F ~20)
- W = Average energy per ion pair: W ~ 25 eV
- Q = Energy release in decay of ¹³⁶Xe: ~2500 keV $\delta E/E = 2.8 \times 10^{-3} FWHM$ (HPXe)

N = Q/W ~100,000 primary electrons $\sigma_N = (F \cdot N)^{1/2} ~120$ electrons rms! Need gain with very low noise/fluctuations! Avalanche gain <u>cannot</u> meet this objective

Fluctuations in Electroluminescence (EL)

EL: linear gain process: electrons excite, atoms radiate Uncorrelated fluctuations: $\sigma = ((F + G) \cdot N)^{1/2}$

EL can provide G less than F:

- **G** for EL contains three terms:
 - 1. Fluctuations in **n**_{uv} (UV photons per e):
 - 2. Fluctuations in n_{pe} (detected photons/e):
 - 3. Fluctuations in photo-detector single PE response:

$$\mathbf{G} = \sigma^2 = 1/(\mathbf{n}_{uv}) + (\mathbf{1} + \sigma^2_{pmt})/\mathbf{n}_{pe})$$

For $G \le F = 0.15 \implies n_{pe} \ge 10/e \implies \Sigma n_{pe} \ge 10^6$ @Q-value Equivalent noise: <u>much less</u> than 1 electron rms!

Separated-function TPC:



Topology: "spaghetti, with meatballs"



Slide: NEXT collaboration

Various HPXe efforts

"Gotthard" TPC: first xenon 0-ν ββ experiment Coimbra: (Portugal) electroluminescence Russia: several pioneering achievements Beppo-SAX satellite: 7-PMT 5-bar TPC Texas A&M: 7-PMT 20 bar HPXe TPC BNL-Temple: HPXe scintillation decay time EXO - gas: Ba⁺⁺ ion tagging, tracking, ... LBNL/LLNL/TAMU/NEXT...

EL in 4.5 bar of Xenon (Russia - 1997)



Fig. 1. Schematic diagram of the gas scintillation drift chamber with 19 PMT matrix readout.





Europe: Beppo-SAX satellite: a HPXe TPC in space!



7-PMT, 20 bar TAMU HPXe TPC







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Europe:

NEXT collaboration

Spain/Portugal/France...

funded: 5M €!

to develop & construct a 100 kg enriched HPXe TPC for $0-\nu \beta\beta$ decay search at Canfranc Laboratory within 5 years

The NEXT Collaboration

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Medium-scale TPC @ LBNL

Separated function TPC

- Diameter of active volume: 30 cm
- Length: 60 cm (drift length: ~30 cm)
- -20 bars pure xenon (M_{Xenon} ≤ 2.5 kg)
- -28 mm PMTs detect t₀, measure energy, and track
- $N_{PMT} = 61$ per end (BaBar PMTs requested)
- Scale matched to 2500 keV Q-value of ¹³⁶Xe
- AMANDA ADCs requested for signal capture



Medium-scale EL TPC

- Primary Goal #1: Energy resolution
 - $\delta E/E \le 5 \ge 10^{-3}$ FWHM at Q-value (2480 keV)

- Must be demonstrated at MeV energies!

• Primary Goal #2: <u>3 -D tracking</u>

Multiple scattering ⇒ complex topologies
 Verify meatball recognition efficiency

• Other Goals: nuclear/electron recoil discrimination, Compton imaging, ...

FY10 Scenario

- LBNL, TAMU, LLNL, NEXT + ... collaborate
- "Xenophilia" workshop, November 2009:
 how does the community want to proceed?
- small-scale TPC operation + simulations:
 experience + optimized design of m-s TPC
- medium-scale TPC built and commissioned
 initial operation at/near ground level
- study of performance & design issues at ton scale
 radiopurity, background rejection, gas handling, ...

FY11 - FY12 Scenario

- Transition to operation at modest depth
- Data analysis, technical refinements +1-2 years
- Performance justifies serious 1000+ kg system
- Substantial new collaboration formed...

