



Search for Charged Lepton Flavor Violation at J-PARC - COMET Experiment -

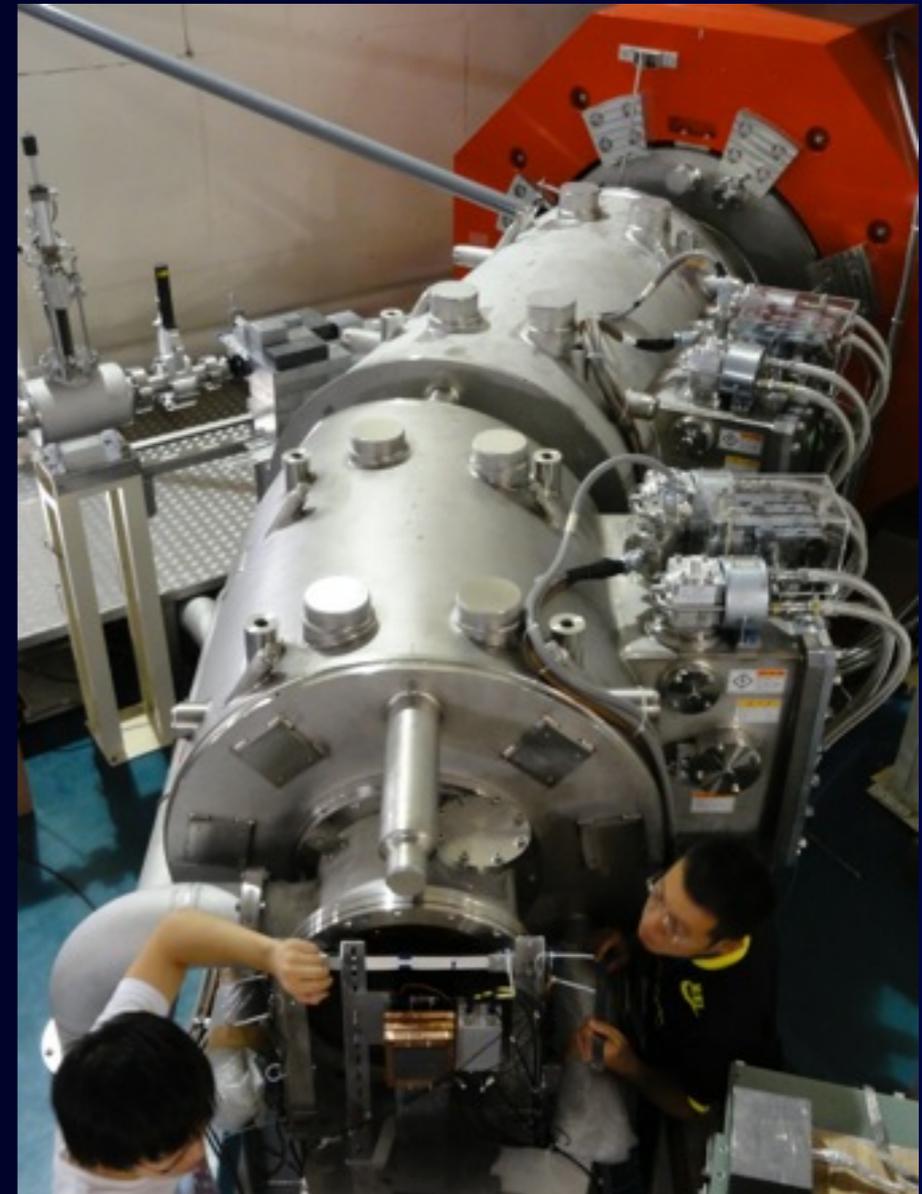
Yoshitaka Kuno
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LBNL

Outline



- Why Flavor Physics ?
Why Charged Lepton Flavor Violation (CLFV) ?
- CLFV Experiments
 - Muon to electron conversion in a muonic atom
- COMET (at J-PARC)
 - for sensitivity of $<10^{-16}$ (x10000)
- COMET Phase-I (at J-PARC)
 - for sensitivity of $<10^{-14}$ (x100)
- Summary



Why Flavor Physics ?
Why Charged Lepton Flavor
Violation (CLFV)?



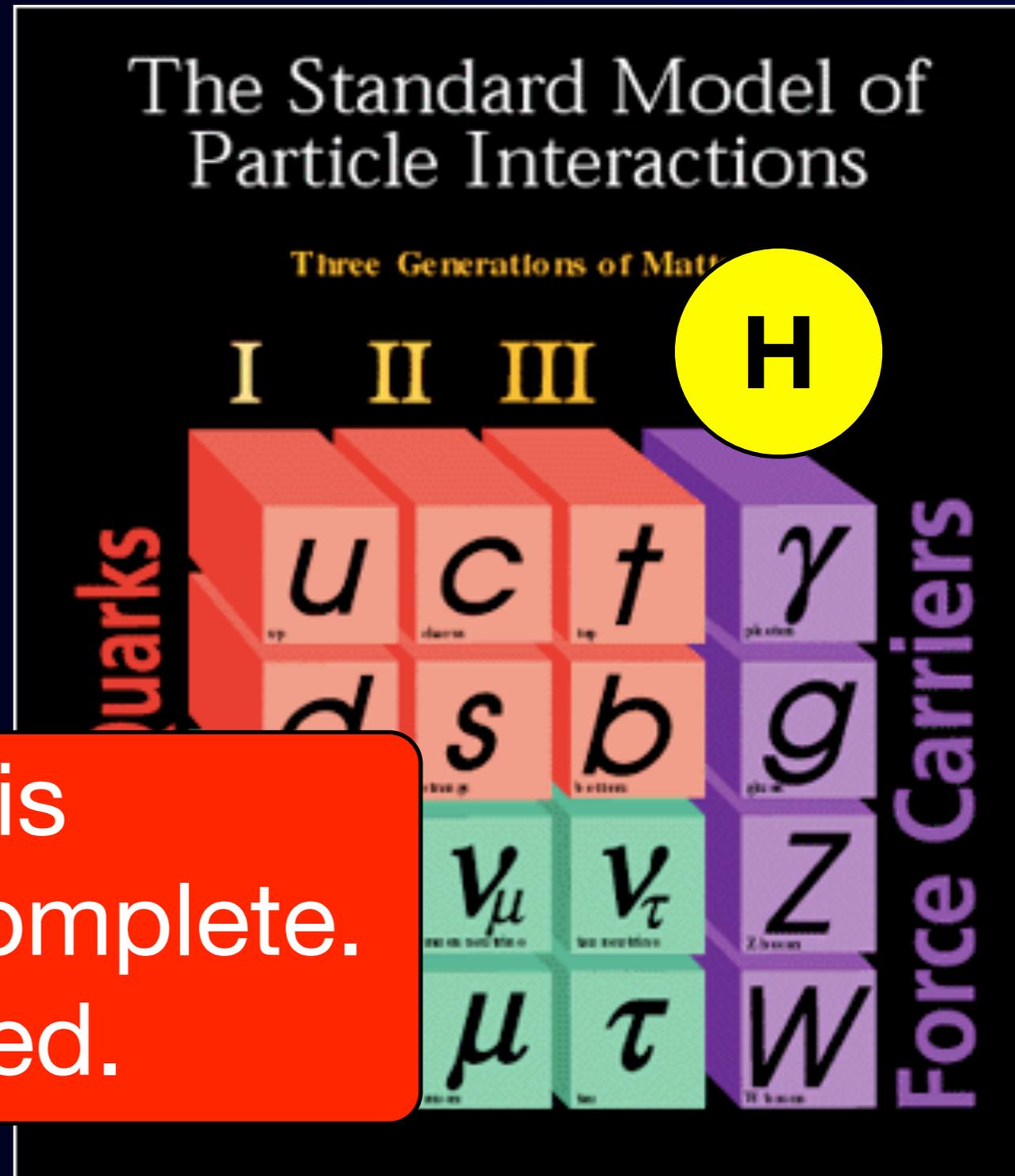
The Standard Model has the Higgs boson, but no new particles are found yet...



The discovery of the Higgs boson has been made.

The Standard Model can explain most of the experimental results. However, there are many undetermined parameters and issues.

The Standard Model is considered to be incomplete. New Physics is needed.



Flavor Physics and New Physics

flavor
structure



Effective Lagrangian in
the Standard Model (SM)

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{gauge}} + \mathcal{L}_{\text{sym. break.}}$$

$$V(\Phi) = -\mu^2 \Phi^\dagger \Phi + \lambda (\Phi^\dagger \Phi)^2 + Y^{ij} \Psi_L^i \Psi_R^j \Phi + \frac{g_{ij}}{\Lambda} \Psi_L^i \Psi_L^{jT} \Phi \Phi^T,$$

Origin of flavor

(1) what determines the observed pattern of masses and mixing angle of quarks and leptons ?

(2) which sources of flavor symmetry breaking are accessible at low energy ?

Ques.(1) is difficult to address owing to the lack of theoretical guidance.
Ques.(2) can be answered by a series of high-precision measurements

search for new physics

New Physics Search in Charged Lepton Flavor



with new physics contributions

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \frac{C_{\text{NP}}}{\Lambda^2} O_{ij}^{(6)},$$

Λ is the energy scale of new physics

Charged Lepton Flavor

For instance, $\mu \rightarrow e\gamma$ ($B < 5.7 \times 10^{-13}$),

Charged Lepton
Flavor Violation

$$\frac{C_{\text{NP}}}{\Lambda^2} O_{ij}^{(6)} \rightarrow \frac{C_{\mu e}}{\Lambda^2} \bar{e}_L \sigma^{\rho\nu} \mu_R \Phi F_{\rho\nu}$$

$$\Lambda > 2 \times 10^5 \text{ TeV} \times (C_{\mu e})^{\frac{1}{2}}.$$

$$\Lambda > O(10^5) \text{ TeV}$$

The constraint in CLFV is even more severe than in the quark flavor.
The SM contribution to muon CLFV is small, of the order of $O(10^{-54})$.

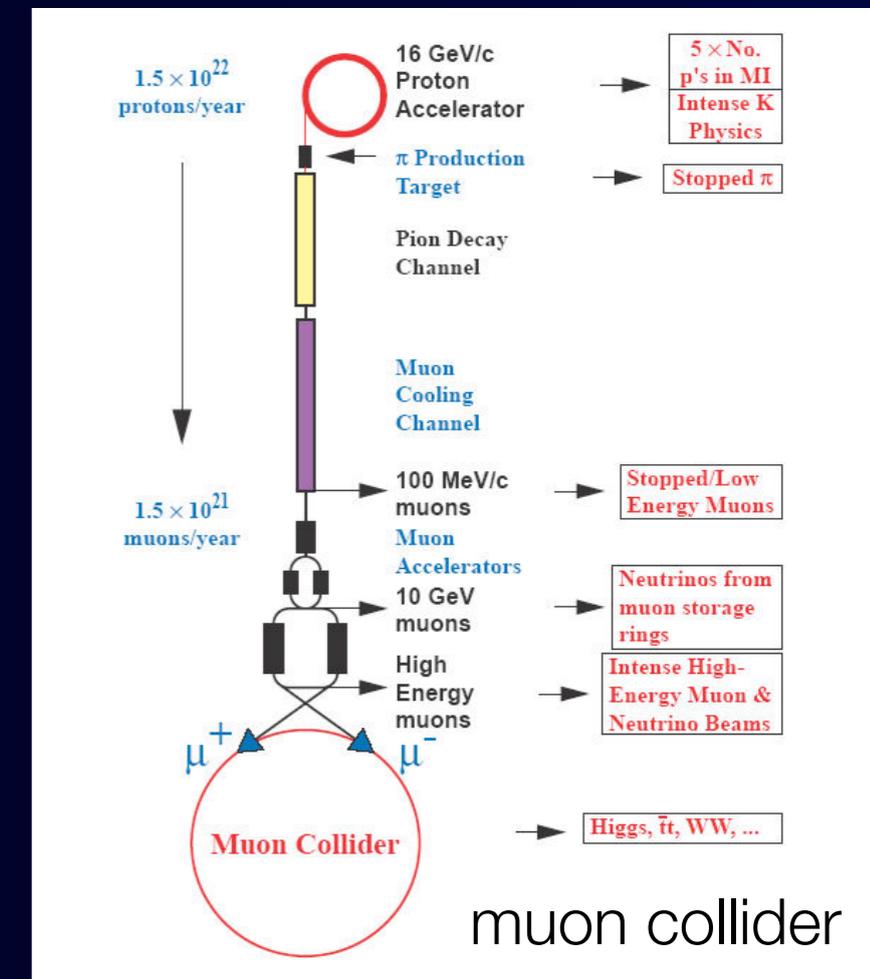
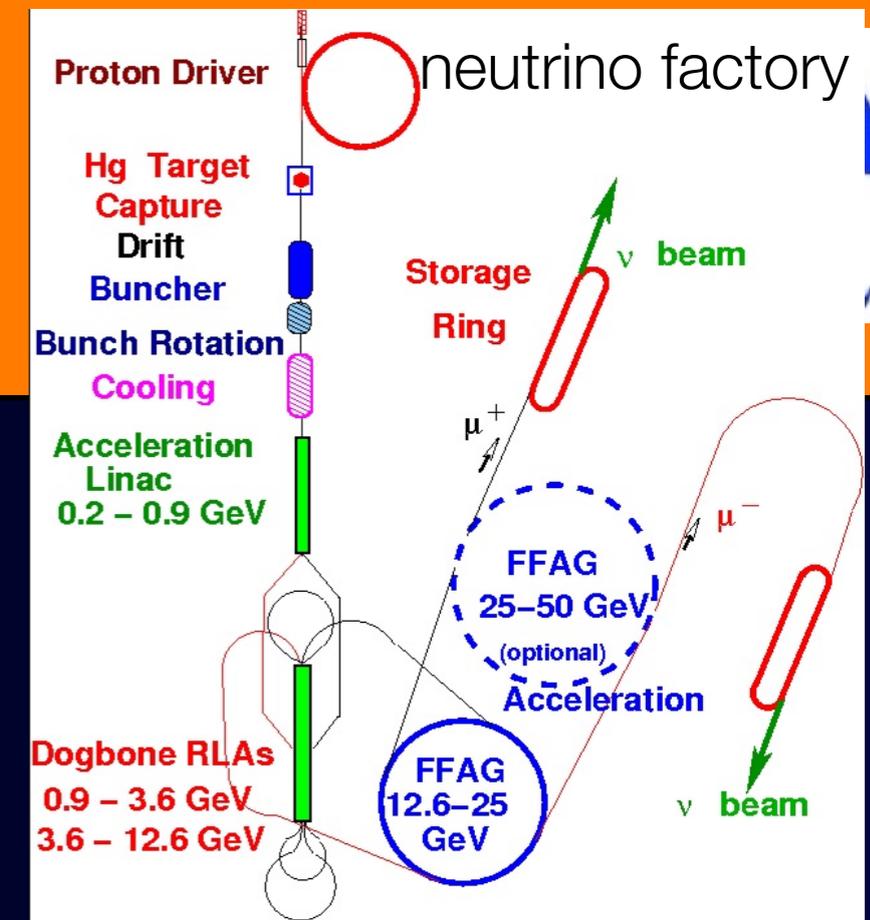
Which Rare Decays at Low Energy ?

- Processes which are forbidden or highly suppressed in the Standard Model would be the best ones to search for new physics beyond the Standard Model.
- **Flavor Changing Neutral Current Process (FCNC)**
- **FCNC in the quark sector**
 - $b \rightarrow s\gamma$, $K \rightarrow \pi\nu\nu$, etc.
 - Allowed in the Standard Model.
 - Need to study deviations from the SM predictions.
 - Uncertainty of more than a few % (from QCD) exists.
- **FCNC in the lepton sector**
 - $\mu \rightarrow e\gamma$, $\mu + N \rightarrow e + N$, etc. (**lepton flavor violation = LFV**)
 - Not allowed in the Standard Model ($\sim 10^{-50}$ with neutrino mixing)
 - Need to study deviations from none
 - clear signature and high sensitivity

Why Muons, not Taus?

- A number of taus available at B factories are about 1-10 taus/sec. At super-B factories, about 100 taus/sec are considered. Also some of the decay modes are already background-limited.
- A number of muons available now, which is about 10^8 muons/sec at PSI, is the largest. Next generation experiments aim 10^{11} - 10^{12} muons/sec. **With the technology of the front end of muon colliders and/or neutrino factories**, about 10^{13} - 10^{14} muons/sec are considered.

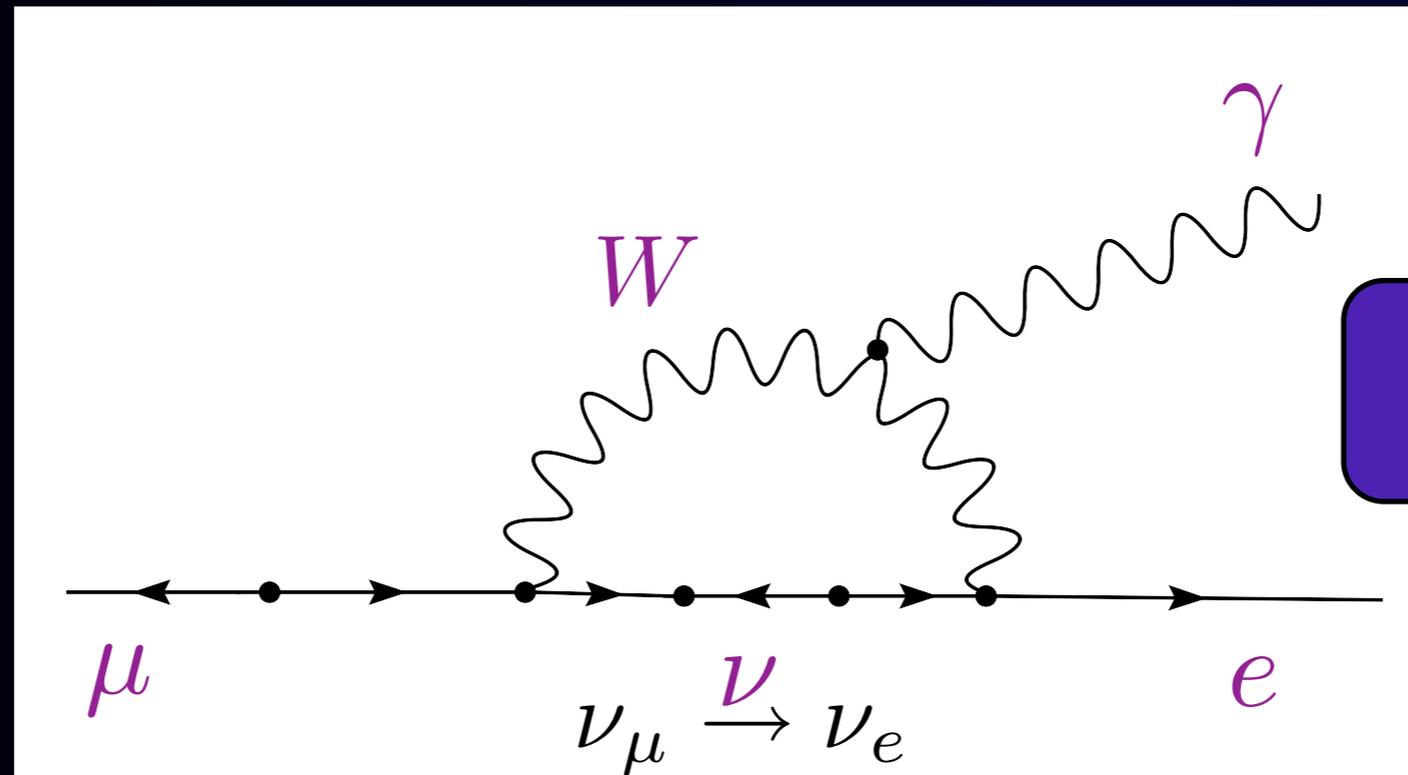
a larger window to search for new physics for muons than taus



No SM Contribution in Charged Lepton Flavor Violation (CLFV)



$$B(\mu \rightarrow e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_l (V_{MNS})_{\mu l}^* (V_{MNS})_{el} \frac{m_{\nu_l}^2}{M_W^2} \right|^2$$



BR $\sim O(10^{-54})$

Observation of CLFV would indicate a clear signal of physics beyond the SM with massive neutrinos.

Sensitivity to New Physics at High Energy

CLFV is unique and important to find new physics beyond the SM.

CLFV is sensitive to NP at high energy scale Λ .

others

amplitude

$$|A_{SM} + \varepsilon_{NP}|^2 \sim |A_{SM}|^2 + \underline{2\text{Re}(A_{SM}\varepsilon_{NP})} + |\varepsilon_N|^2$$

subject to uncertainty of SM prediction

CLFV

rate

$$|A_{SM} + \varepsilon_{NP}|^2 \sim \cancel{|A_{SM}|^2} + \cancel{2\text{Re}(A_{SM}\varepsilon_{NP})} + \underline{|\varepsilon_N|^2}$$

could go higher energy scale

$$R \propto \frac{1}{\Lambda^4}$$

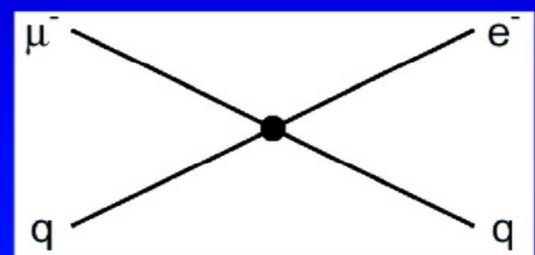
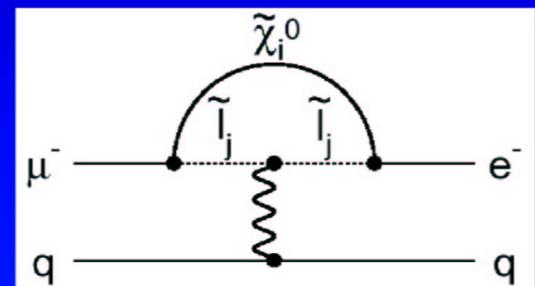
CLFV for muons can be improve by a factor of 10,000 or more, implying 10 times in energy reach.

Various Models Predict CLFV.....

Sensitivity to Different Muon Conversion Mechanisms

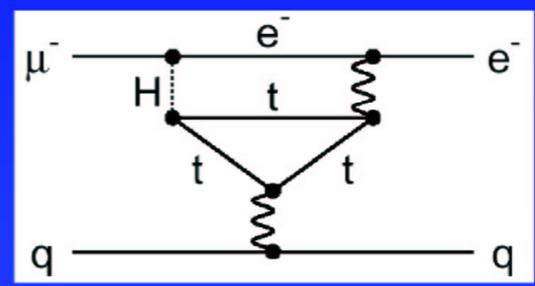
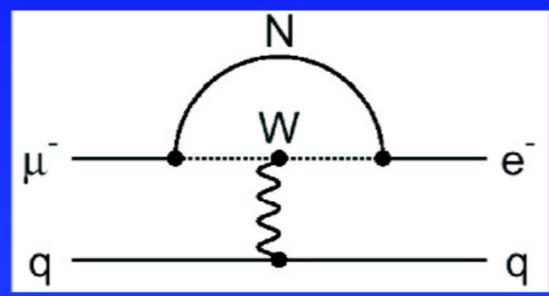


Supersymmetry
Predictions at 10^{-15}



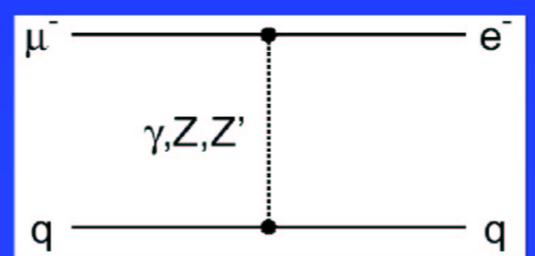
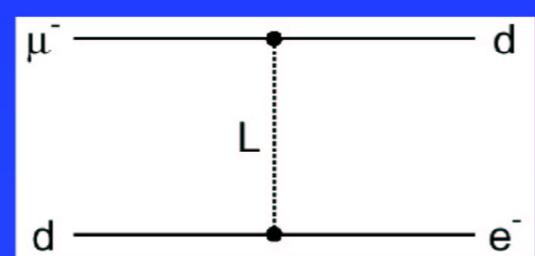
Compositeness
 $\Lambda_c = 3000 \text{ TeV}$

Heavy Neutrinos
 $|U_{\mu N}^* U_{eN}|^2 = 8 \times 10^{-13}$



Second Higgs doublet
 $g_{H\mu e} = 10^{-4} \times g_{H\mu\mu}$

Leptoquarks
 $M_L = 3000 (\lambda_{\mu d} \lambda_{ed})^{1/2} \text{ TeV}/c^2$



Heavy Z',
Anomalous Z
coupling
 $M_Z = 3000 \text{ TeV}/c^2$
 $B(Z \rightarrow \mu e) < 10^{-17}$

After W. Marciano

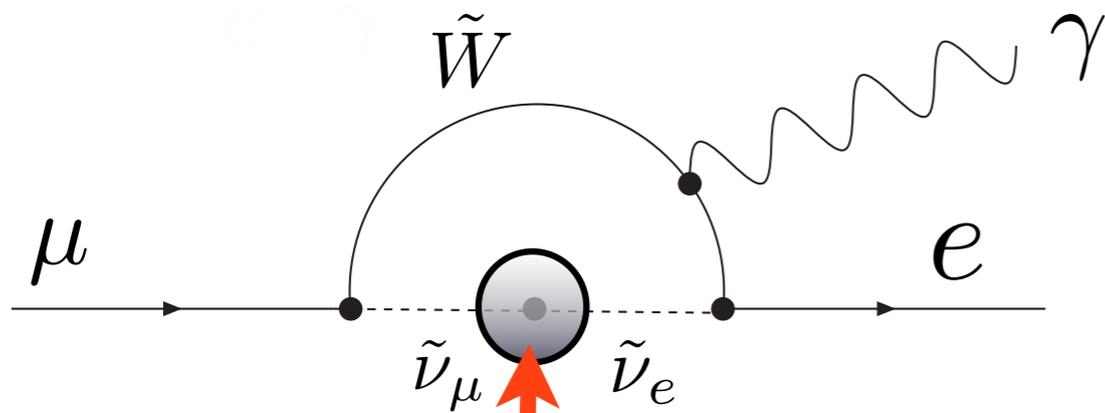
Example of Sensitivity to NP in High Energy Scale : SUSY models



For loop diagrams,

$$\text{BR}(\mu \rightarrow e\gamma) = 1 \times 10^{-11} \times \left(\frac{2\text{TeV}}{\Lambda}\right)^4 \left(\frac{\theta_{\mu e}}{10^{-2}}\right)^2 \quad y = \frac{g^2}{16\pi^2} \theta_{\mu e}$$

> sensitive to TeV energy scale with reasonable mixing



example diagram for SUSY (~TeV)

Physics at about 10^{16} GeV

slepton mixing
(from RGE)

$$(m_{\tilde{L}}^2)_{21} \sim \frac{3m_0^2 + A_0^2}{8\pi^2} h_t^2 V_{td} V_{ts} \ln \frac{M_{GUT}}{M_{R_s}}$$

$$(m_L^2)_{21} \sim \frac{3m_0^2 + A_0^2}{8\pi^2} h_\tau^2 U_{31} U_{32} \ln \frac{M_{GUT}}{M_R}$$

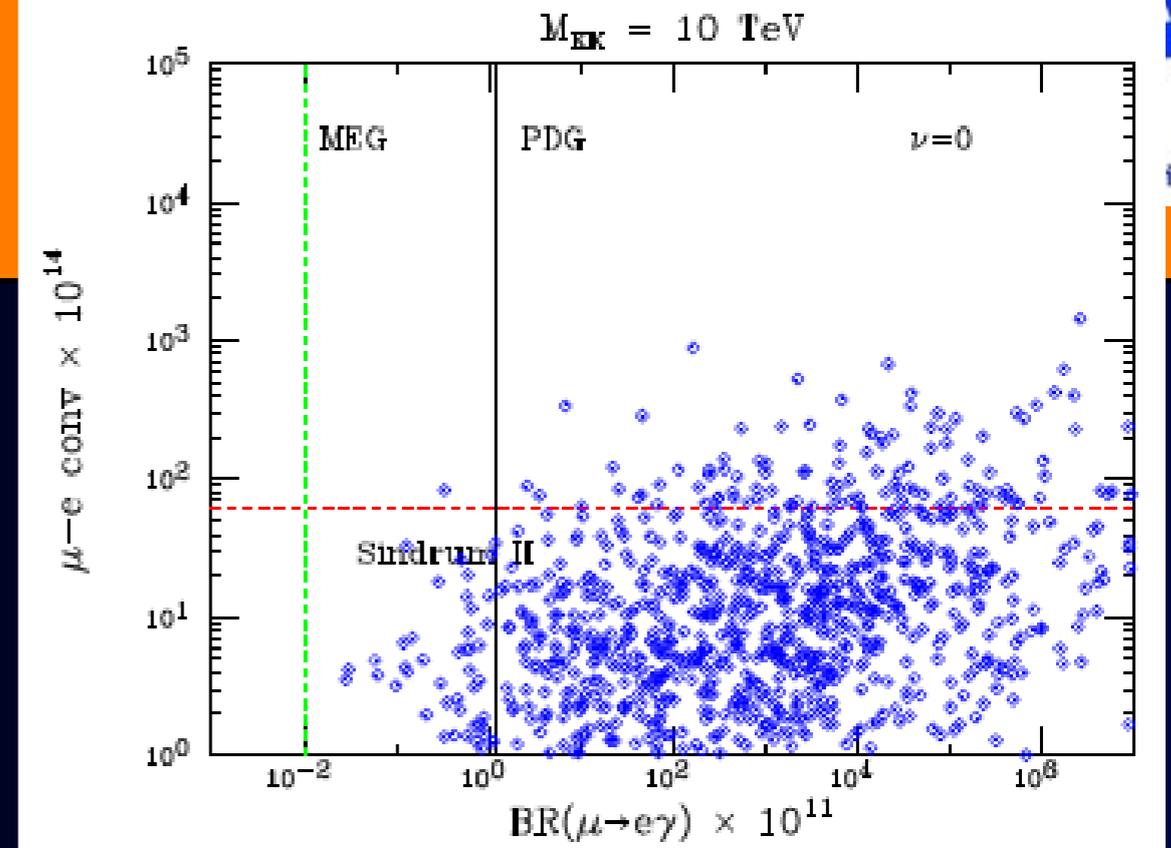
SUSY-GUT model

SUSY neutrino
seesaw model

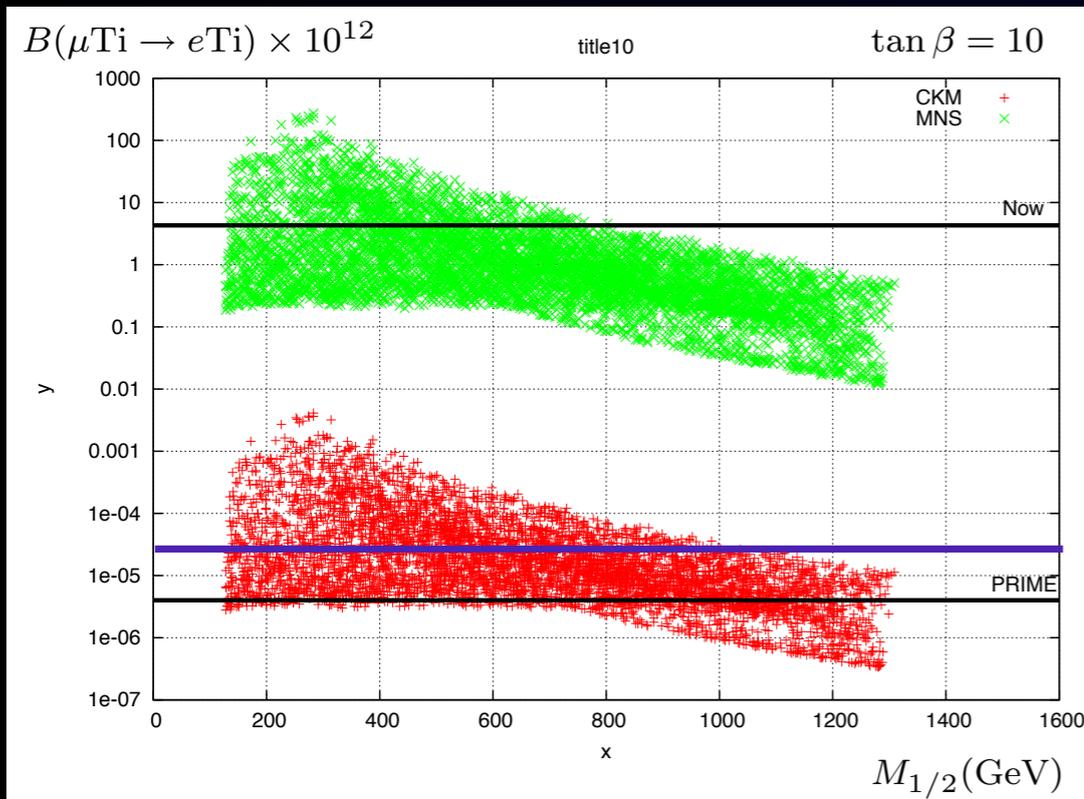
CLFV Predictions

Various BSM models predict sizable muon CLFV, as well as tau CLFV.

extra dimension model

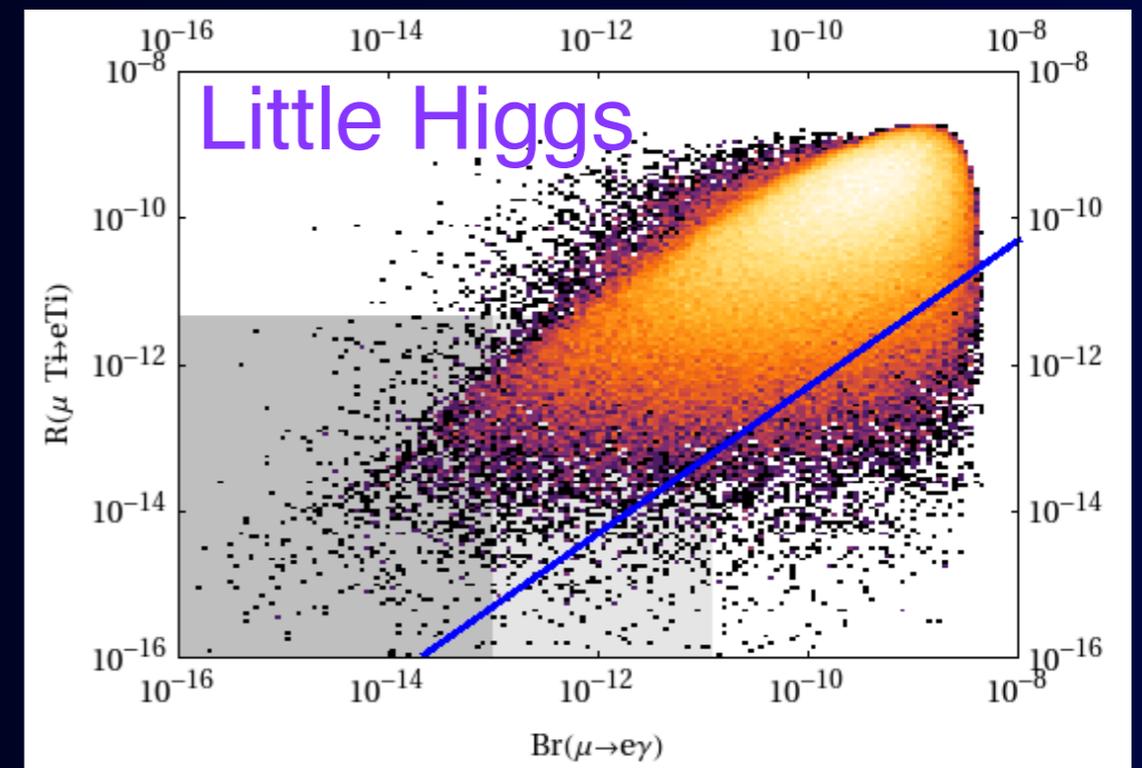


SUSY model



10^4

little Higgs model



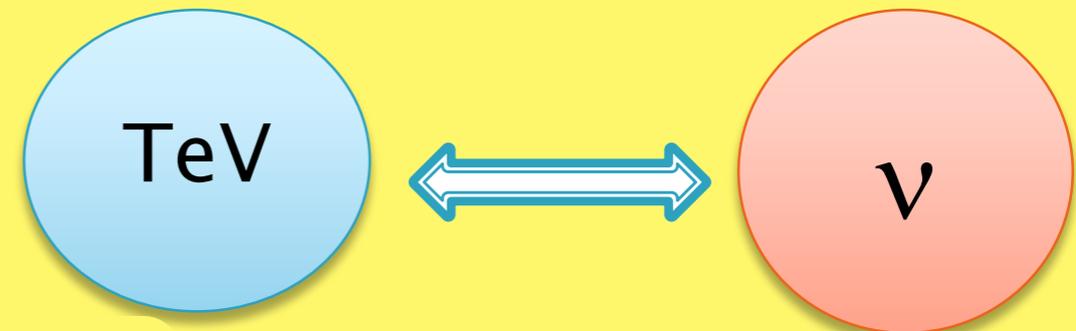
CLFV and Neutrino Mass Generation

Scale of the electroweak
symmetry breaking

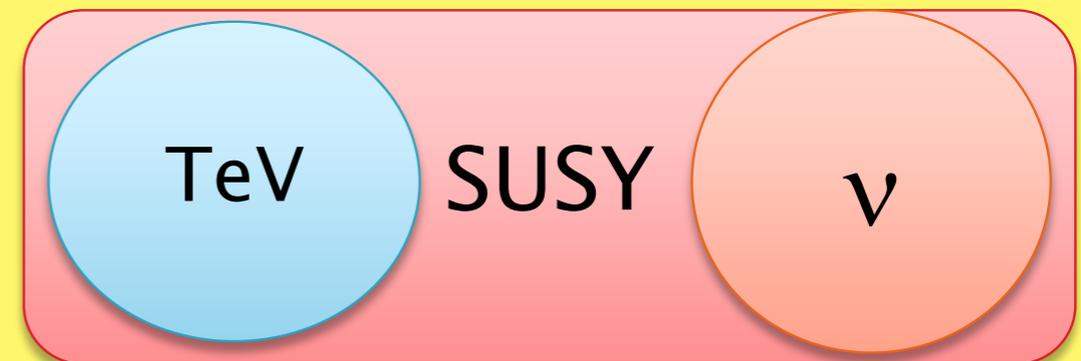
Scale of the neutrino
mass generation

If two scales are well separated,
LFVs are suppressed.

$$\text{CLFV} \sim \mathcal{O}(10^{-54})$$

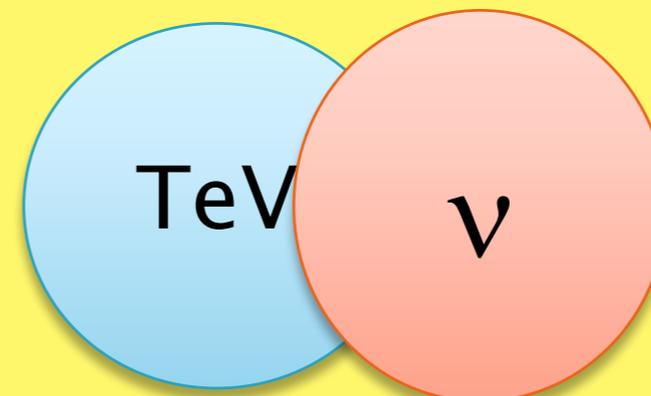


In supersymmetric models,
large LFV signals are expected
even if two scales are separated.



If two scales are close,
large LFVs are expected.

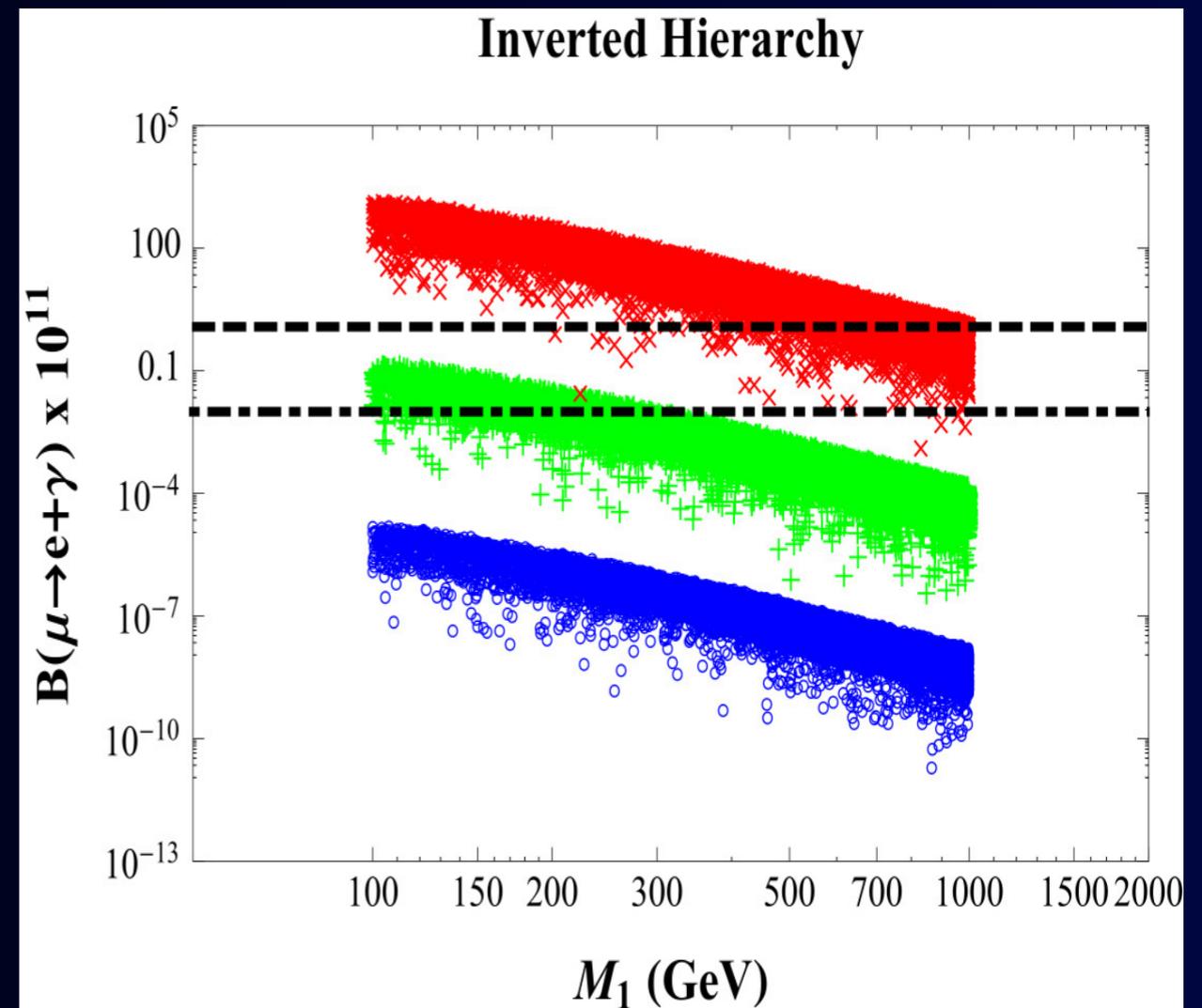
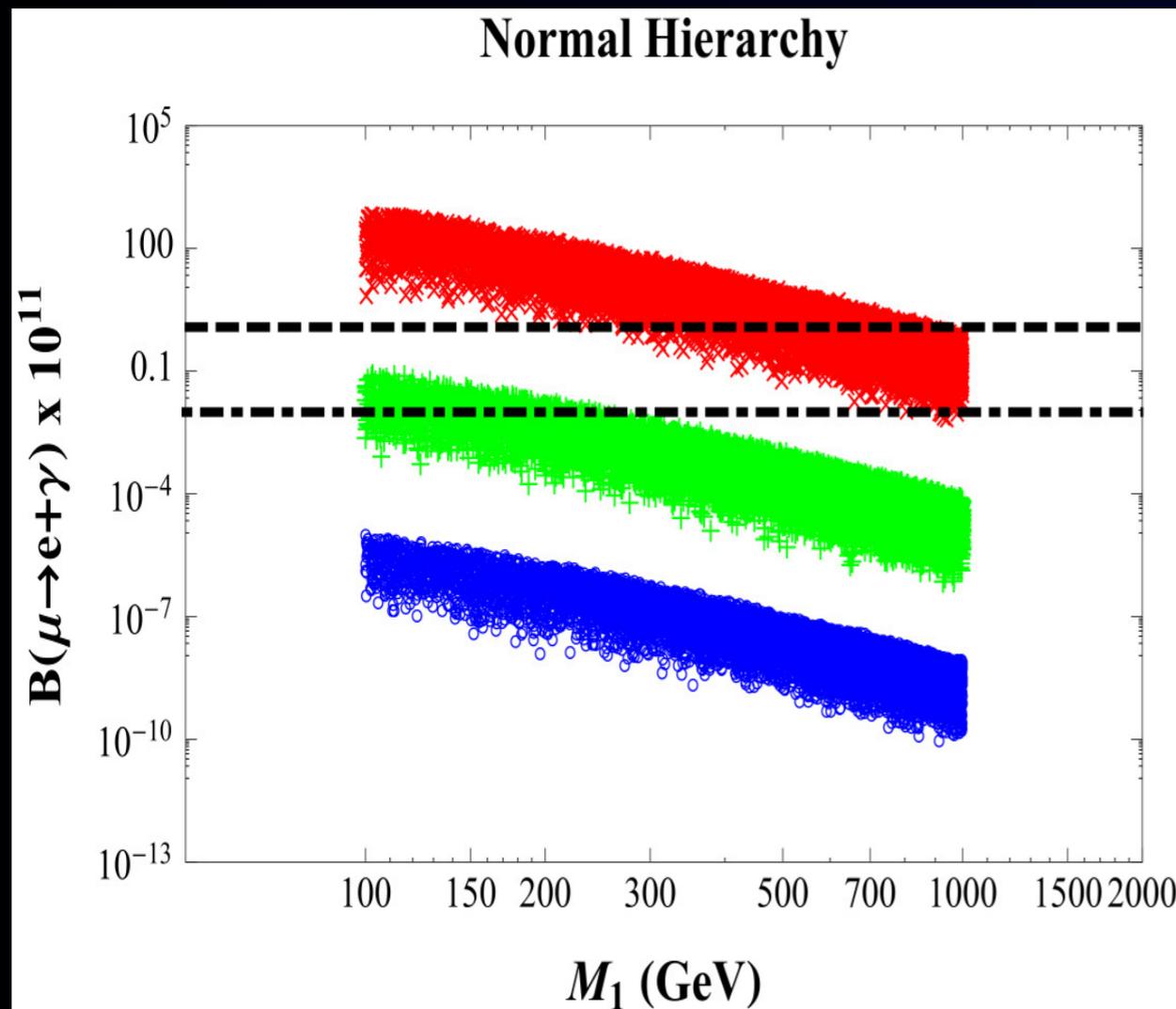
Neutrino mass from loop
Triplet Higgs for neutrino mass
Left-right symmetric model



1

2

CLFV with TeV Seesaw (Type-I)



TeV seesaw type-I models predict sizable branching ratio of CLFV with right-handed neutrino mass of $O(\text{TeV})$.

“DNA of New Physics” (a la Prof. Dr. A.J. Buras)



W. Altmannshofer, A.J. Buras, S. Gori, P. Paradisi and D.M. Straub

	AC	RVV2	AKM	δ LL	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	★★★★	★	★	★	★	★★★★	?
ϵ_K	★	★★★★	★★★★	★	★	★★	★★★★
$S_{\psi\phi}$	★★★★	★★★★	★★★★	★	★	★★★★	★★★★
$S_{\phi K_S}$	★★★★	★★	★	★★★★	★★★★	★	?
$A_{CP}(B \rightarrow X_s \gamma)$	★	★	★	★★★★	★★★★	★	?
$A_{7,8}(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★★★★	★★★★	★★	?
$A_9(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★	★	★	?
$B \rightarrow K^{(*)} \nu \bar{\nu}$	★	★	★	★	★	★	★
$B_s \rightarrow \mu^+ \mu^-$	★★★★	★★★★	★★★★	★★★★	★★★★	★	★
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	★	★	★	★	★	★★★★	★★★★
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	★	★	★	★	★	★★★★	★★★★
$\mu \rightarrow e \gamma$	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★
$\tau \rightarrow \mu \gamma$	★★★★	★★★★	★	★★★★	★★★★	★★★★	★★★★
$\mu + N \rightarrow e + N$	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★
d_n	★★★★	★★★★	★★★★	★★	★★★★	★	★★★★
d_e	★★★★	★★★★	★★	★	★★★★	★	★★★★
$(g-2)_\mu$	★★★★	★★★★	★★	★★★★	★★★★	★	?

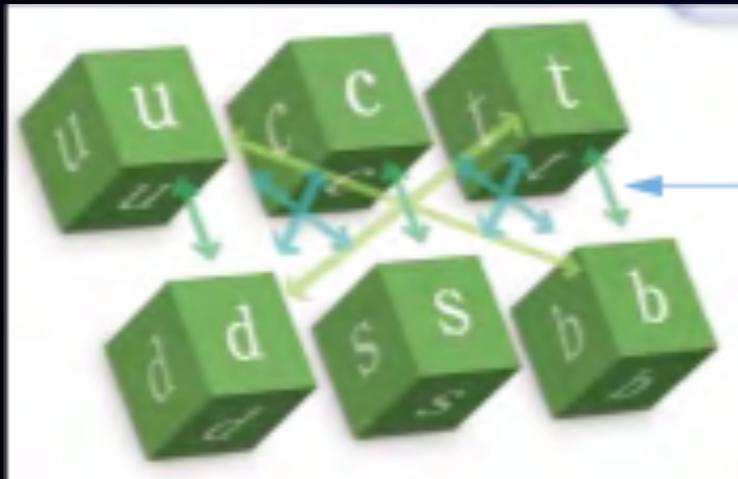
The pattern of measurement:
 ★ ★ ★ large effects
 ★ ★ visible but small effects
 ★ unobservable effects
 is characteristic,
 often uniquely so,
 of a particular model

GLOSSARY	
AC [10]	RH currents & U(1) flavor symmetry
RVV2 [11]	SU(3)-flavored MSSM
AKM [12]	RH currents & SU(3) family symmetry
δLL [13]	CKM-like currents
FBMSSM [14]	Flavor-blind MSSM
LHT [15]	Little Higgs with T Parity
RS [16]	Warped Extra Dimensions

These are a subset of a subset listed by Buras and Girschbach
 MFV, CMFV, 2HDM_{MFV}, LHT, SM4, SUSY flavor. SO(10) – GUT,
 SSU(5)_{HN}, FBMSSM, RHMfV, L-R, RS₀, gauge flavor,

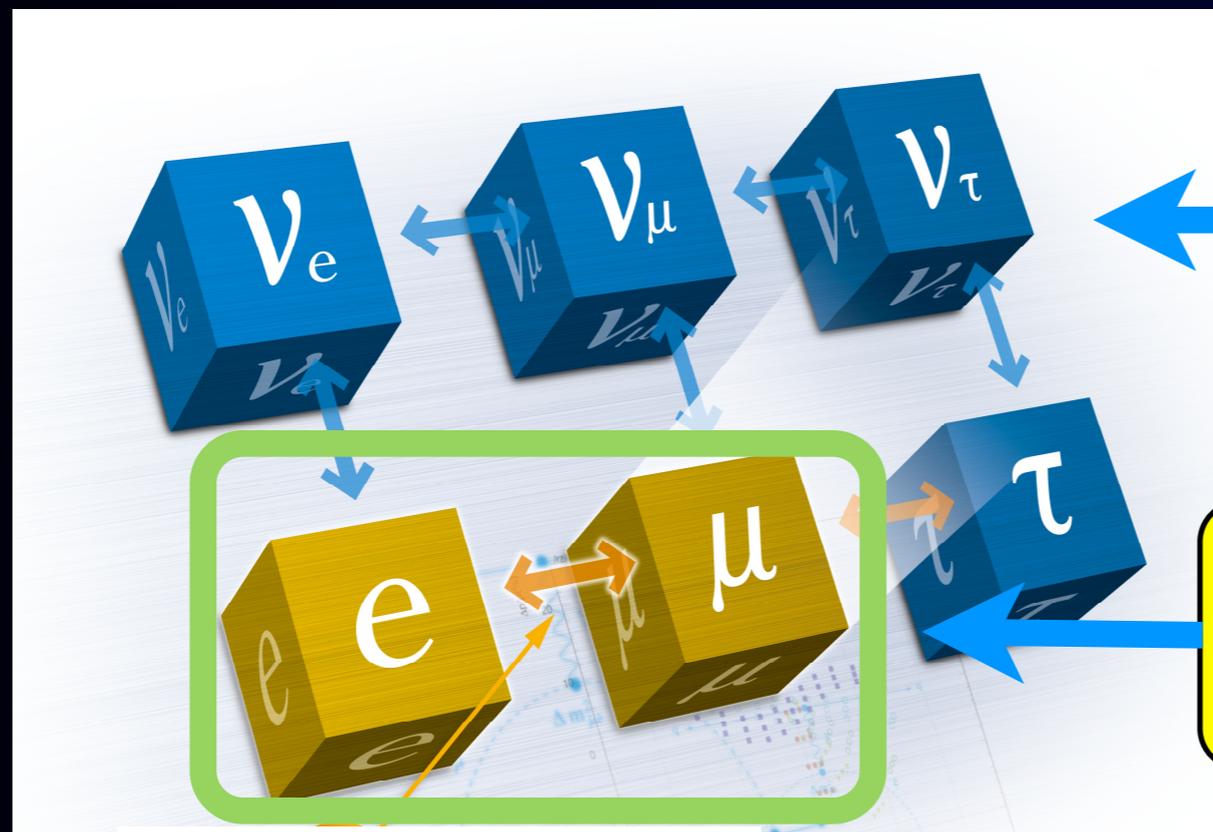
Quarks, Neutrinos, and Charged Leptons

Quarks



Quark mixing
observed

Lepton



Neutrino mixing
observed

Charged lepton mixing
not observed.