HPXe TPC Summary

- Potential for superior performance is high
 x11 better intrinsic energy resolution than LXe
- Performance of HPXe needs demonstration
 EL scheme works better in larger systems
- Simultaneous WIMP/ $0\nu\beta\beta$ search possible - No compromise necessary for either purpose









CUORE: Cryogenic "calorimeters"

- CUORICINO: 40.7kg TeO₂ (34% abundant ¹³⁰Te)
 - $T^{0v}_{1/2} \ge 2.4 \times 10^{24} \text{ yr} (90\% \text{ C.L.})$
 - $< m_v > \le 0.2 0.9 \text{ eV}$
 - Resolution: 7.5 keV FWHM at Q = 2529 keV!
- CUORE ~1000 crystals, 720 kg



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EXO-200: expected E resolution

ionization and scintillation are strongly "anti-correlated" in LXe



 $\delta E/E = 33 \times 10^{-3} @ Q_{0\nu\beta\beta} FWHM$ - predicted

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Backgrounds for the $\beta\beta0\nu$ search



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NEXT Collaboration

Nr Discrimination in HPXe with TAMU 7-PMT TPC



neutrons

Why Xenon for $0-\nu \beta\beta$ search?

- Only inert gas with a $0-\nu \beta\beta$ candidate
- No long-lived Xe radio-isotopes
- No need to grow crystals no surfaces
- Can be easily re-purified in place (recirculation)
- ¹³⁶Xe enrichment easy (natural abundance 8.9%)
- Gas Phase advantages:
 - Purification easier
 - No liquid level or temperature challenges
 - Event topology available
 - Excellent energy resolution (not demonstrated!)

"Gotthard TPC"

Pioneer TPC detector for 0-v $\beta\beta$ decay search

- Pressurized TPC, to 5 bars
- Enriched 136 Xe (3.3 kg) + 4% CH₄
- MWPC readout plane, wires ganged for energy
- No scintillation detection \Rightarrow no TPC start signal!
 - No measurement of drift distance
- $\delta E/E \sim 80 \text{ x } 10^{-3} \text{ FWHM} (1592 \text{ keV})$

 \Rightarrow 66 x 10⁻³ FWHM (2480 keV)

Reasons for this less-than-optimum resolution are not clear...

Possible: uncorrectable losses to electronegative impurities

Possible: undetectable losses to quenching $(4\% \text{ CH}_4)$

But: ~30x topological rejection of γ interactions!

Silicon Photomultiplier "SiPM"



Figure 1 Schematic of a single microcell (left), schematic of part of an SPM array of microcells (center) and photo of a portion of the SPM microcells (right).

SiPM from Hamamatsu, "MPPC"



SiPM photoelectron spectrum



Electro-Luminescence (EL) is the key (Gas Proportional Scintillation)

- Electrons drift in low electric field region
- Electrons then enter a high electric field region
- Electrons gain energy, excite xenon, lose energy
- Xenon generates VUV (175 nm)
- Electron starts over, gaining energy again
- Linear growth of signal with voltage
- Photon generation up to ~1000/e, but <u>no</u> ionization
- Early history irrelevant, \Rightarrow **fluctuations are small**

More Virtues of Electroluminescence

- Immune to <u>microphonics</u>
- Absence of positive ion <u>space charge</u>
- <u>Linearity</u> of gain versus pressure, HV
- <u>Isotropic</u> signal dispersion in space
- Trigger, energy, and tracking functions accomplished with <u>optical</u> detectors

Molecular Chemistry of Xenon

- Scintillation:
 - Excimer formation: $Xe^{*} + Xe \rightarrow Xe_{2}^{*} \rightarrow hv + Xe$
 - Recombination: $Xe^+ + e^- \rightarrow Xe^* \rightarrow$
- Density-dependent processes also exist:

$$Xe^{*} + Xe^{*} \rightarrow Xe^{**} \rightarrow Xe^{+} + e^{-} + heat$$

- Two excimers are consumed!
- More likely for both high ρ + high ionization density
- Quenching of both ionization and scintillation can occur! $Xe^* + M \rightarrow Xe + M^* \rightarrow Xe + M + heat (similarly for Xe_2^*, Xe^{**}, Xe_2^{*+}...)$ $Xe^+ + e^-(hot) + M \rightarrow Xe^+ + e^-(cold) + M^* \rightarrow$

$$Xe^+ + e^-(cold) + M + heat \rightarrow e^-(cold) + Xe^+ \rightarrow Xe^*$$