Charged Lepton Flavor Violation (CLFV)

Nobel Prize-winning
class research

CLFV Experiments



Present Limits and Expectations in Future

process	present limit	future	
$\mu \rightarrow e\gamma$	$<5.7 \times 10$	<10	MEG at PSI
$\mu \rightarrow eee$	$<1.0 \times 10$	<10	Mu3e at PSI
$\mu N \rightarrow eN$ (in Al)	none	<10	Mu2e / COMET
$\mu N \rightarrow eN$ (in Ti)	$<4.3 \times 10$	<10	PRISM
$\tau \rightarrow e\gamma$	$<1.1 \times 10$	<10	superKEKB
$\tau \rightarrow eee$	$<3.6 \times 10$	<10	superKEKB
$\tau \rightarrow \mu\gamma$	$<4.5 \times 10$	<10	superKEKB
$\tau \rightarrow \mu\mu\mu$	$<3.2 \times 10$	<10	superKEKB/LHCb

List of cLFV Processes with Muons

$\Delta L=1$

- $\mu^+ \rightarrow e^+ \gamma$
- $\mu^+ \rightarrow e^+ e^+ e^-$
- $\mu^- + N(A, Z) \rightarrow e^- + N(A, Z)$
- $\mu^- + N(A, Z) \rightarrow e^+ + N(A, Z - 2)$

← this talk

$\Delta L=2$

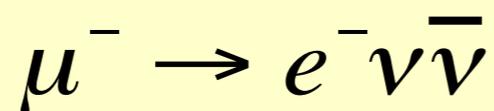
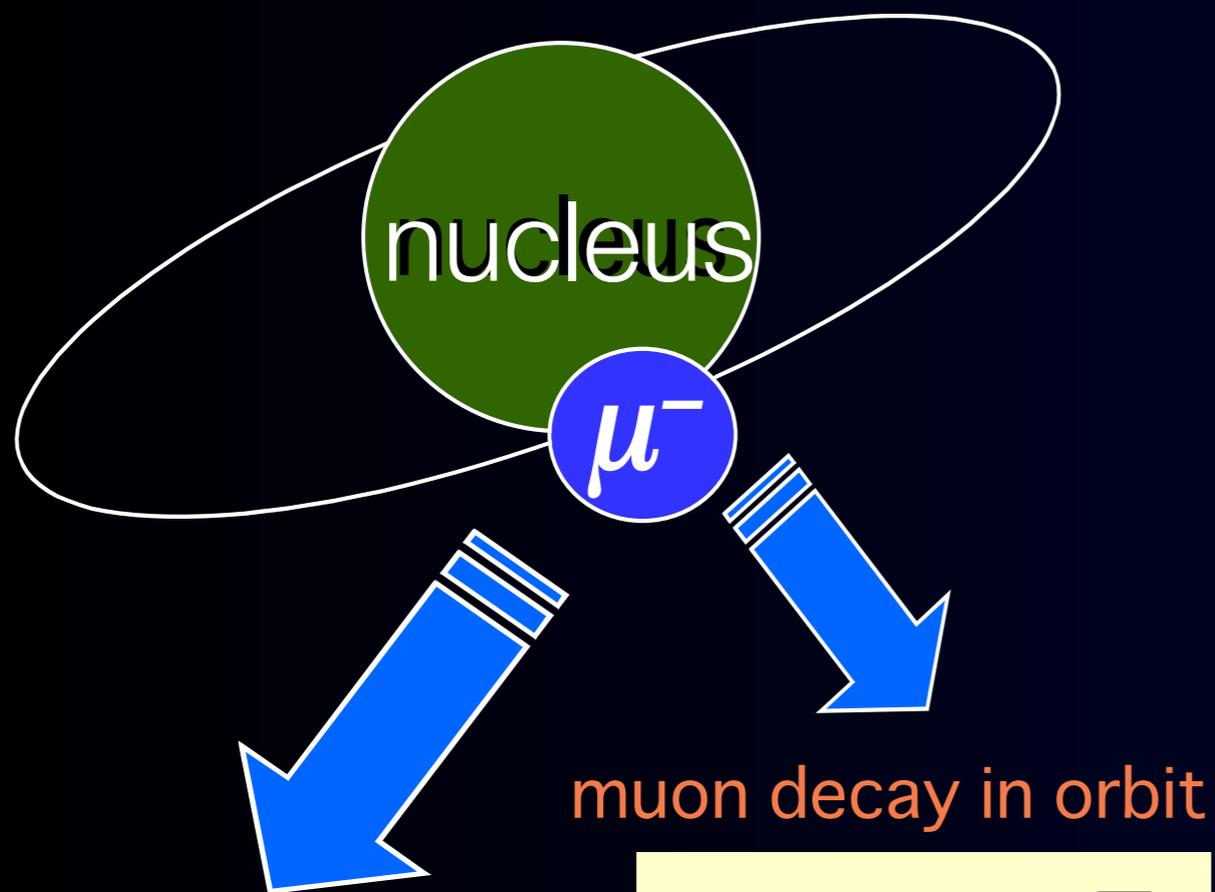
- $\mu^+ e^- \rightarrow \mu^- e^+$
- $\mu^- + N(A, Z) \rightarrow \mu^+ + N(A, Z - 2)$
- $\nu_\mu + N(A, Z) \rightarrow \mu^+ + N(A, Z - 1)$
- $\nu_\mu + N(A, Z) \rightarrow \mu^+ \mu^+ \mu^- + N(A, Z - 1)$



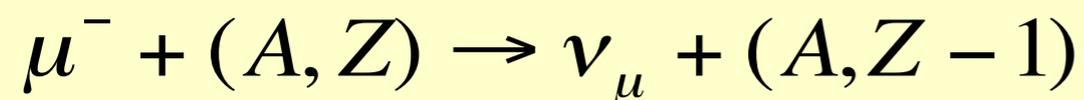
$\mu \rightarrow e$ conversion
in
a muonic atom

What is Muon to Electron Conversion?

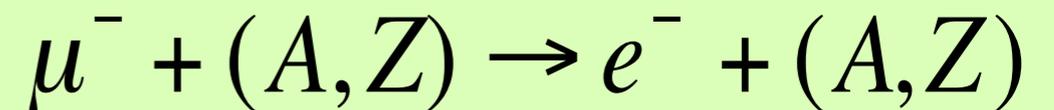
1s state in a muonic atom



nuclear muon capture



Neutrino-less muon nuclear capture



Event Signature :

a single mono-energetic electron of 105 MeV

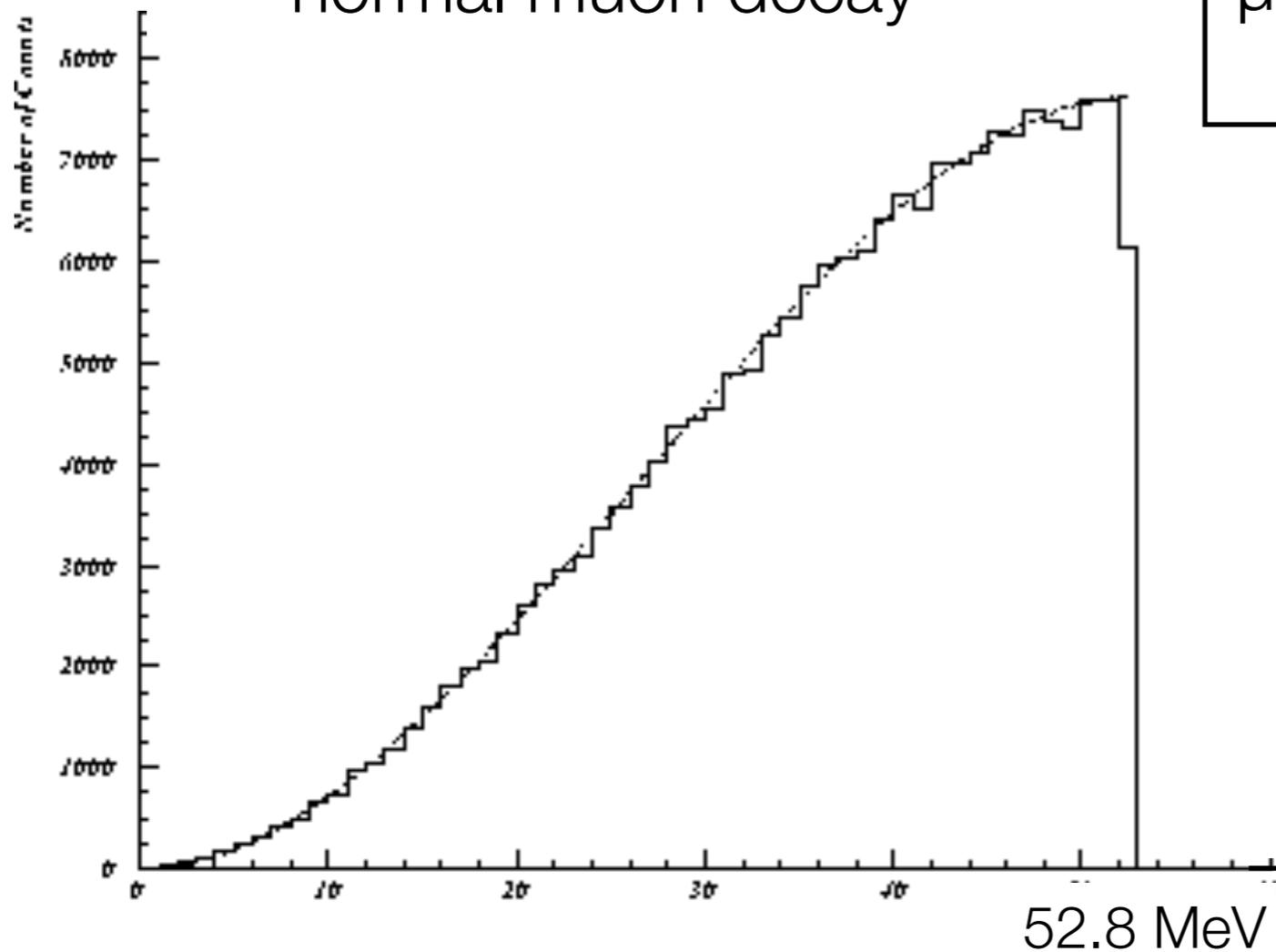
Backgrounds:

- (1) physics backgrounds
ex. muon decay in orbit (DIO)
- (2) beam-related backgrounds
ex. radiative pion capture,
muon decay in flight,
- (3) cosmic rays, false tracking

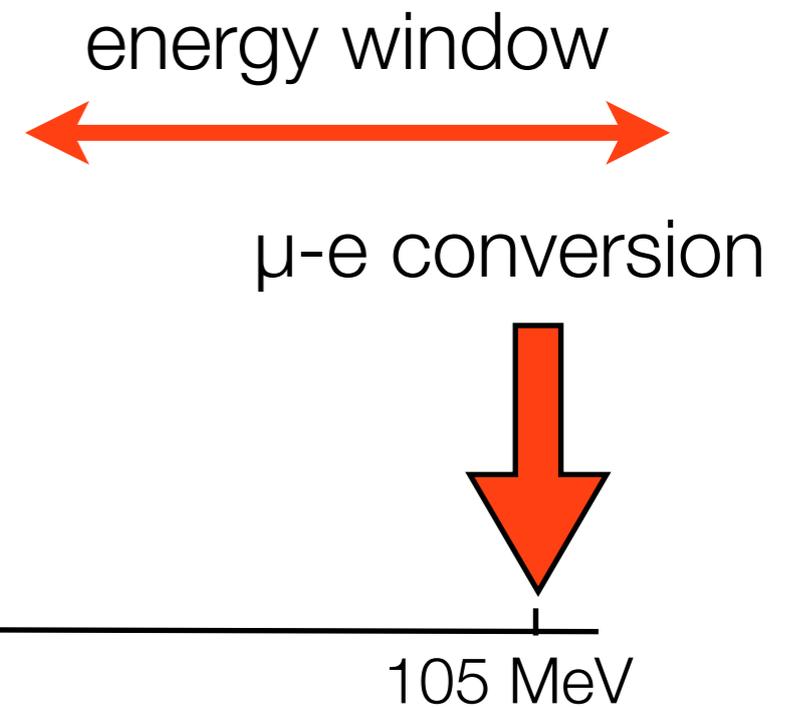
μ -e Conversion Signal and Normal Muon Decays



normal muon decay



μ -e conversion and muon Michel decays are well separated.



electron momentum spectrum

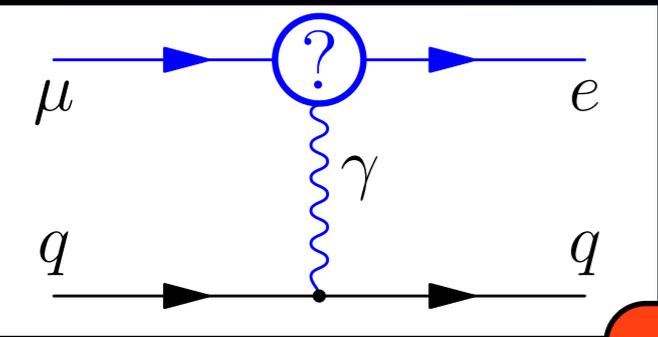
High Intensity beam can be used only for μ -e conversion

Physics Sensitivity: $\mu \rightarrow e\gamma$ vs. μ -e conversion

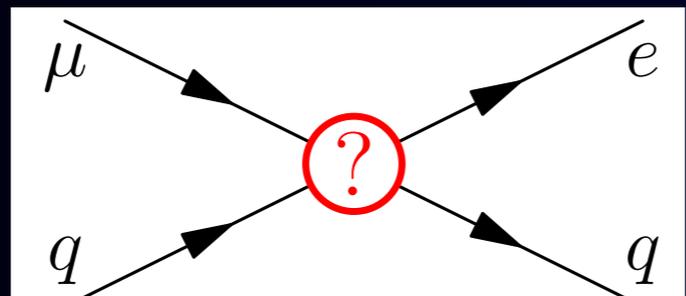


$$L_{\text{CLFV}} = \frac{1}{1 + \kappa} \frac{m_\mu}{\Lambda^2} \bar{\mu}_R \sigma^{\mu\nu} e_L F_{\mu\nu} + \frac{\kappa}{1 + \kappa} \frac{1}{\Lambda^2} (\bar{\mu}_L \gamma^\mu e_L) (\bar{q}_L \gamma_\mu q_L)$$

Photonic (dipole) interaction



Contact interaction

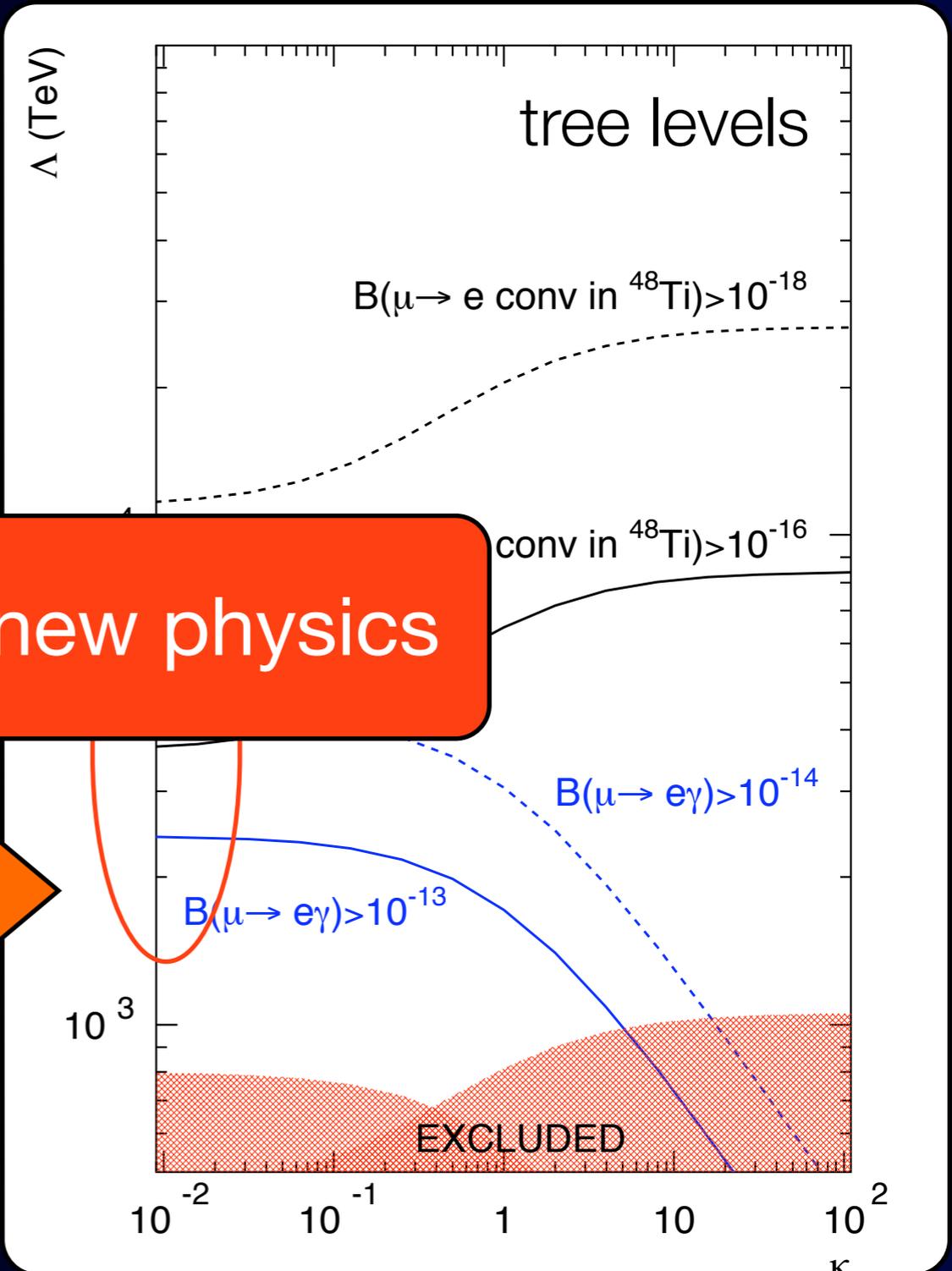


more sensitive to new physics

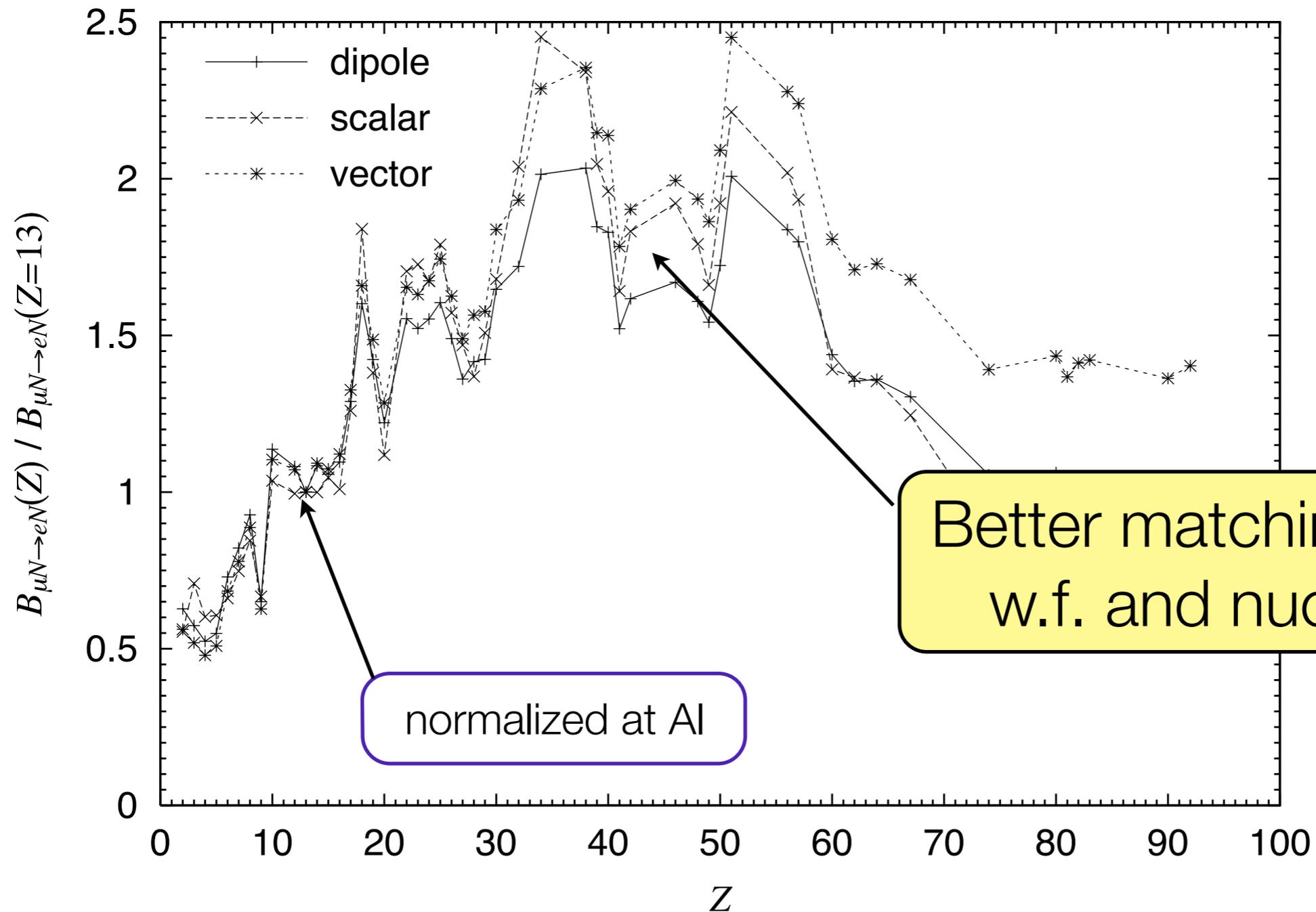
if photonic contri

$$\frac{B(\mu N \rightarrow eN)}{B(\mu \rightarrow e\gamma)} = \frac{G_F^2 m_\mu^4}{96\pi^3 \alpha} \times 3 \times 10^{12} B(A, Z) \sim \frac{B(A, Z)}{428}$$

- for aluminum, about 1/390~0.003
- for titanium, about 1/230



μ -e Conversion : Target dependence (discriminating effective interaction)



Experimental Comparison between $\mu \rightarrow e\gamma$ and μ -e Conversion



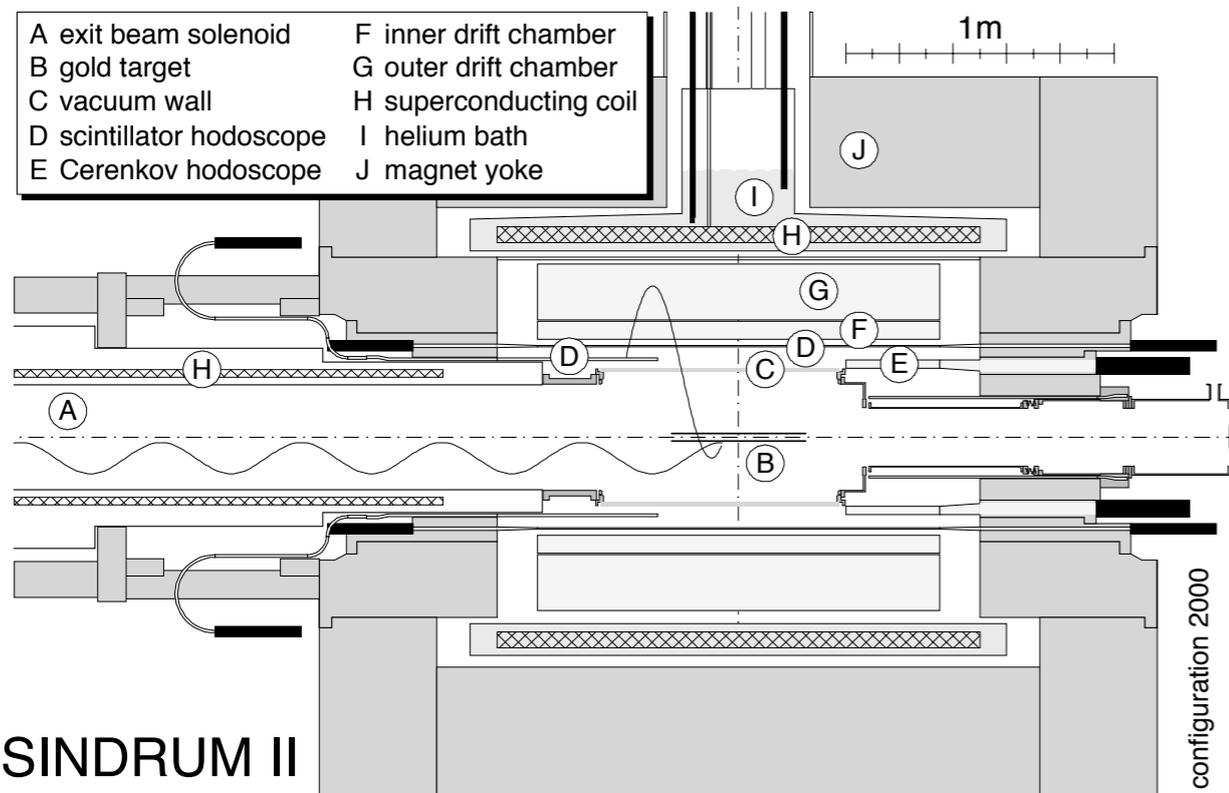
	background	challenge	beam intensity
• $\mu \rightarrow e\gamma$	accidentals	detector resolution	limited
• μ -e conversion	beam	beam background	no limitation

- $\mu \rightarrow e\gamma$:
 - Accidental background is given by $(\text{rate})^2$.
 - The detector resolutions have to be improved, but difficult.
 - The ultimate sensitivity would be about 10^{-14} .
- μ -e conversion :
 - A higher beam intensity can be taken because of no accidentals.
 - Improvement of a muon beam can be possible.
 - high intensity and high purity

μ -e conversion might be a next step.

Previous Measurements

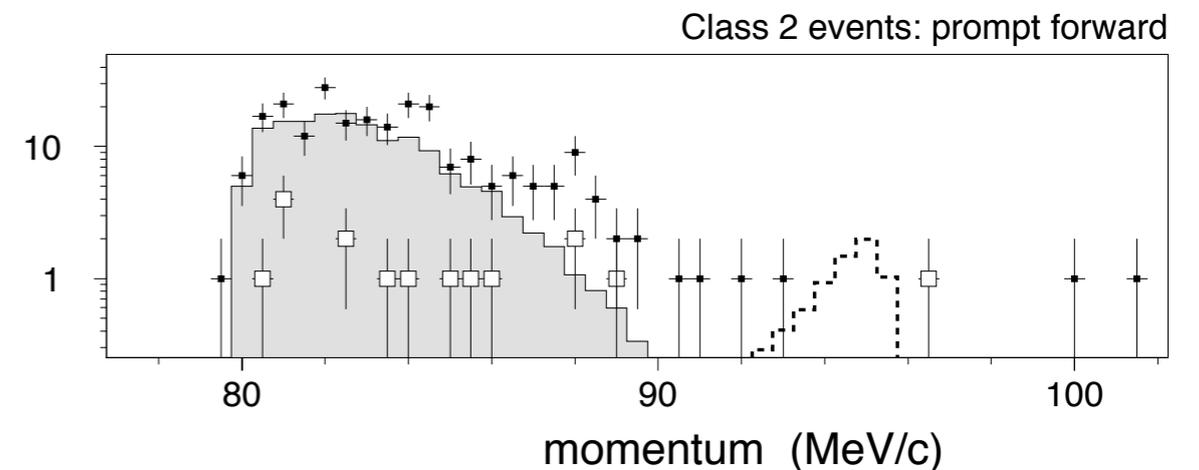
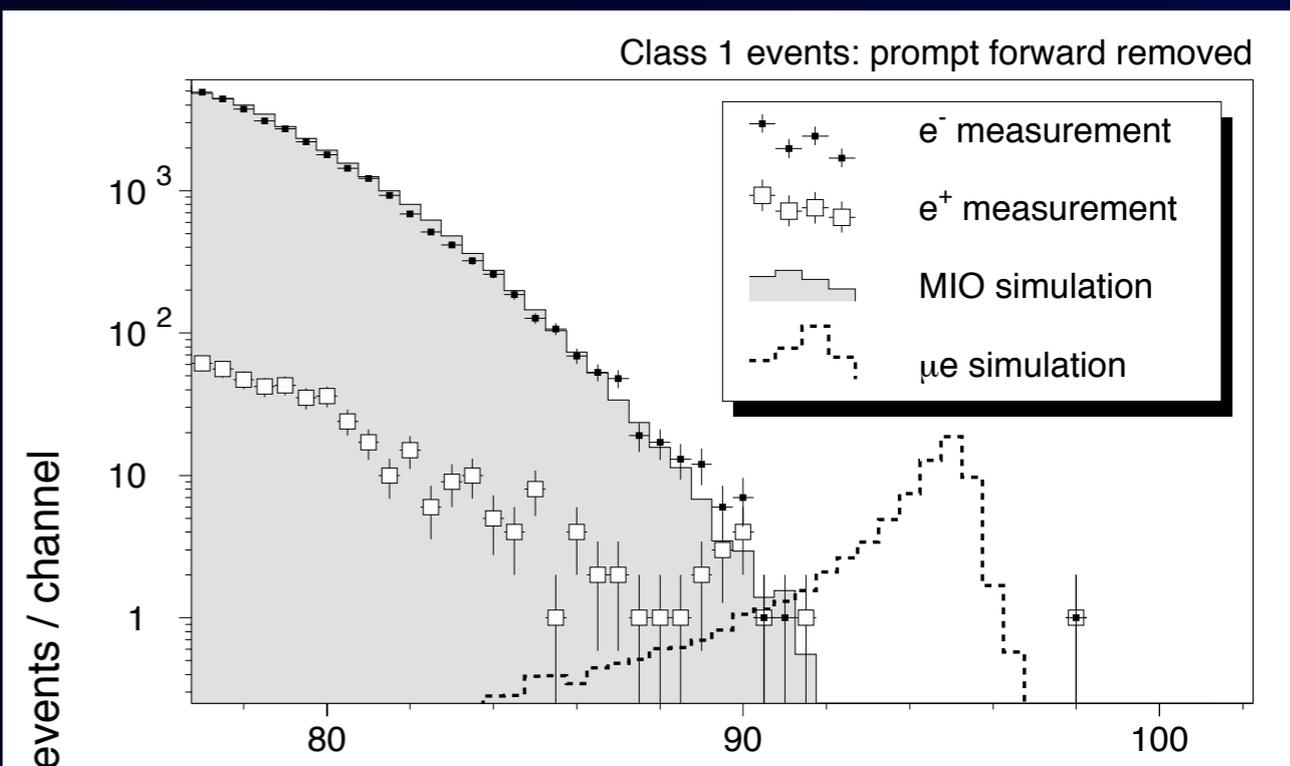
SINDRUM-II (PSI)



PSI muon beam intensity $\sim 10^{7-8}/\text{sec}$ beam from the PSI cyclotron. To eliminate beam related background from a beam, a beam veto counter was placed. But, it could not work at a high rate.

Published Results (2004)

$$B(\mu^- + Au \rightarrow e^- + Au) < 7 \times 10^{-13}$$



Improvements for Signal Sensitivity

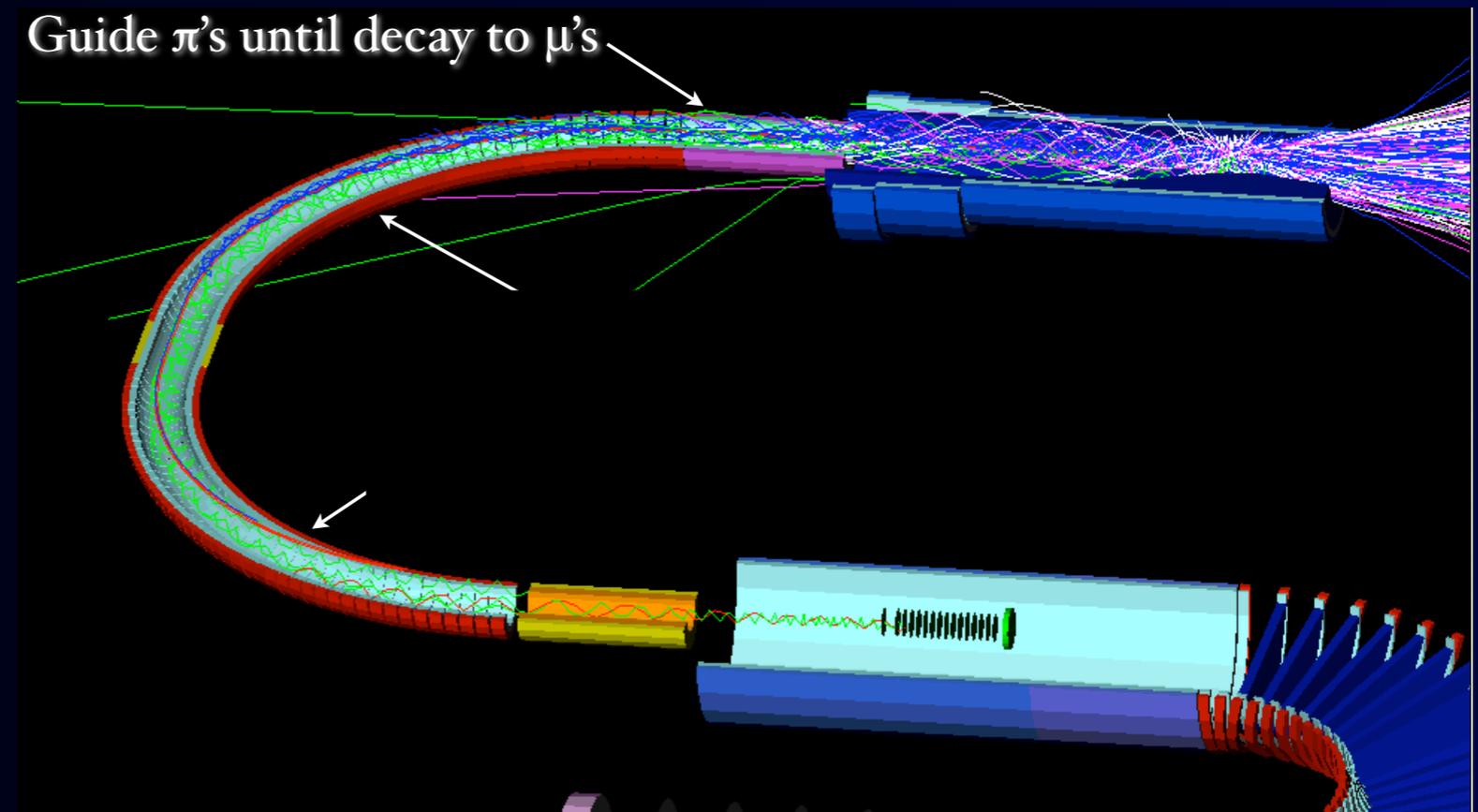
To achieve a single sensitivity of 10^{-17} , we need

10^{11} muons/sec (with 10^7 sec running)

whereas the current highest intensity is 10^8 /sec at PSI.

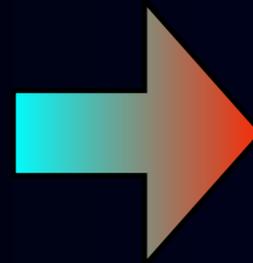
Pion Capture and
Muon Transport by
Superconducting
Solenoid System

(10^{11} muons for 50
kW beam power)



Improvements for Background Rejection

Beam-related
backgrounds

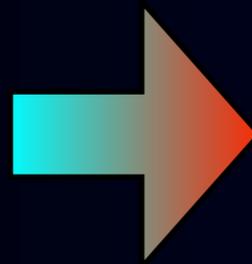


Beam pulsing with
separation of $1\mu\text{sec}$

measured
between beam
pulses

proton extinction = #protons between pulses/#protons in a pulse $< 10^{-9}$

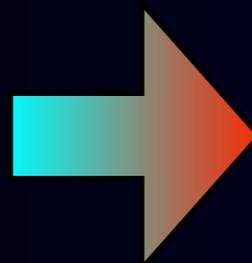
Muon DIO
background



low-mass trackers in
vacuum & thin target

improve
electron energy
resolution

Muon DIF
background

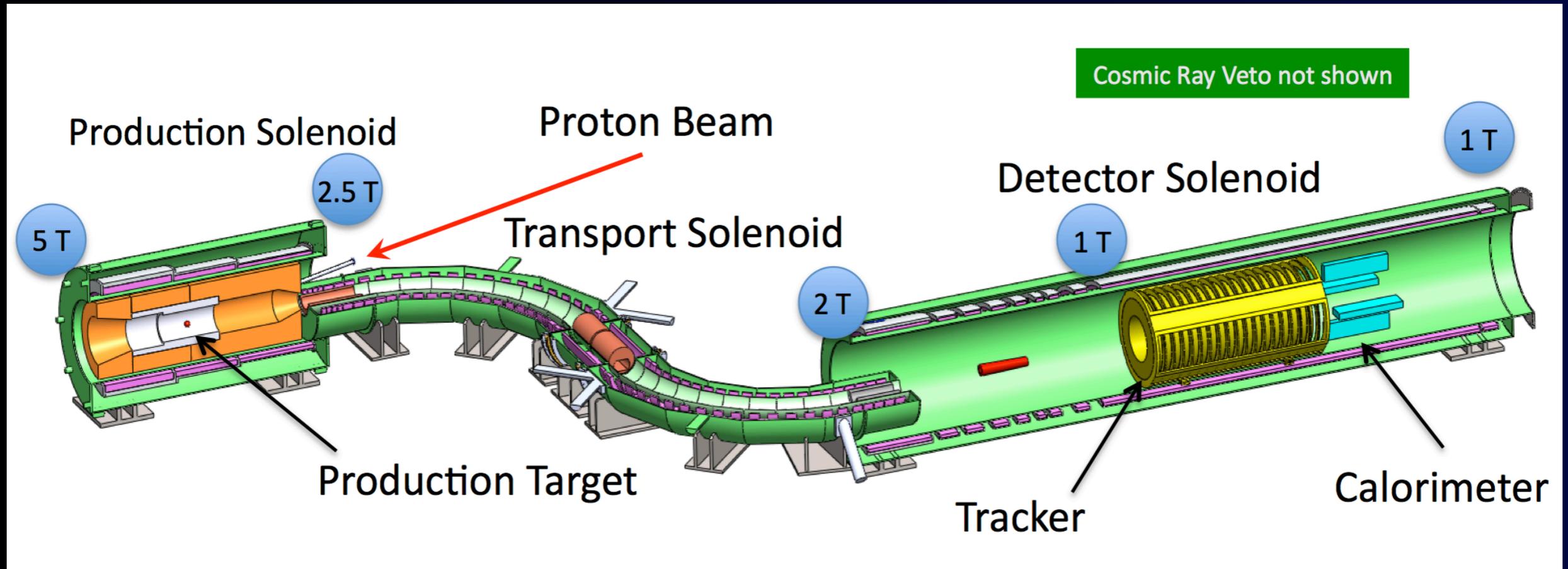
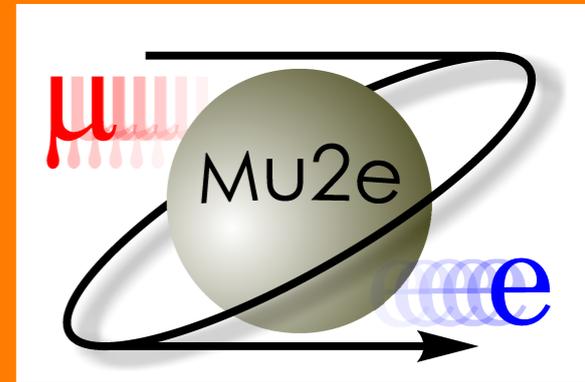


curved solenoids for
momentum selection

eliminate
energetic muons
($>75\text{ MeV}/c$)

base on the MELC proposal at Moscow Meson Factory

μ -e conversion : Mu2e at Fermilab



$$B(\mu^- + Al \rightarrow e^- + Al) = 5 \times 10^{-17} \quad (\text{S.E.})$$

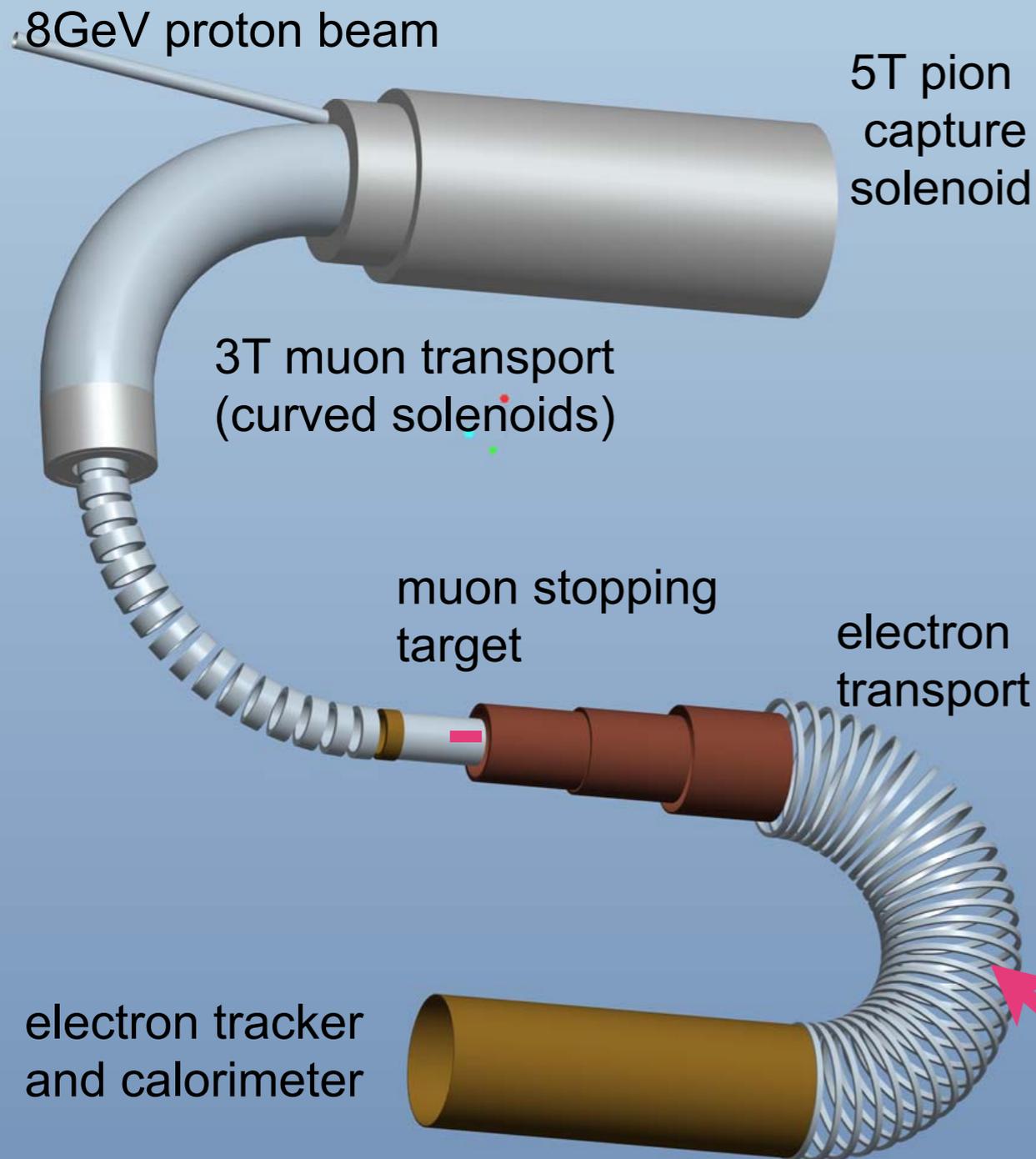
$$B(\mu^- + Al \rightarrow e^- + Al) < 10^{-16} \quad (90\% \text{C.L.})$$

- Reincarnation of MECO at BNL.
- Antiproton buncher ring is used to produce a pulsed proton beam.
- Approved in 2009, and CD0 in 2009, and CD1 in 2011.
- Data taking starts in about 2019.

COMET



What is COMET (E21) at J-PARC



Experimental Goal of COMET

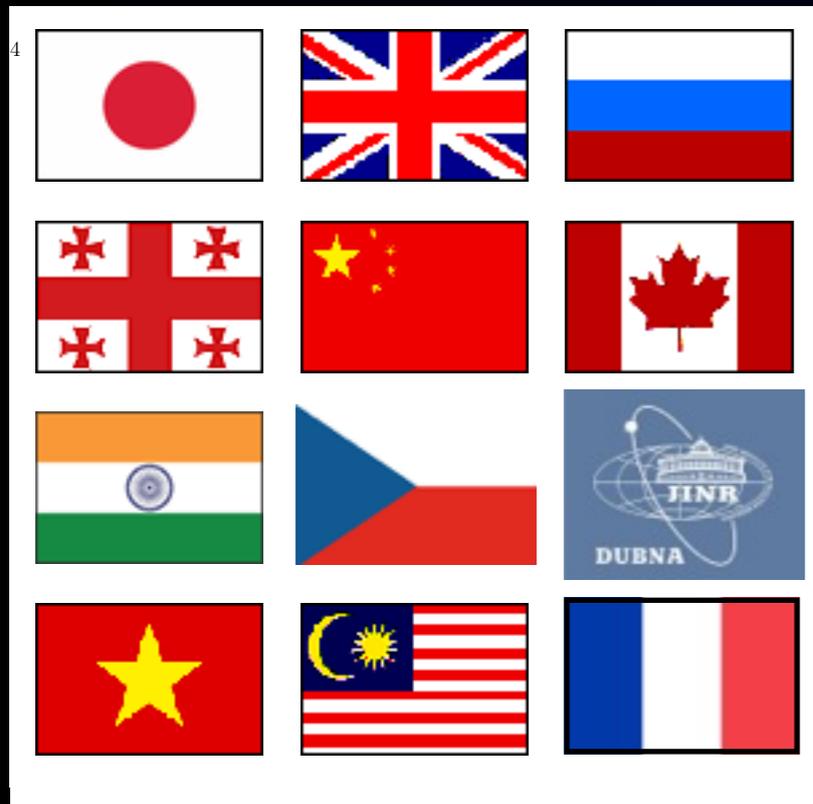
$$B(\mu^- + Al \rightarrow e^- + Al) = 2.6 \times 10^{-17}$$

$$B(\mu^- + Al \rightarrow e^- + Al) < 6 \times 10^{-17} \quad (90\% C.L.)$$

- 10^{11} muon stops/sec for 56 kW proton beam power.
- 2×10^7 running time (~ 1 year)
- C-shape muon beam line
- C-shape electron transport followed by electron detection system.
- Stage-1 approved in 2009.

Electron transport with curved solenoid would make momentum and charge selection.

COMET Collaboration is international.



129 collaborators
28 institutes, 11 countries

The
COMET
Japan
group
funded.

The
COMET
China
group
funded.

The
COMET
JINR group
funding
underway.

The COMET Collaboration

R. Akhmetshin³, K. Akuma¹⁷, M. Aoki²², R. B. Appleby¹⁹, Y. Arimoto¹², Y. Bagaturia⁷, W. Bertsche¹⁹, A. Bondar³, D. Bryman², B. Chiladze⁵, M. Danilov¹⁰, W. daSilva¹⁶, P. Dauncey⁸, G. Devidze⁵, P. Dornan⁸, A. Drutskoy¹⁰, S. Dymov¹¹, A. Edmonds²⁵, L. Epshteyn³, P. Evtoukhovich¹¹, G. Fedotov³, Y. Fukao¹², M. Gersabeck¹⁹, D. Grigoriev³, K. Hasegawa¹², I. H. Hasim²², O. Hayashi²², M. I. Hossain¹⁸, Z. Ibrahim¹⁷, F. Idris¹⁷, Y. Igarashi¹², F. Ignatov³, M. Ikeno¹², S. Ishimoto¹², T. Itahashi²², S. Ito²², T. Iwami²², Y. Iwashita¹³, X. Jiang⁴, P. Jonsson⁸, V. Kalinnikov¹¹, F. Kapusta¹⁶, H. Katayama²², K. Kawagoe¹⁵, V. Kazanin³, B. Khazin³, A. Khvedelidze¹¹, M. Koike²⁶, G. Kozlov¹¹, B. Krikler⁸, A. Kulikov¹¹, Y. Kuno²², Y. Kuriyama¹⁴, A. Kurup⁸, B. Lagrange¹⁴, M. Lancaster²⁵, H. B. Li⁴, W. Li⁴, A. Liparteliani⁵, G. Macharashvili¹¹, Y. Makida¹², Y. Matsumoto²², T. Mibe¹², S. Mihara¹², A. Moiseenko¹¹, Y. Mori¹⁴, N. Mosulishvili⁵, E. Motuk²⁵, Y. Nakai¹⁵, T. Nakamoto¹², T. H. Nam²², J. Nash⁸, M. Nioradze⁵, H. Nishiguchi¹², T. Numao²⁴, T. Ogitsu¹², K. Okamoto²², C. Omori¹², K. Ooishi¹⁵, T. Ota²³, H. Owen¹⁹, R. Palmer¹, C. Parkes¹⁹, J. Pasternak⁸, A. Popov³, V. Rusinov¹⁰, A. Ryzhenkov³, B. Sabirov¹¹, N. Saito¹², H. Sakamoto²², P. Sarin⁶, K. Sasaki¹², A. Sato²², J. Sato²³, D. Shemyakin³, V. Shmakova¹¹, M. Sugano¹², W. Tajudeen¹⁷, Y. Takubo¹², M. Tanaka¹², C. V. Tao²¹, E. Tarkovsky¹⁰, Y. Tevzadze⁵, N. D. Thong²², V. Thuan⁹, J. Tojo¹⁵, M. Tomizawa¹², I. Trekov⁵, N. M. Truong²², Z. Tsmalaidze¹¹, N. Tsverava¹¹, S. Tygier¹⁹, T. Uchida¹², Y. Uchida⁸, K. Ueno¹², S. Umasankar⁶, E. Velicheva¹¹, A. Volkov¹¹, M. Warren²⁵, M. Wing²⁵, C. Wu⁴, G. Xia¹⁹, K. Yai²², A. Yamamoto¹², M. Yamanaka²⁰, M. Yoshida¹², Y. Yoshii¹², K. Yoshimura¹², T. Yoshioka¹⁵, Y. Yuan⁴, Y. Yudin³, Y. Zhang⁴

¹Department of Physics, Brookhaven National Laboratory, USA ²University of British Columbia, Vancouver, Canada ³Budker Institute of Nuclear Physics (BINP), Novosibirsk, Russia ⁴Institute of High Energy Physics (IHEP), China ⁵Institute of High Energy Physics of I. Javakishvili State University (HEPI-TSU), Tbilisi, Georgia ⁶Indian Institute of Technology, Bombay, India ⁷Ilia State University (ISU), Tbilisi, Georgia ⁸Imperial College London, UK ⁹Institute for Nuclear Science and Technology, Vietnam ¹⁰Institute for Theoretical and Experimental Physics (ITEP), Russia ¹¹Joint Institute for Nuclear Research (JINR), Dubna, Russia ¹²High Energy Accelerator Research Organization (KEK), Tsukuba, Japan ¹³Institute for Chemical Research, Kyoto University, Kyoto, Japan ¹⁴Research Reactor Institute, Kyoto University, Kyoto, Japan ¹⁵Kyushu University, Fukuoka, Japan ¹⁶Laboratory of Nuclear and High Energy Physics (LPNHE), CNRS-IN2P3 and University Pierre and Marie Curie (UPMC), Paris, France ¹⁷University of Malaya, Malaysia ¹⁸University Technology Malaysia, Johor, Malaysia ¹⁹University of Manchester, UK ²⁰Nagoya University, Nagoya, Japan ²¹College of Natural Science, National Vietnam University, Vietnam ²²Osaka University, Osaka, Japan ²³Saitama University, Japan ²⁴TRIUMF, Canada ²⁵University College London, UK ²⁶Utsunomiya University, Utsunomiya, Japan

Proton Beam

J-PARC@Tokai

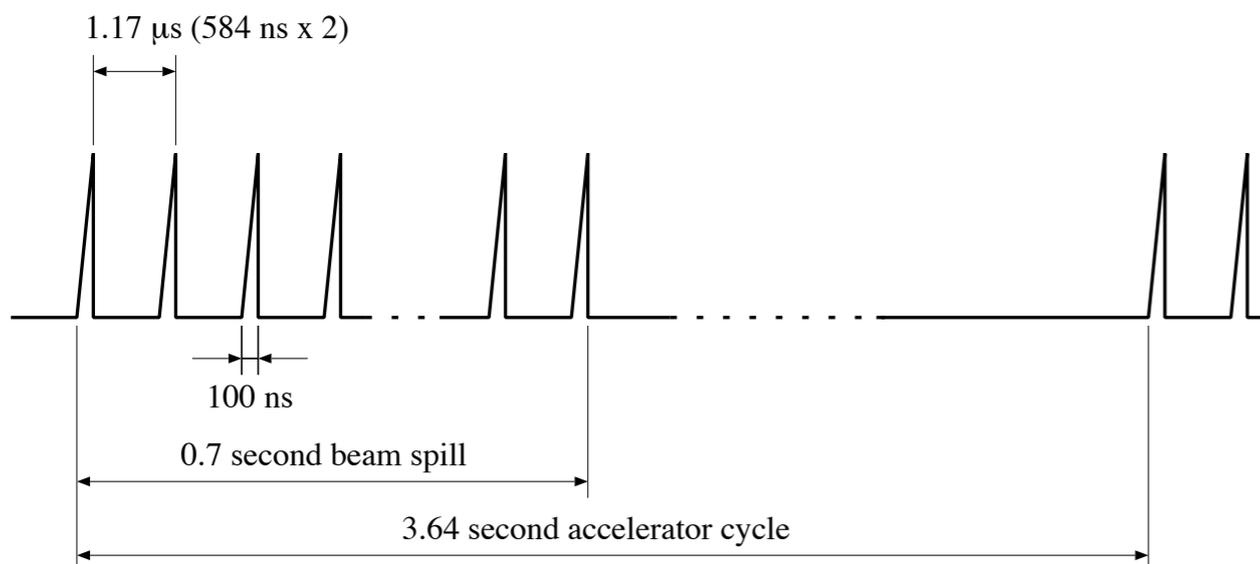
COMET
Exp. Area

Hadron Experimental Hall

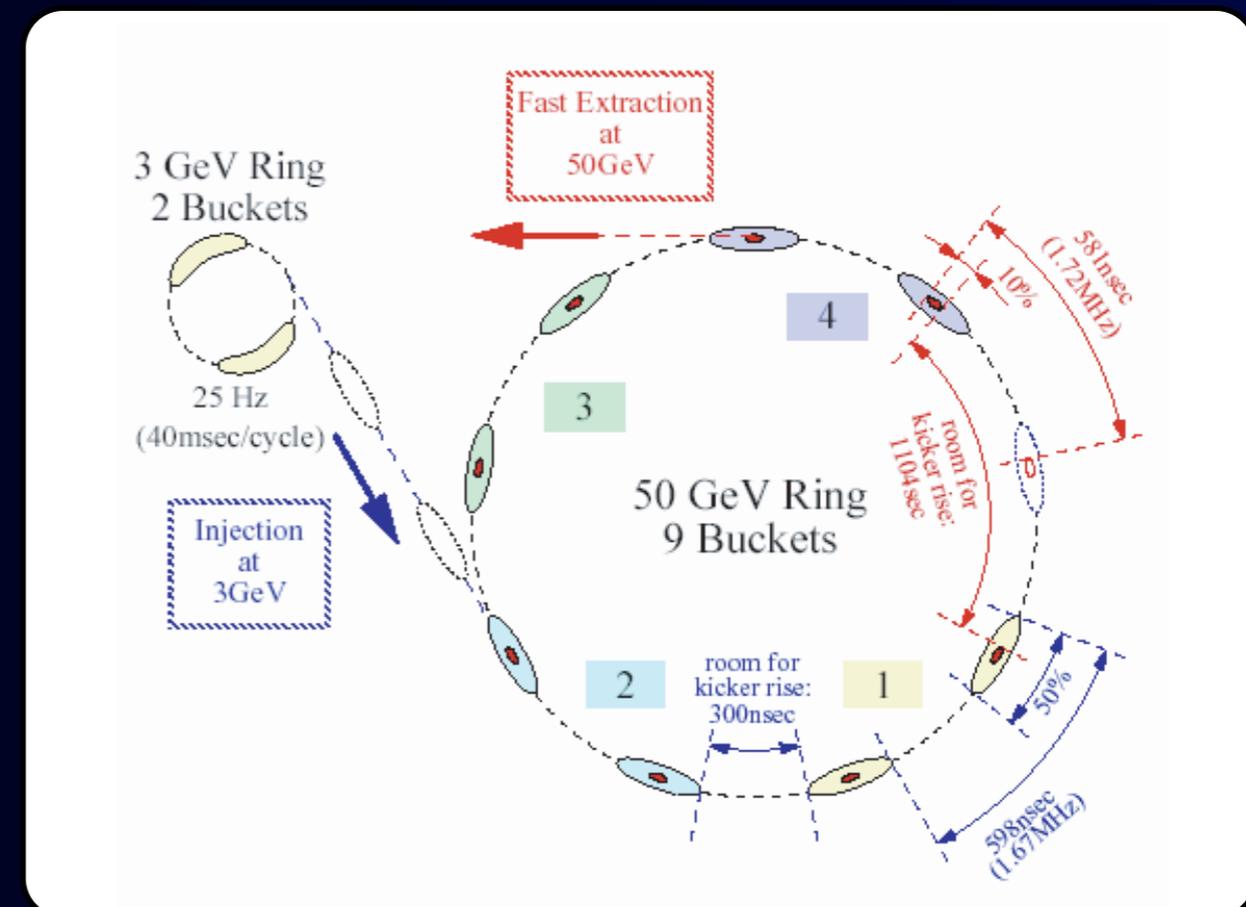


Proton Beam at J-PARC

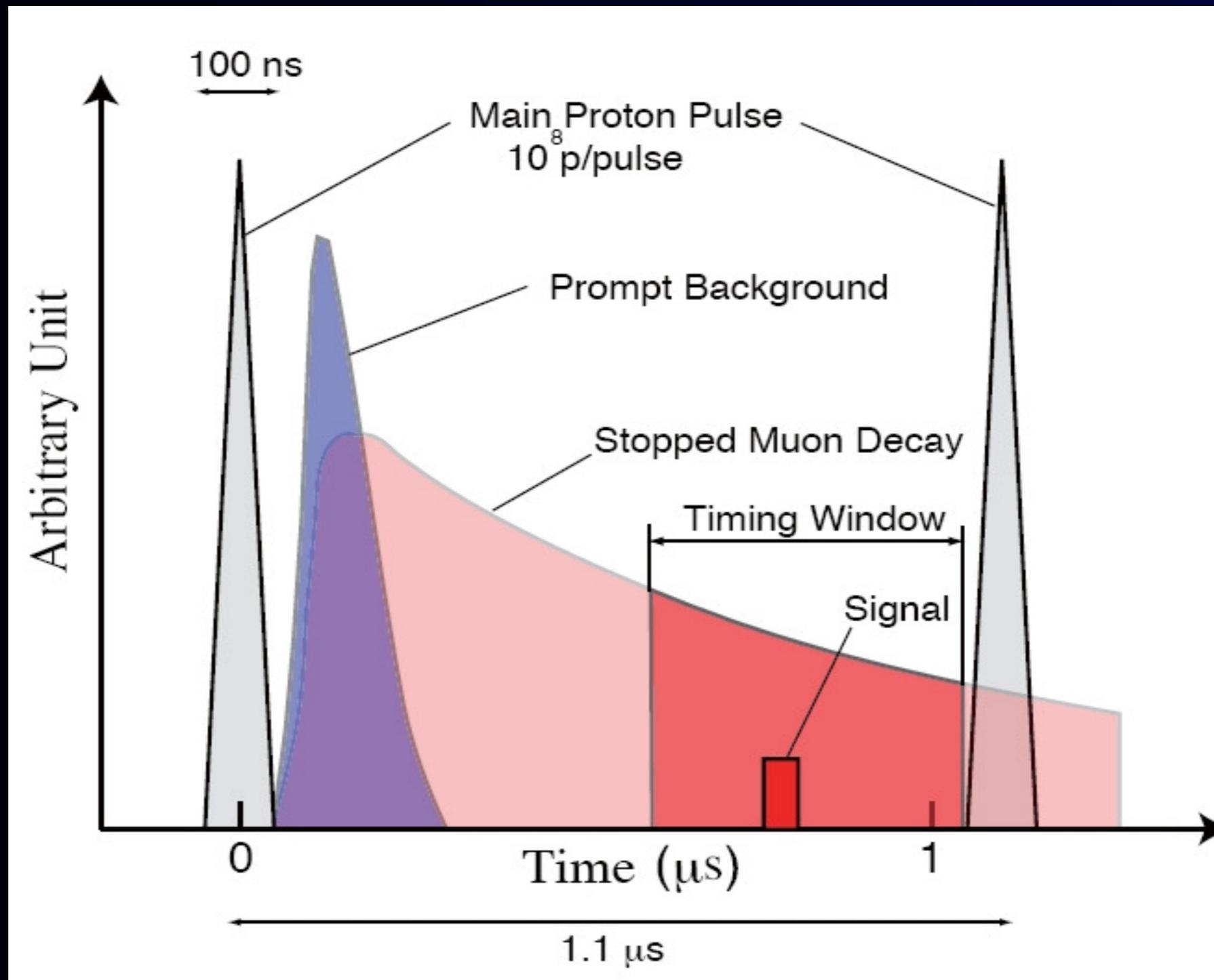
- A pulsed proton beam is needed to reject beam-related prompt background.
- Time structure required for proton beams.
 - Pulse separation is $\sim 1\mu\text{sec}$ or more (muon lifetime).
 - Narrow pulse width ($<100\text{ nsec}$)



- Pulsed beam from slow extraction.
 - fill every other rf buckets with protons and make slow extraction
 - spill length (flat top) ~ 0.7



Proton Beam for COMET



Muon Beam

Charged Particle Trajectory in Curved Solenoids



- A center of helical trajectory of charged particles in a curved solenoidal field is drifted by

$$D = \frac{p}{qB} \theta_{bend} \frac{1}{2} \left(\cos \theta + \frac{1}{\cos \theta} \right)$$

D : drift distance

B : Solenoid field

θ_{bend} : Bending angle of the solenoid channel

p : Momentum of the particle

q : Charge of the particle

θ : $\text{atan}(P_T/P_L)$

- This can be used for charge and momentum selection.

- This drift can be compensated by an auxiliary field parallel to the drift direction given by

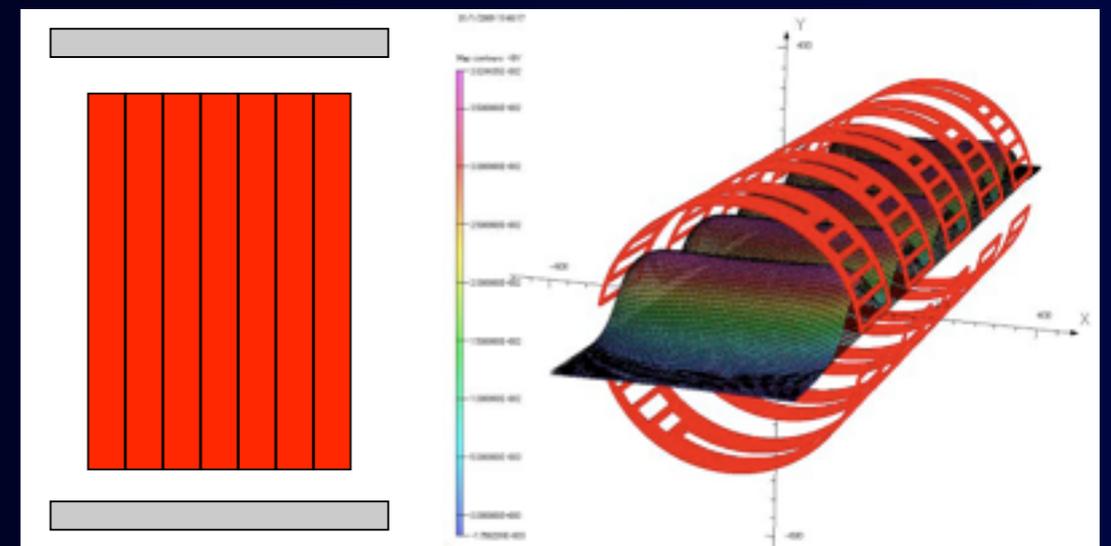
$$B_{comp} = \frac{p}{qr} \frac{1}{2} \left(\cos \theta + \frac{1}{\cos \theta} \right)$$

p : Momentum of the particle

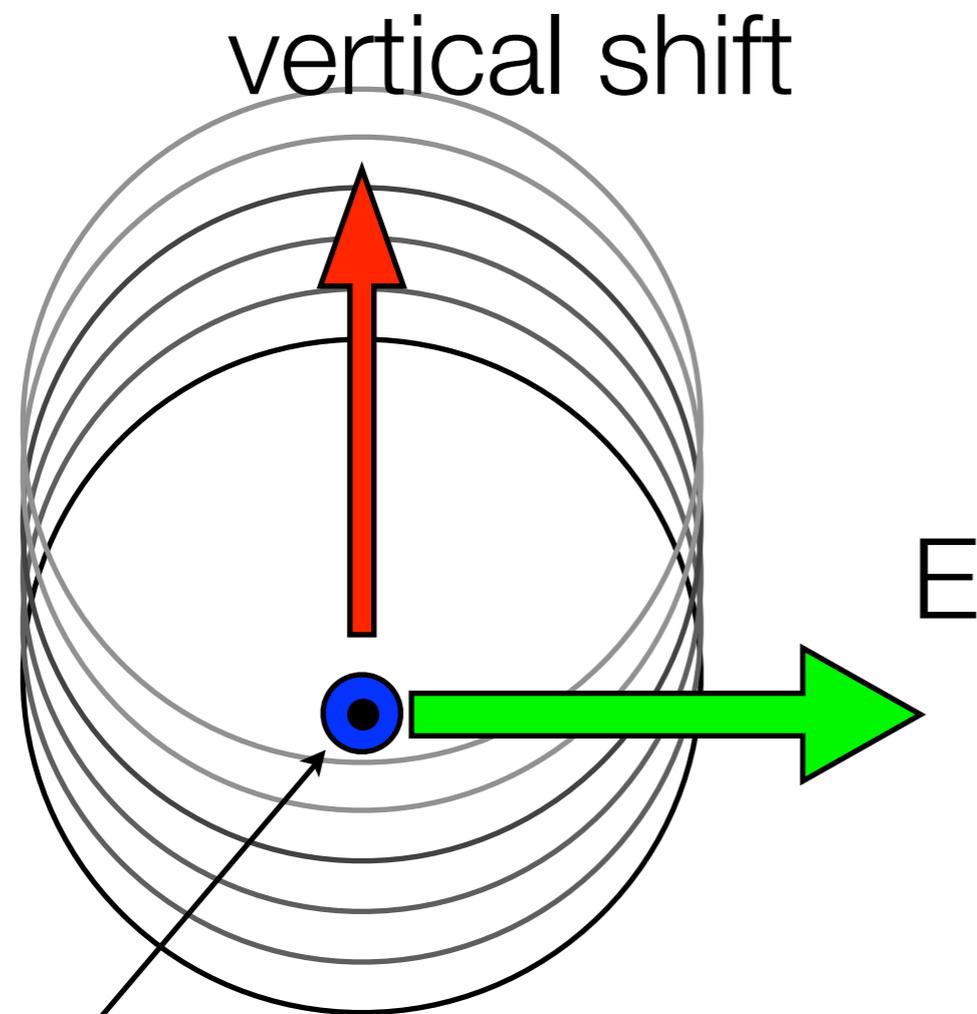
q : Charge of the particle

r : Major radius of the solenoid

θ : $\text{atan}(P_T/P_L)$



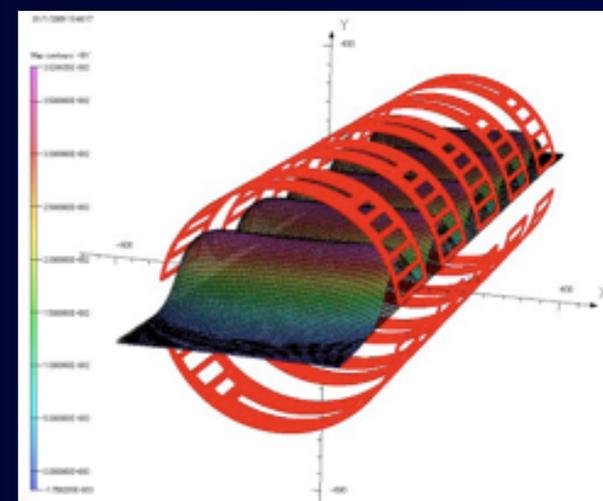
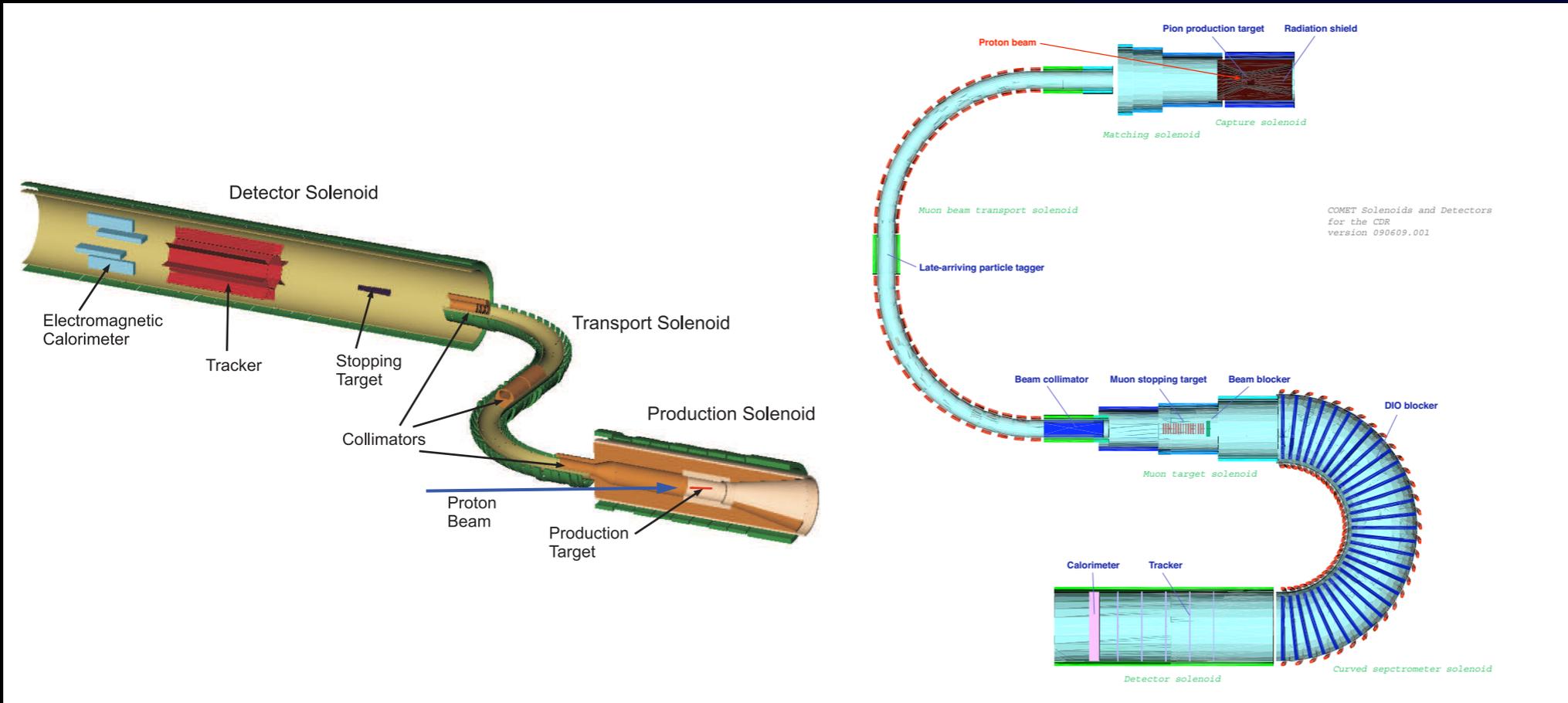
EM Physics for Particle Trajectories in Toroidal Magnetic Field



B (perpendicular to screen)

- For helical trajectory in a curved mag. field, a centrifugal force gives E in the radial direction.
- To compensate a vertical shift, an electric field in the opposite direction shall be applied, or a vertical mag. field that produces the desired electric field by $v \times B$, can be applied.

Mu2e vs. COMET

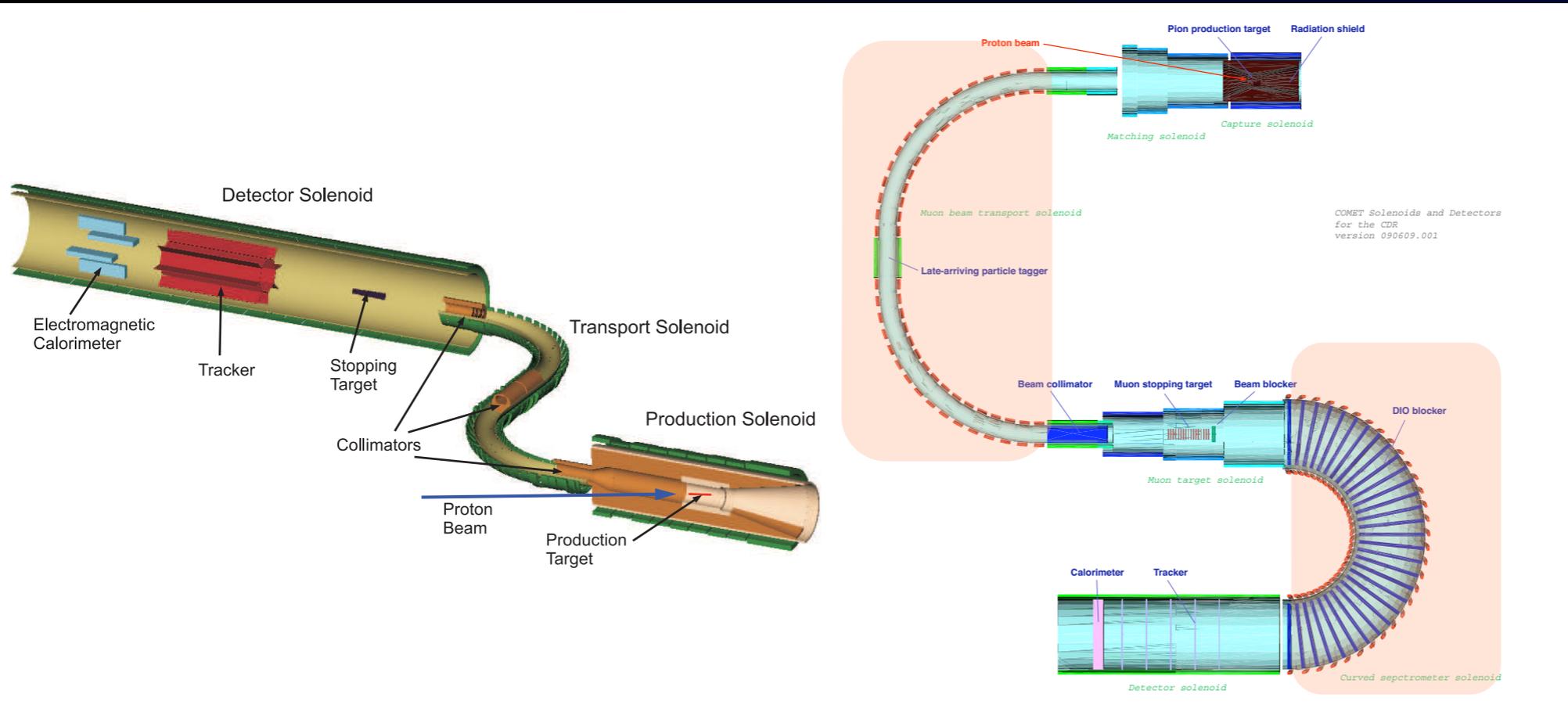


Dipole Coils

COMET curved solenoids have dipole coils on top of the solenoids, to keep muons with momentum of interest in the bending plane.

	Mu2e	COMET
muon beam line	2x 90° bends (opposite direction)	2x 90° bend (same direction)
electron spectrometer	straight solenoid	curved solenoid

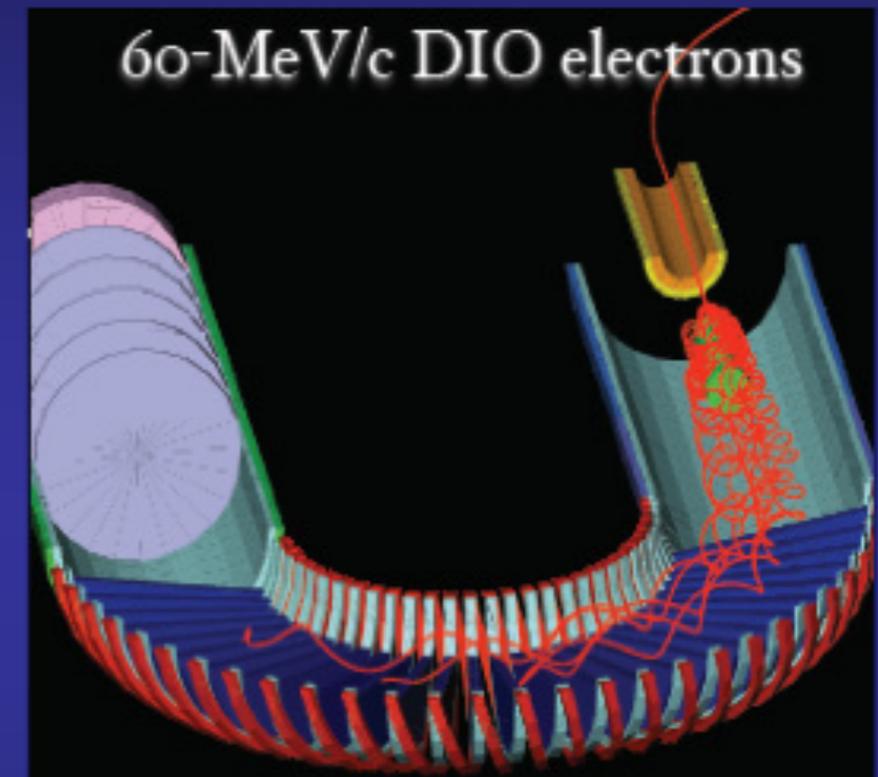
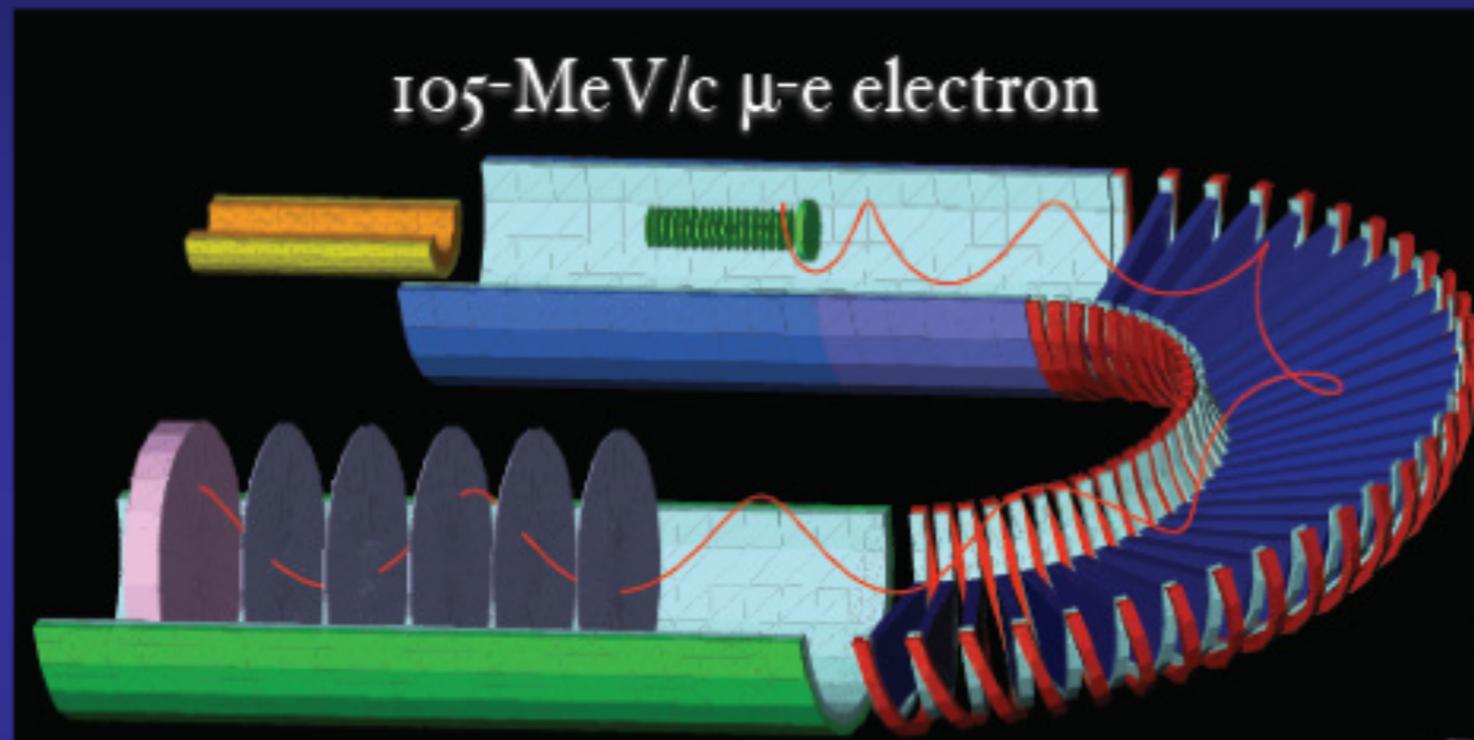
Mu2e vs. COMET



- Select low momentum muons
- eliminate muon decay in flight
- Selection of 100 MeV electrons
- eliminate protons from nuclear muon capture.
- eliminate low energy events to make the detector quiet.

	Mu2e	COMET
muon beam line	2x 90° bends (opposite direction)	2x 90° bend (same direction)
electron spectrometer	straight solenoid	curved solenoid

Electron Spectrometer



- One component that is not included in the Mu2e design.
- 1T solenoid with additional 0.17T dipole field.
- Vertical dispersion of toroidal field allows electrons with $P < 60 \text{ MeV}/c$ to be removed.
 - reduces rate in tracker to $\sim 1 \text{ kHz}$.

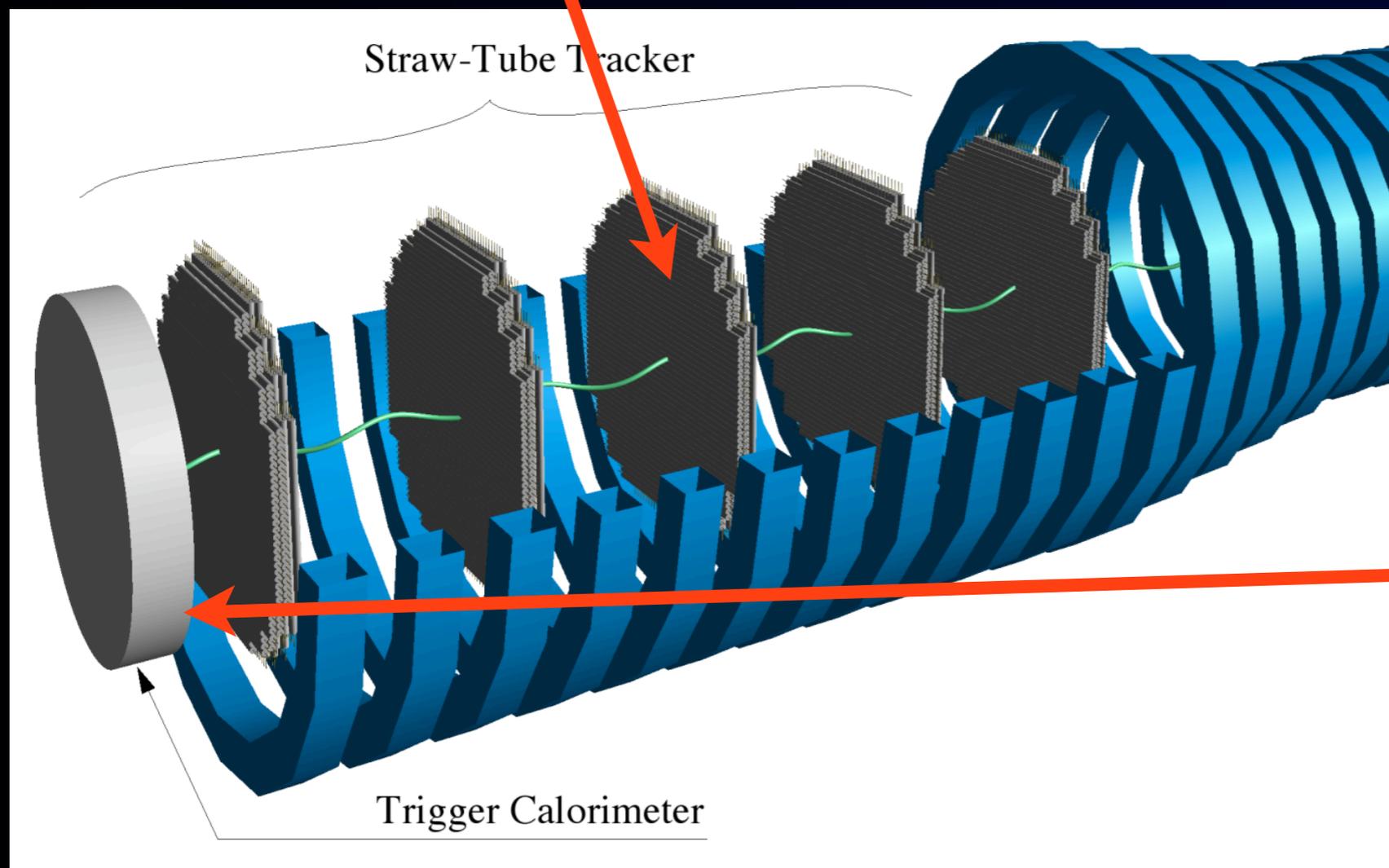
Electron Detection

Electron Tracker to measure electron momentum

- work in vacuum and under a magnetic field.
- Straw tube chambers
 - Straw tubes of 25 μ m thick, 5 mm diameter.
 - five plane has 2 views (x and y) with 2 layers per view.
- Planar drift chambers

Under a solenoidal magnetic field of 1 Tesla.

In vacuum to reduce multiple scattering.



- Electron calorimeter to measure electron energy, make triggers and give additional hit position.
- Candidate are LYSO, GSO
 - MPPC or APD readout

Sensitivity and Backgrounds

Signal Sensitivity (preliminary) - 2×10^7 sec

- Single event sensitivity

$$B(\mu^- + Al \rightarrow e^- + Al) \sim \frac{1}{N_\mu \cdot f_{cap} \cdot A_e},$$

- N_μ is a number of stopping muons in the muon stopping target. It is 2×10^{18} muons.
- f_{cap} is a fraction of muon capture, which is 0.6 for aluminum.
- A_e is the detector acceptance, which is 0.04.

total protons	8.5×10^{18}
muon transport efficiency	0.008
muon stopping efficiency	0.3
# of stopped muons	2.0×10^{18}

$$B(\mu^- + Al \rightarrow e^- + Al) = 2.6 \times 10^{-17}$$

$$B(\mu^- + Al \rightarrow e^- + Al) < 6 \times 10^{-17} \quad (90\% C.L.)$$

Background Rates

Table 1.1. Summary of Estimated Backgrounds

Radiative Pion Capture	0.05
Beam Electrons	< 0.1 [‡]
Muon Decay in Flight	< 0.0002
Pion Decay in Flight	< 0.0001
Neutron Induced	0.024
Delayed-Pion Radiative Capture	0.002
Anti-proton Induced	0.007
Muon Decay in Orbit	0.15
Radiative Muon Capture	< 0.001
μ^- Capt. w/ n Emission	< 0.001
μ^- Capt. w/ Charged Part. Emission	< 0.001
Cosmic Ray Muons	0.002
Electrons from Cosmic Ray Muons	0.002
Total	0.34

[‡] Monte Carlo statistics limited.

beam-related prompt
backgrounds

beam-related delayed
backgrounds

intrinsic physics
backgrounds

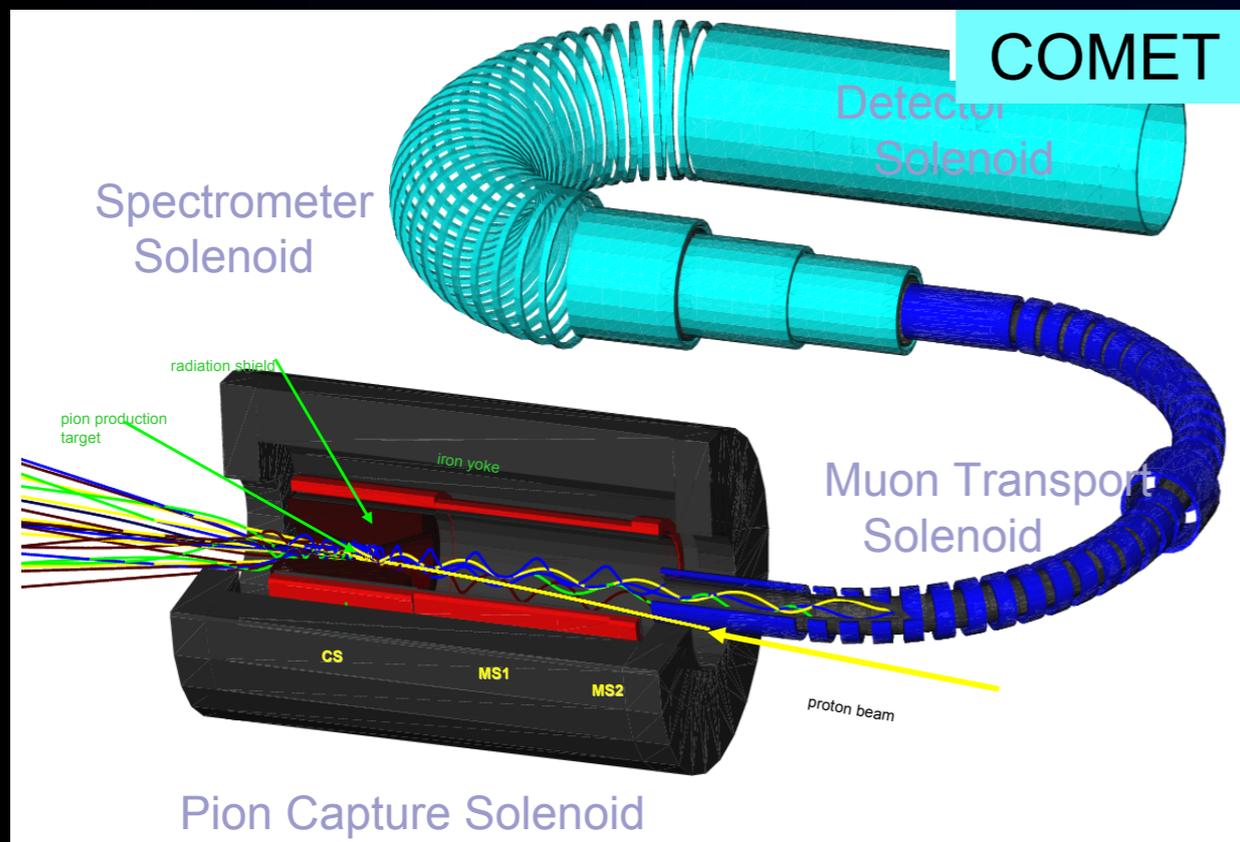
cosmic-ray and other
backgrounds

Expected background events are about 0.34.

COMET Milestones



R&D Milestones for μ -e conversion



$$B(\mu^- + Al \rightarrow e^- + Al) < 10^{-16}$$

single event sensitivity: 2.6×10^{-17}

1 Reduction of Backgrounds

Beam pulsing

measurement is done between beam pulses to reduce beam related backgrounds. And proton beam extinction of $< 10^{-9}$ is required.

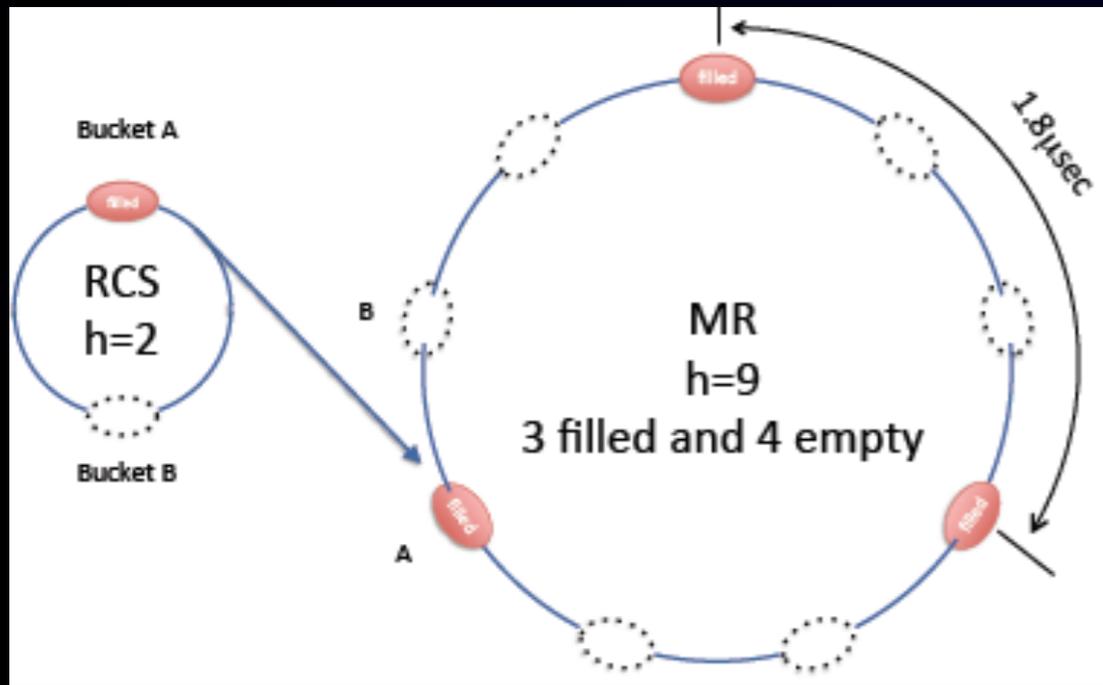
2 Increase of Muon Intensity

Pion capture system $\times 10^3$

high field superconducting solenoid magnets surrounding a pion production target

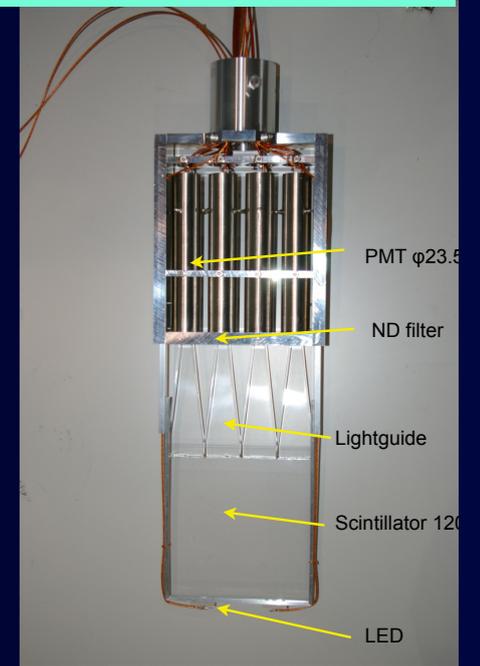
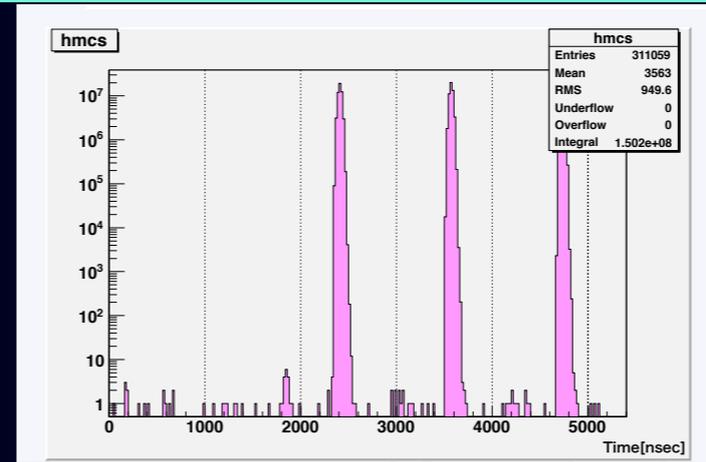
1

Proton Extinction Measurements at J-PARC



Measured at abort beamline (2010)

Measured at secondary beamline (2010)



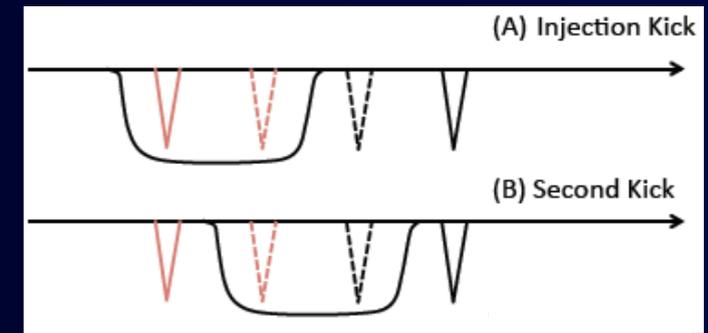
J-PARC MR proton extinction

$\sim O(10^{-7})$

Single Bunch Kicking

Tested at the abort (2010)

x additional $O(10^{-6})$



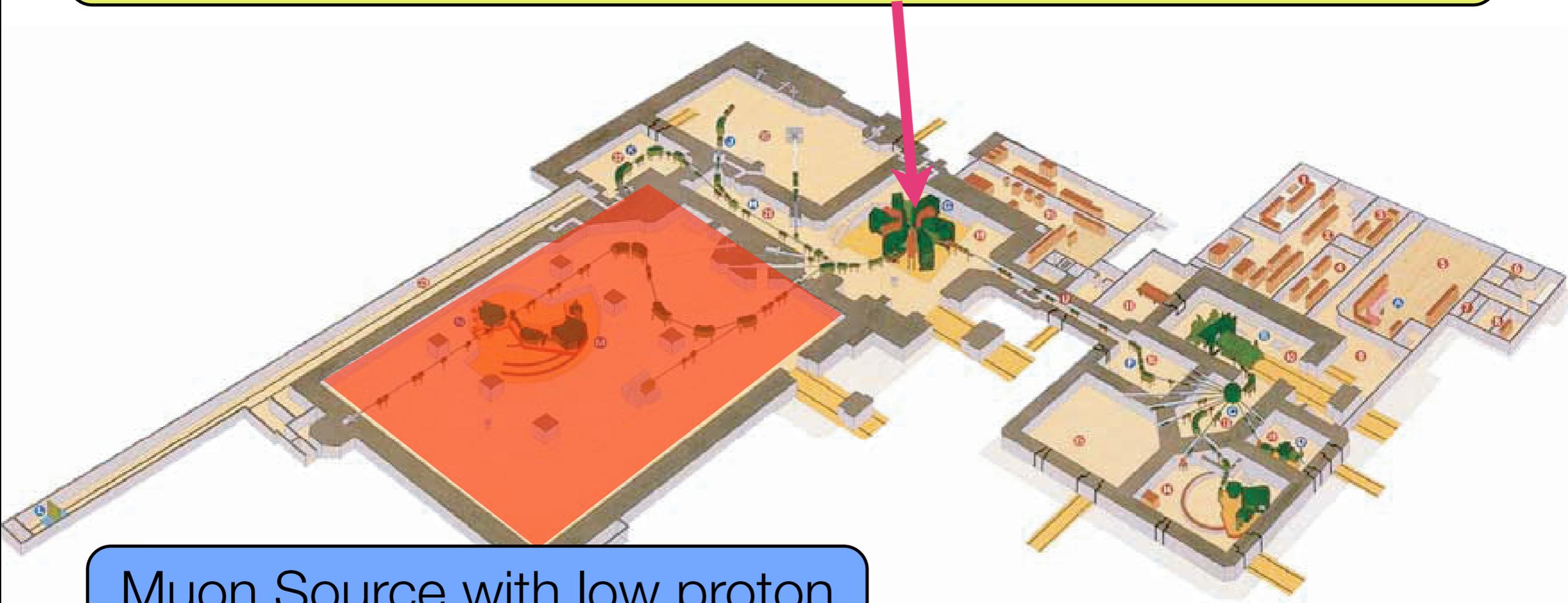
COMET is confident to achieve proton extinction of $<O(10^{-9})$.

2

Research Center for Nuclear Physics (RCNP), Osaka University

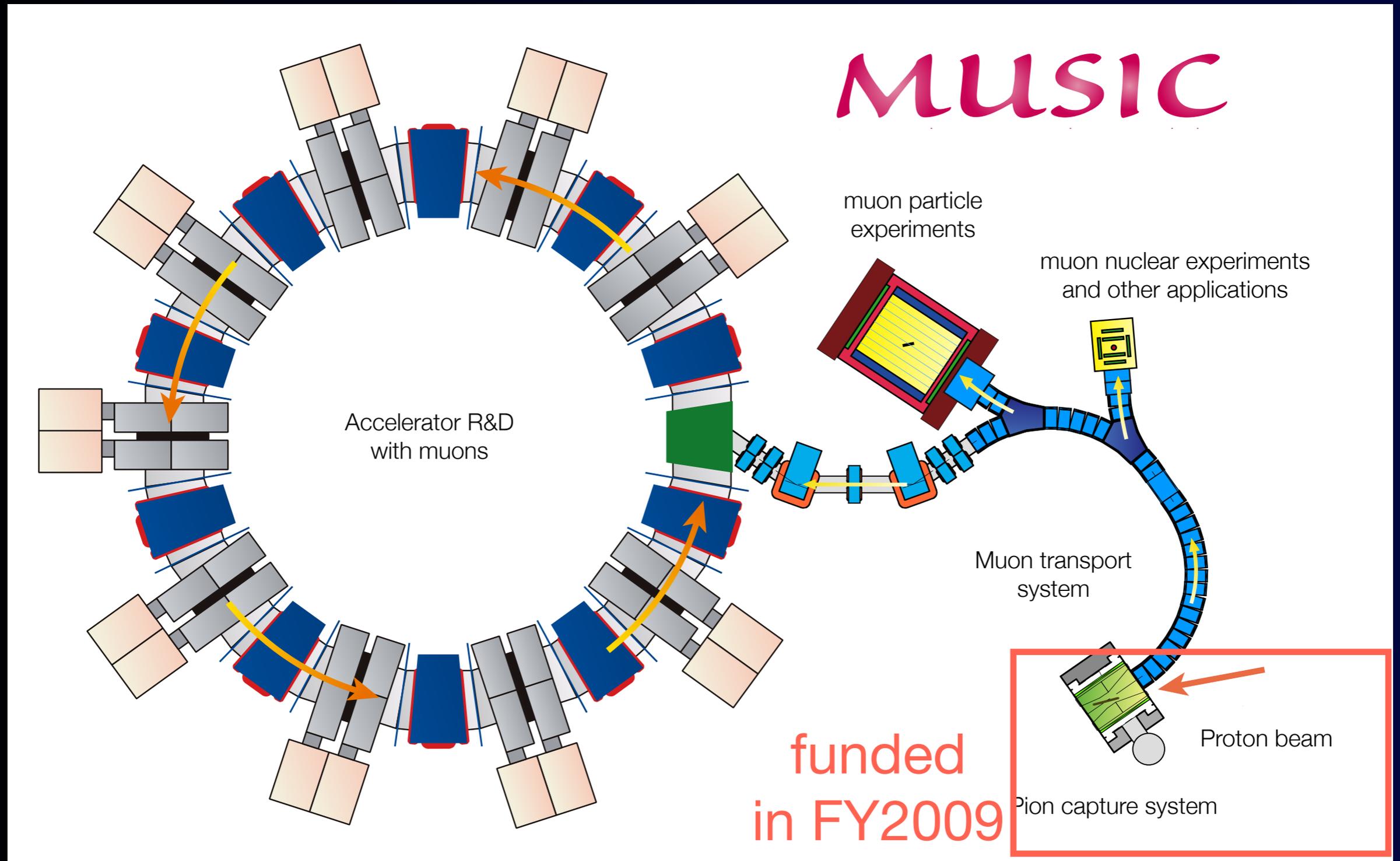


Research Center for Nuclear Physics (RCNP), Osaka University has a cyclotron of 400 MeV with 1 microA. The energy is above pion threshold.



Muon Source with low proton power at Osaka U.?

MuSIC (=Muon Science Innovative Channel)

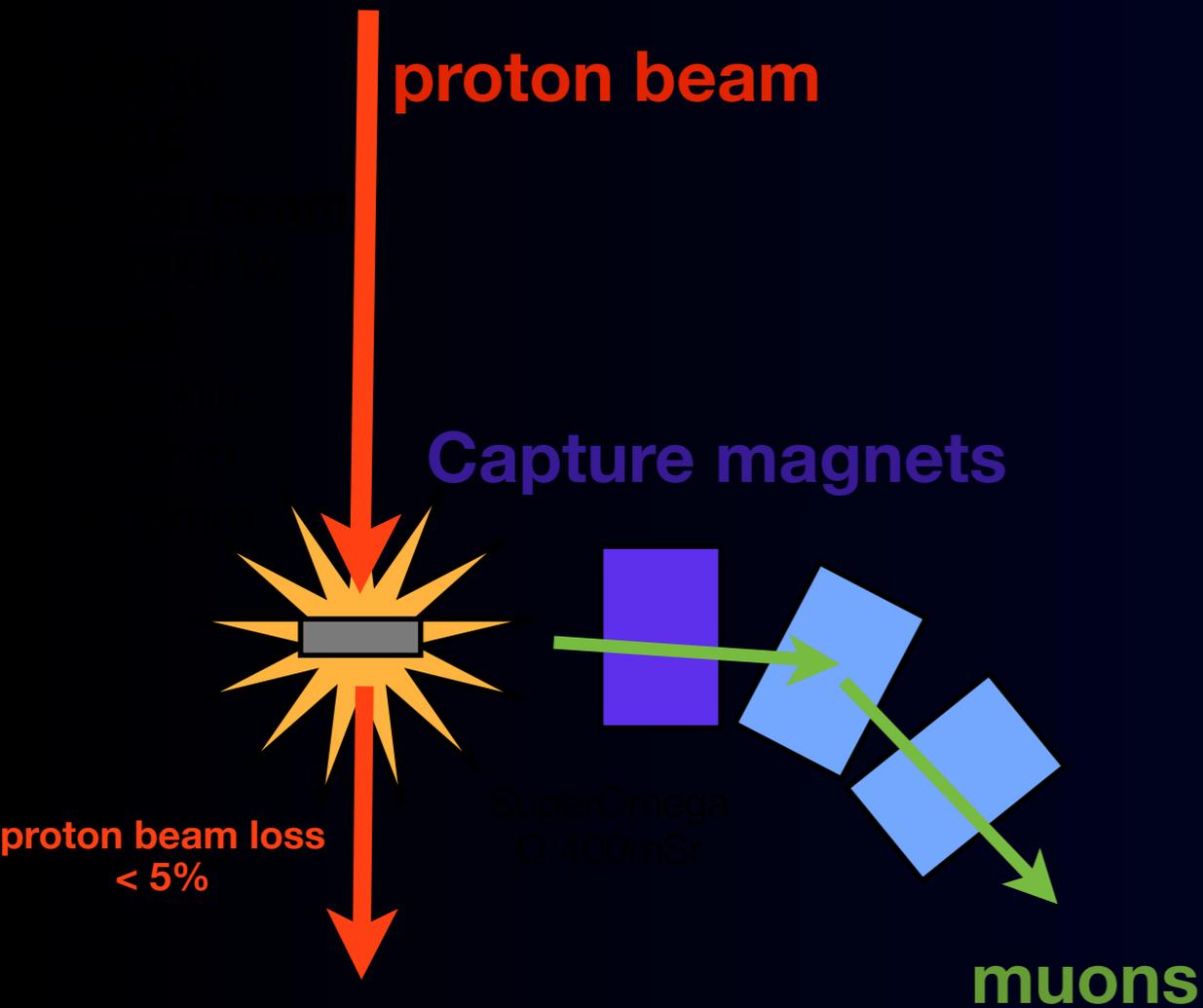


Production and Collection of Pions and Muons

Conventional muon beam line

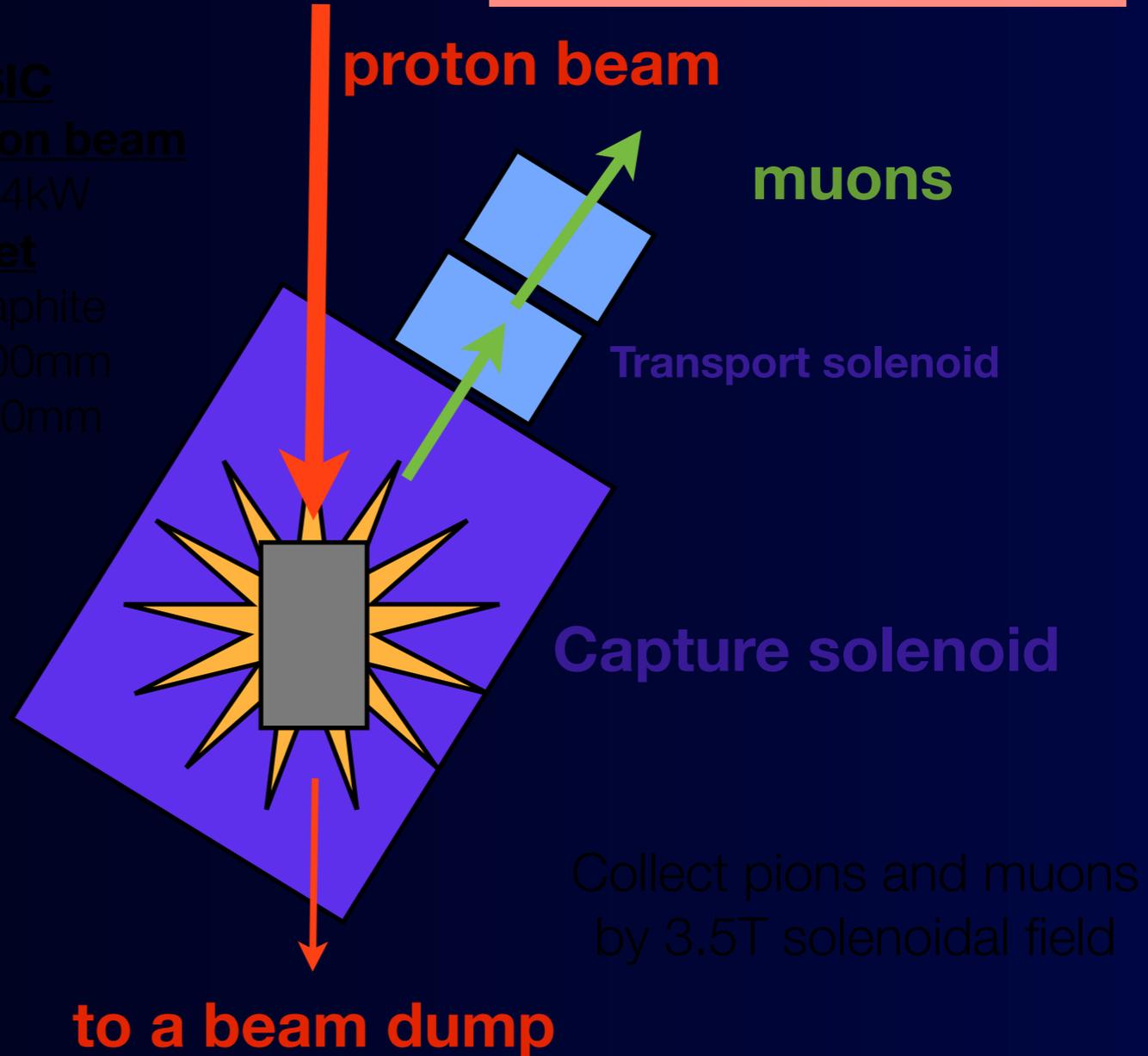
Much efficient

MuSIC, COMET, PRISM, Neutrino factory, Muon collider



MuSIC
proton beam

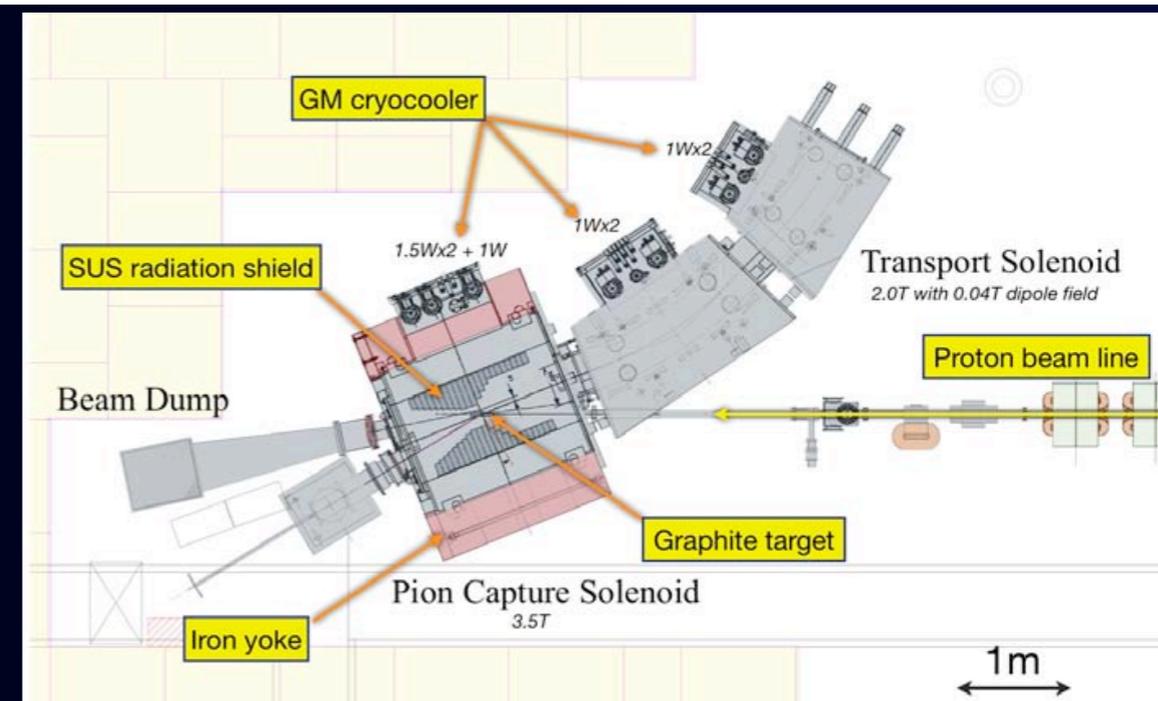
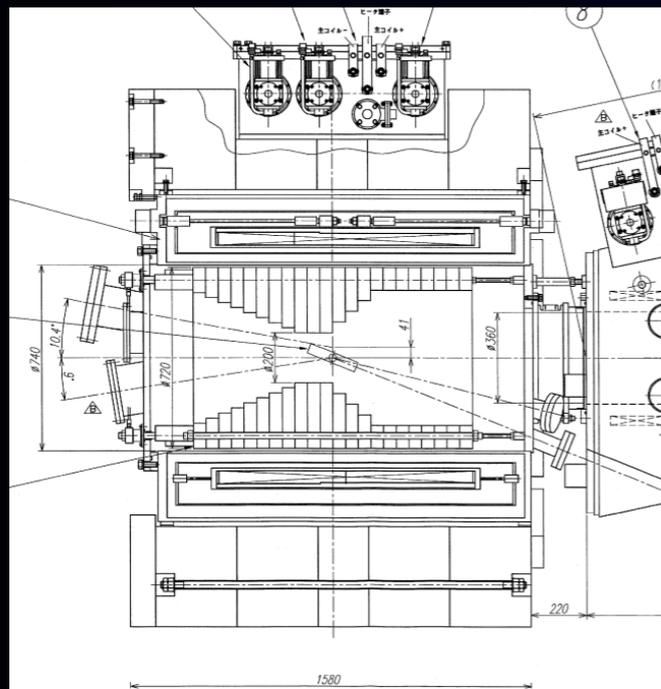
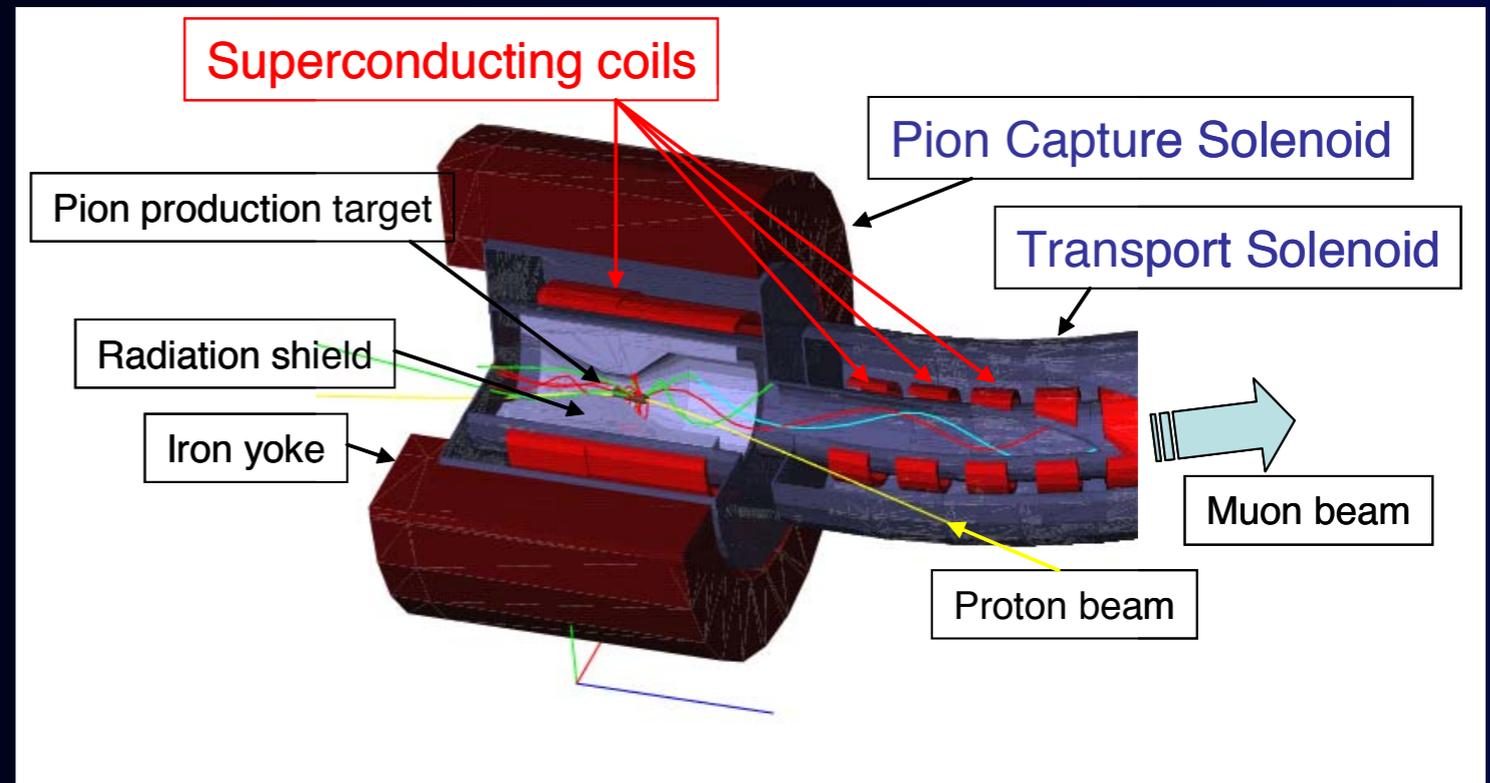
-0.4kW
target
graphite
t200mm
 $\phi 40\text{mm}$

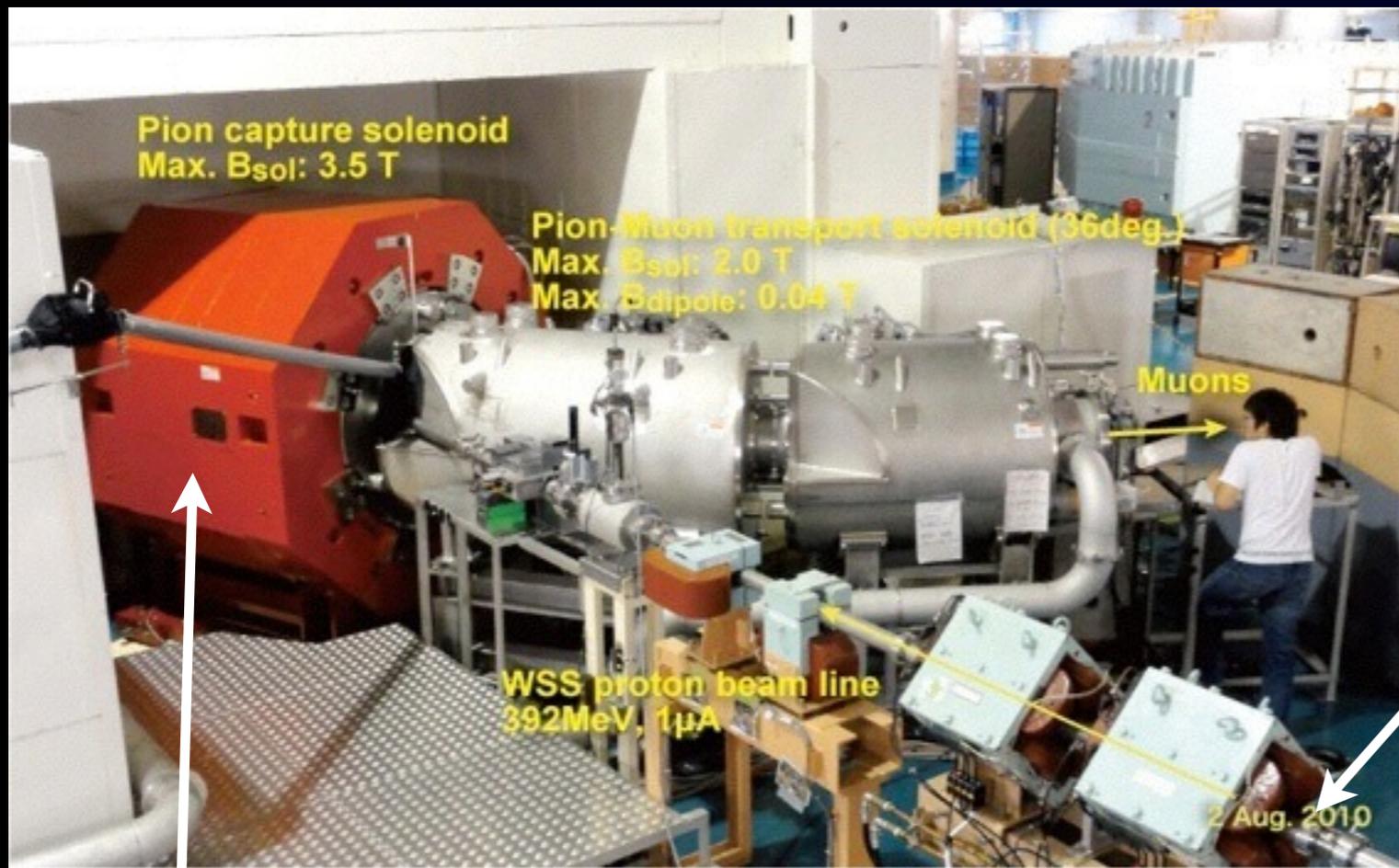


Large solid angle & thick target

Pion Capture System at MuSIC@Osaka-U

- Pion Capture SC Solenoid :
 - 3.5 T at central
 - diameter 740mm
 - SUS radiation shield
- Transport SC solenoids
 - 2 T magnetic field
 - 8 thin solenoids
- Graphite target for pion production

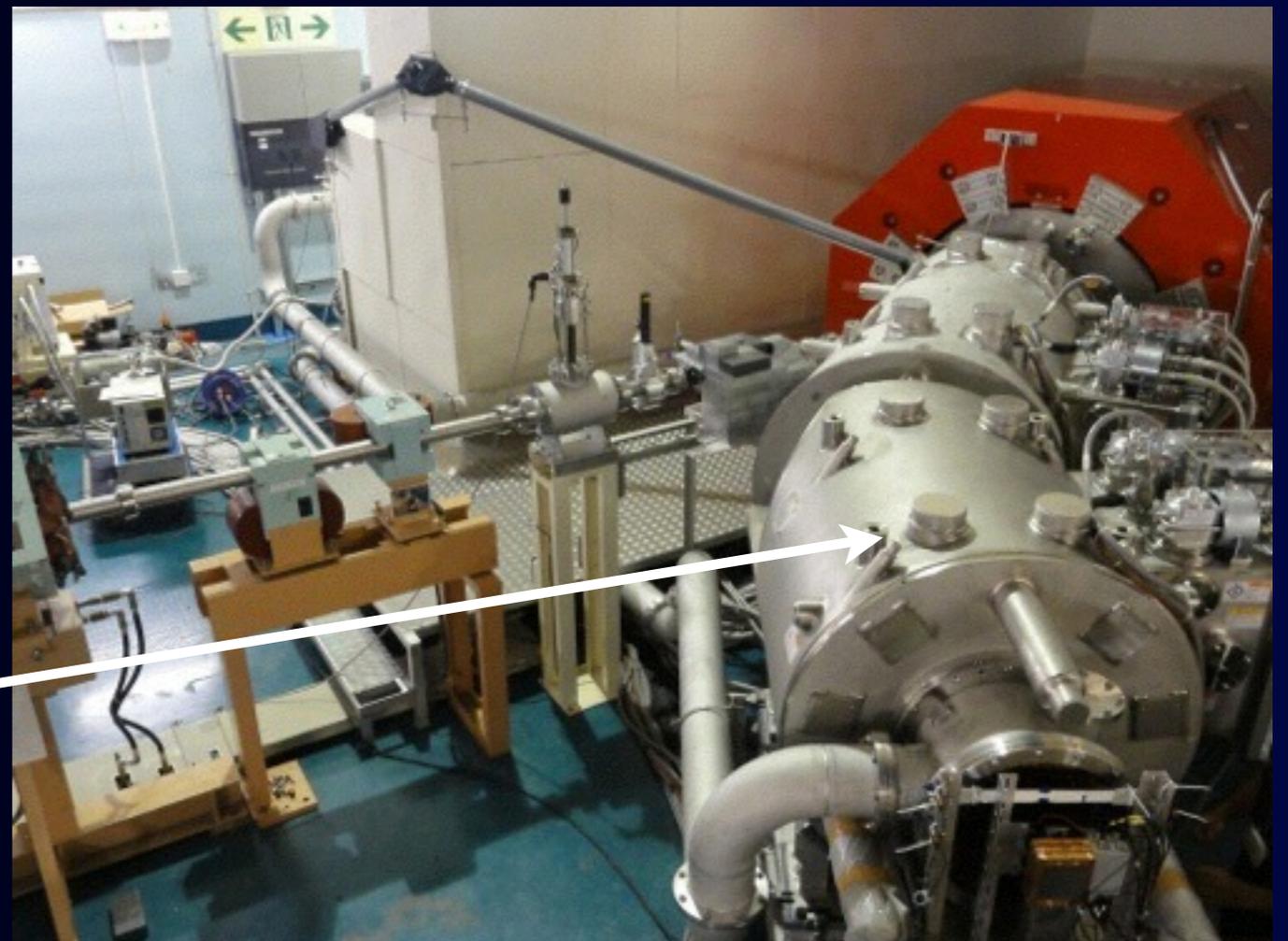




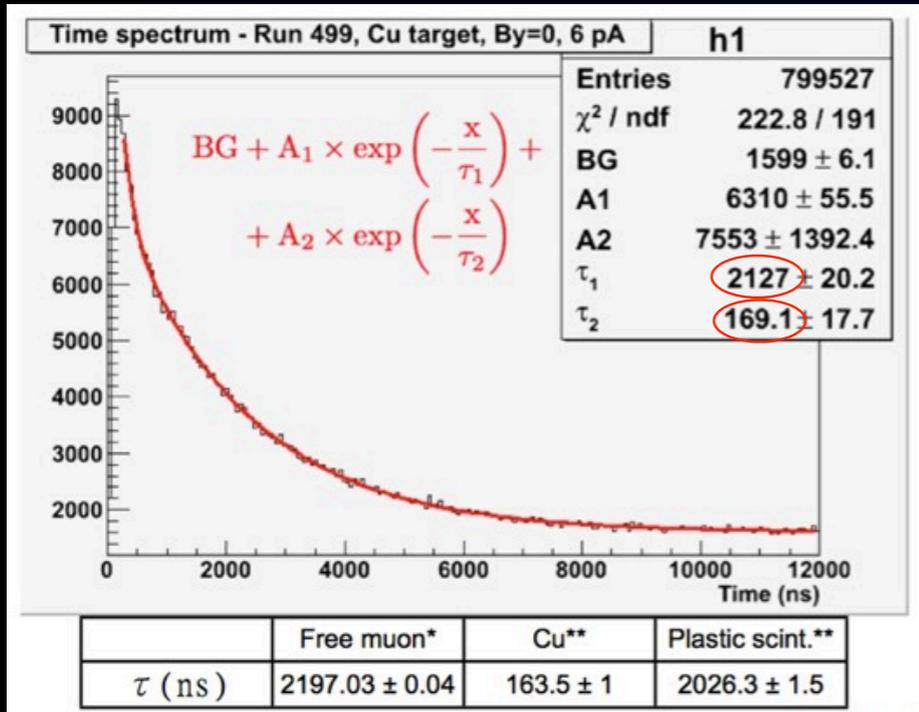
proton beam line

pion capture
superconducting
solenoid

muon transport
superconducting
solenoid



MuSIC Beam Test in 2011



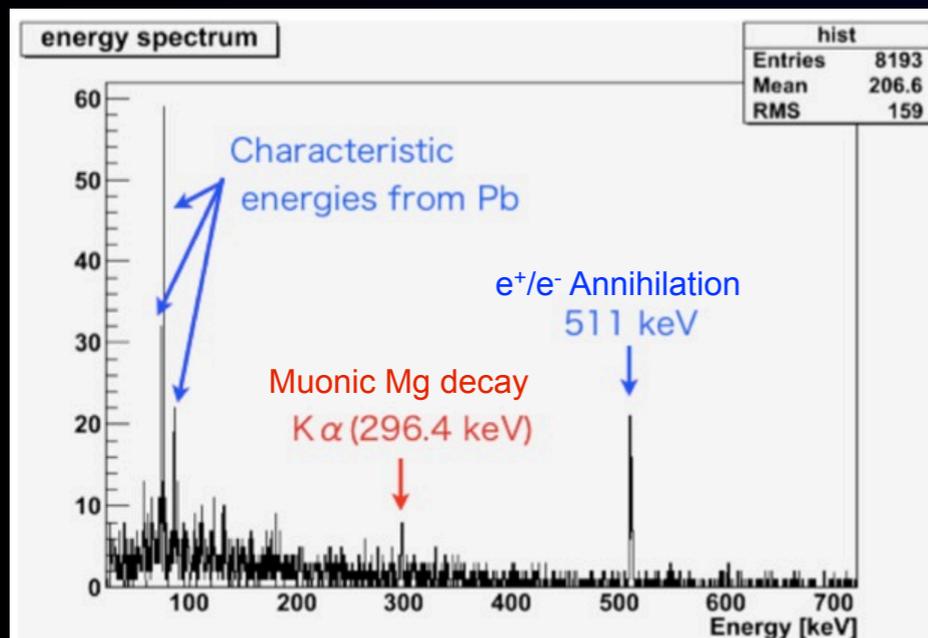
preliminary

MuSIC muon yields

μ^+ : $3 \times 10^8 / \text{s}$ for 400W

μ^- : $1 \times 10^8 / \text{s}$ for 400W

cf. $10^8 / \text{s}$ for 1MW @PSI
Req. of $\times 10^3$ achieved...



Great opportunities to carry out muon particle physics from NOW!

Measurements on June 21, 2011 (62 pA)

Common R&D with COMET and Mu2e

Superconducting Magnet R&D (2010~)

...through the US-Japan Program

R&D of solenoid coils with aluminum-stabilized superconductors

Japan



Prototype coil of aluminum-stabilized superconductors were wound in Japan and sent to FNAL (2010-2012).

U.S.

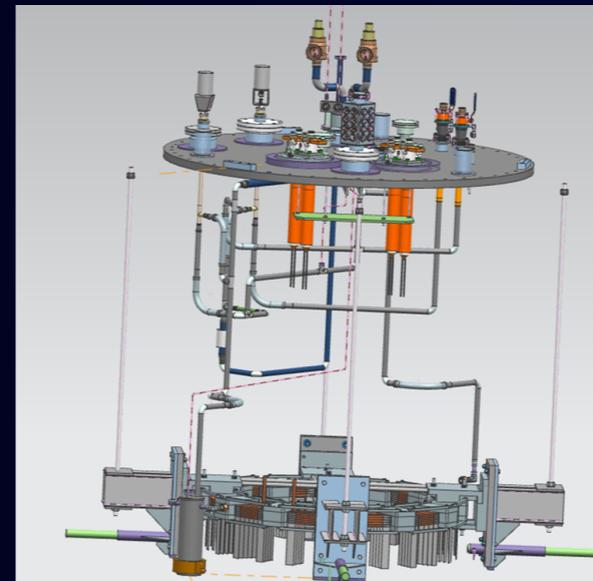


Figure 2. Mechanical support system for the prototype coil



Figure 3. Photo of test cryostat in its Central Helium Liquifier Location

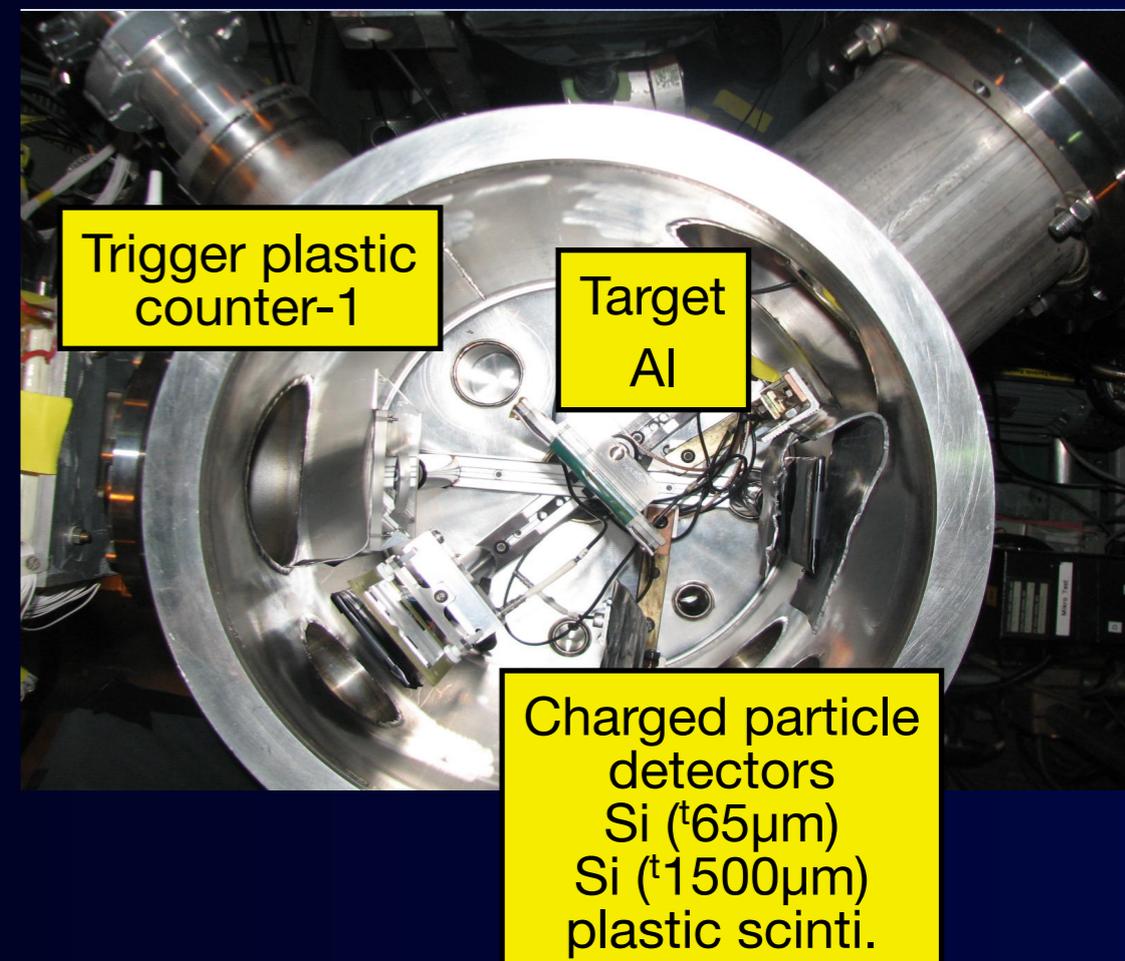
Indirect cooling test bench was prepared at FNAL to test the prototype coil (2013~, not yet?)

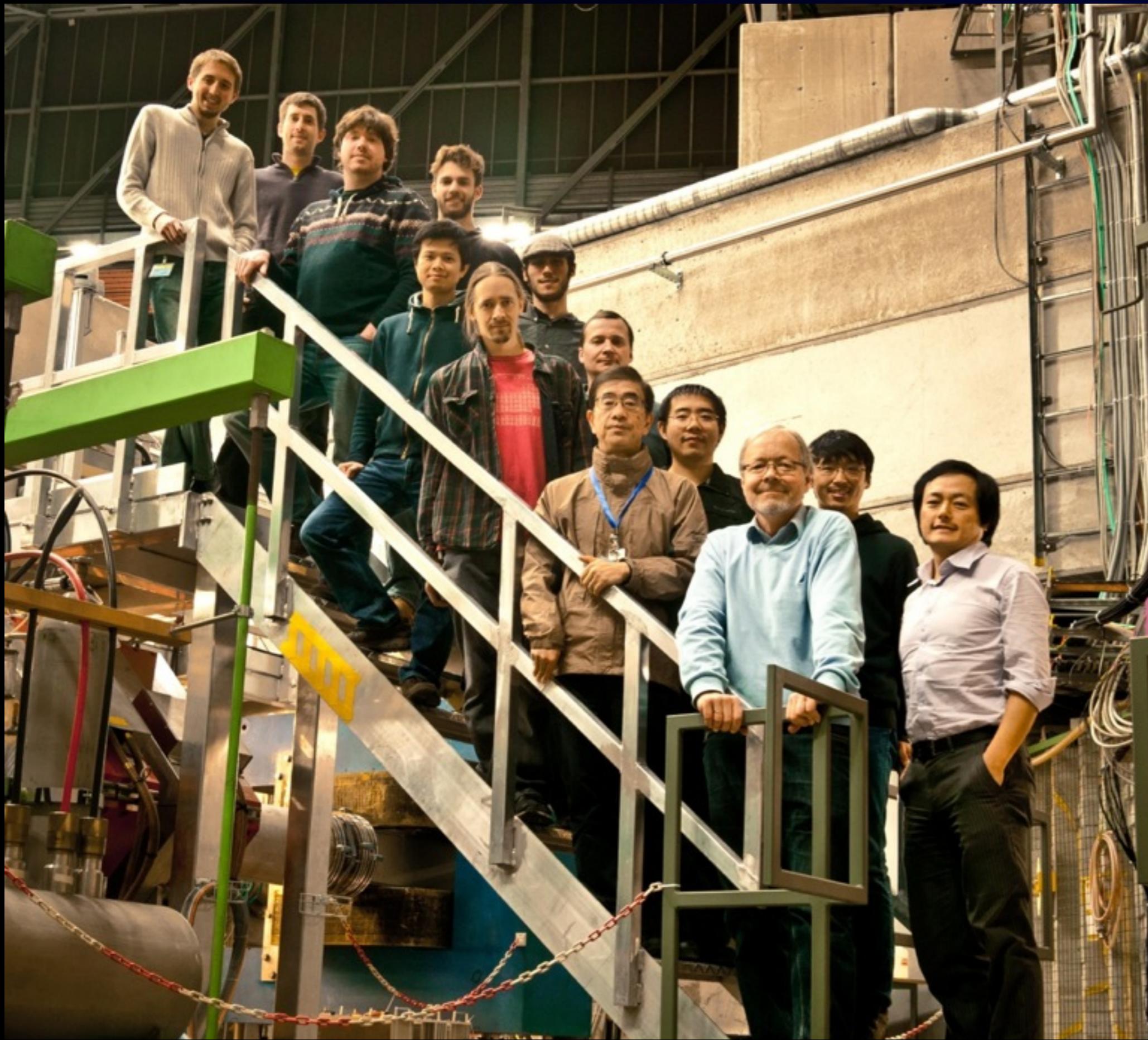
AlCap Experiment at PSI (2013~)

...through the US-Japan Program

Measurements of particle (proton) production after muon capture on Al.

- Proton emission rate after muon capture is important, since it determines single rates of tracking chambers if no charge selection is made before detection.
- That rate for aluminum has not been measured.
- **As a joint effort** of Mu2e and COMET, the **AlCap** experiment (P. Kammel (UW) and YK (Osaka), co-spokespersons) is being done at PSI in December, 2013.
- The measurements of neutron emission will be done in 2014?





AlCap@PSI

Dec. 13



AlCap@PSI

Dec. 13

-  COMET
-  Mu2e

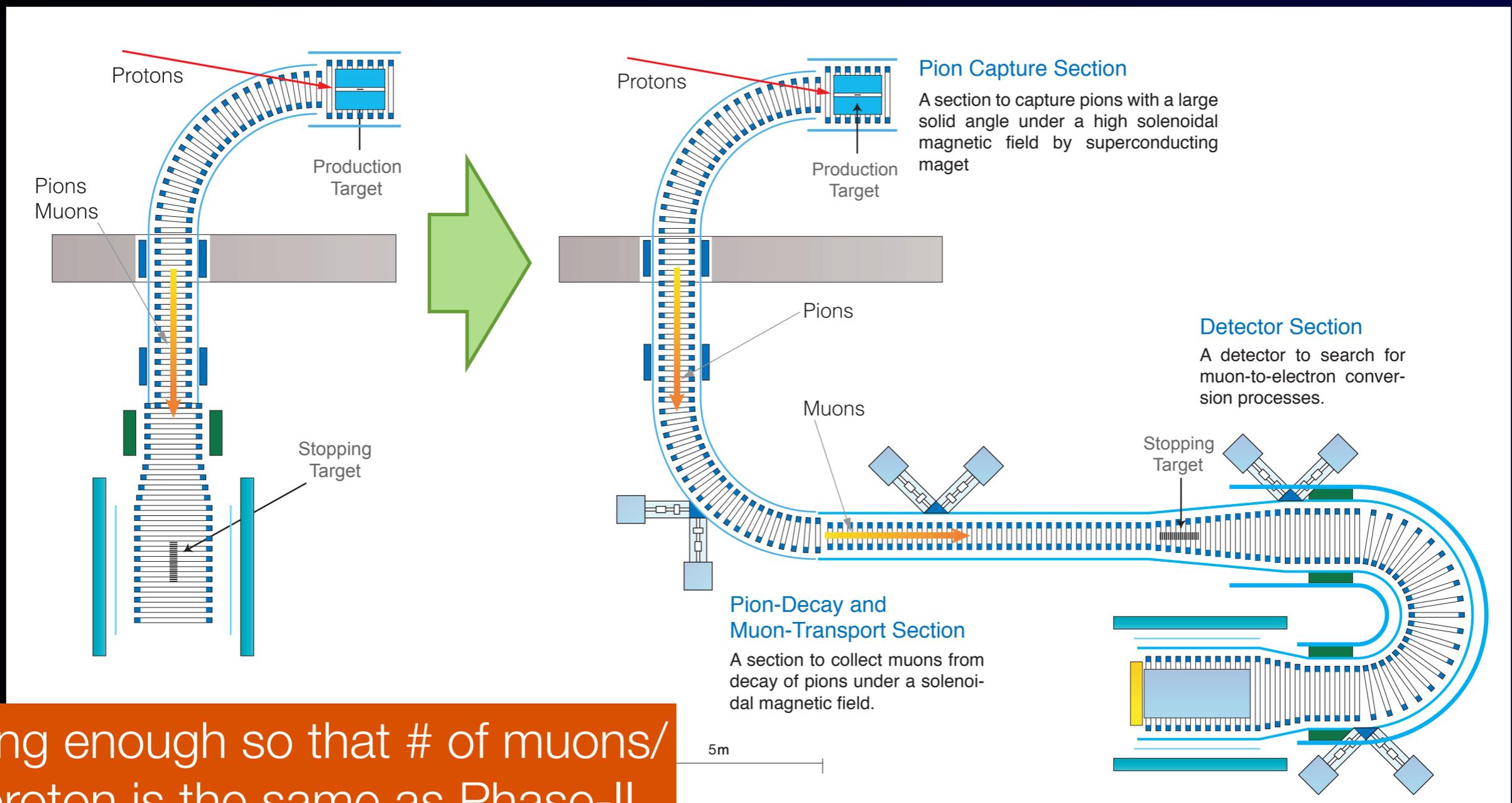
COMET Phase-I



COMET Staged Approach (2012~)

COMET Phase-I

COMET Phase-II



long enough so that # of muons/proton is the same as Phase-II.

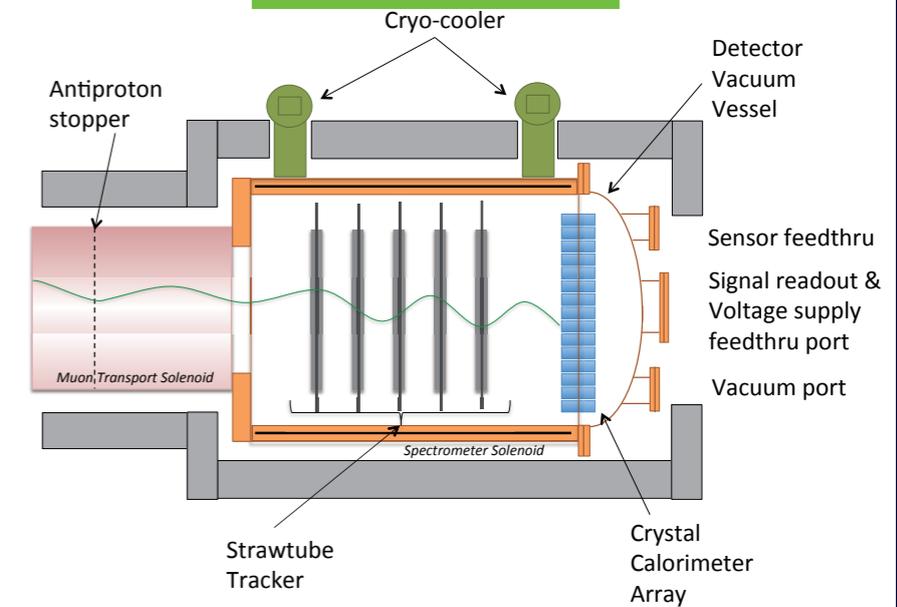
Goals of COMET Phase-I

1

Background Study for COMET Phase-I

direct measurement of potential backgrounds from other sources for the full COMET experiment
 actual COMET beamline construction

StrEcal

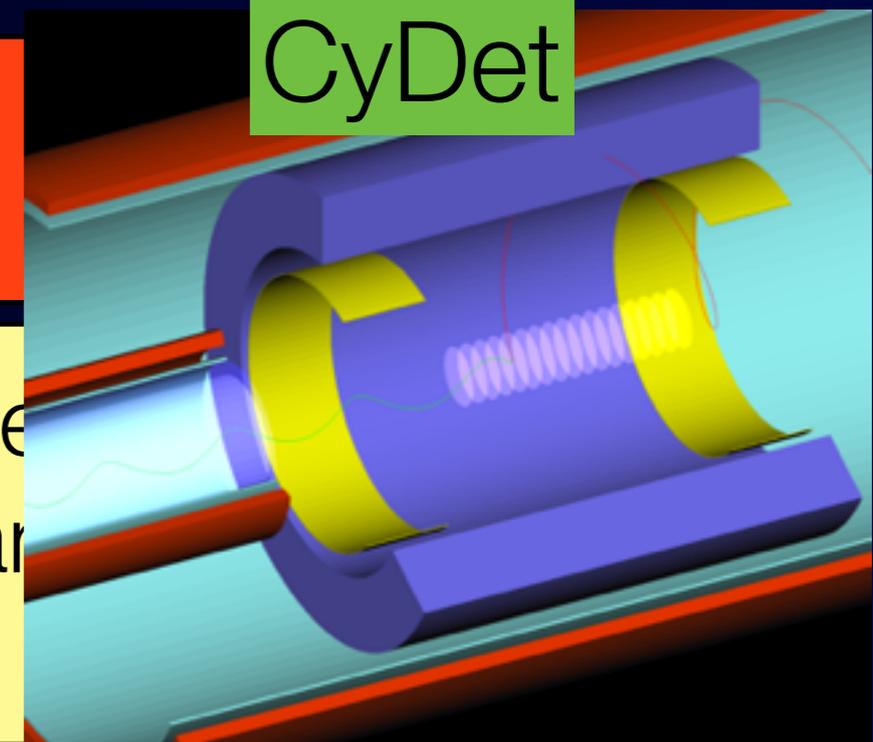


2

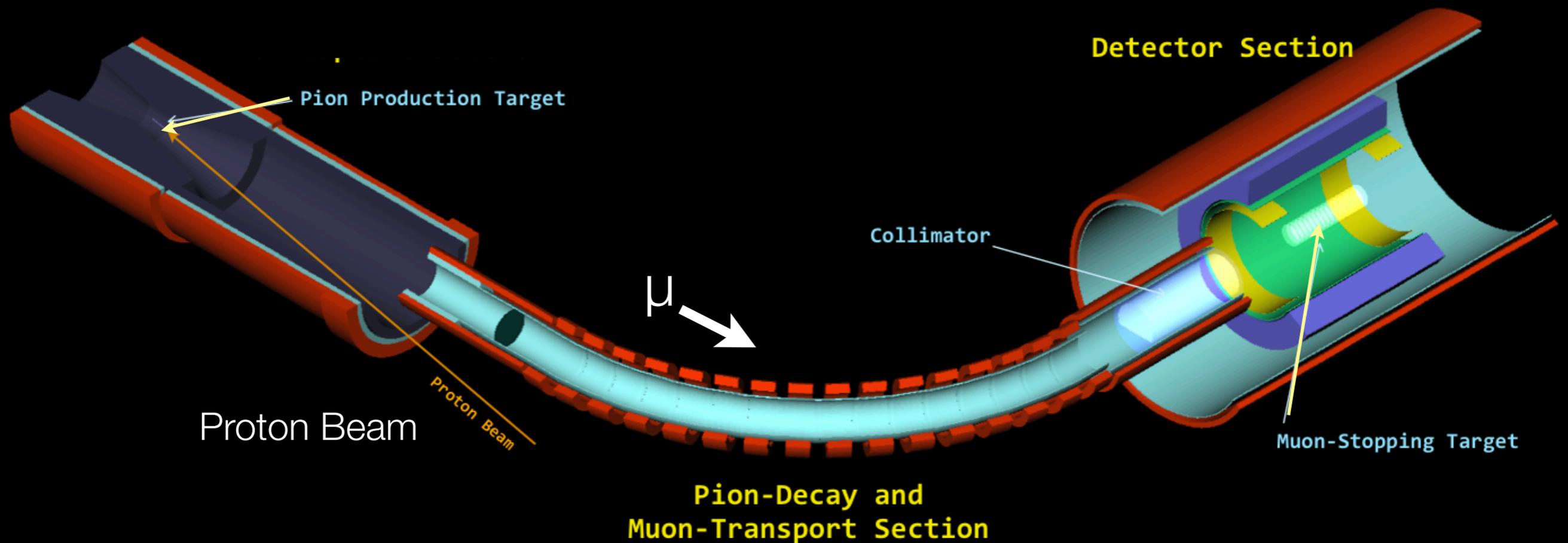
Search for μ -e conversion

a search for $\mu^- - e^-$ conversion at intermediate sensitivity which would be more than 10 times better than the SINDRUM-II limit

CyDet



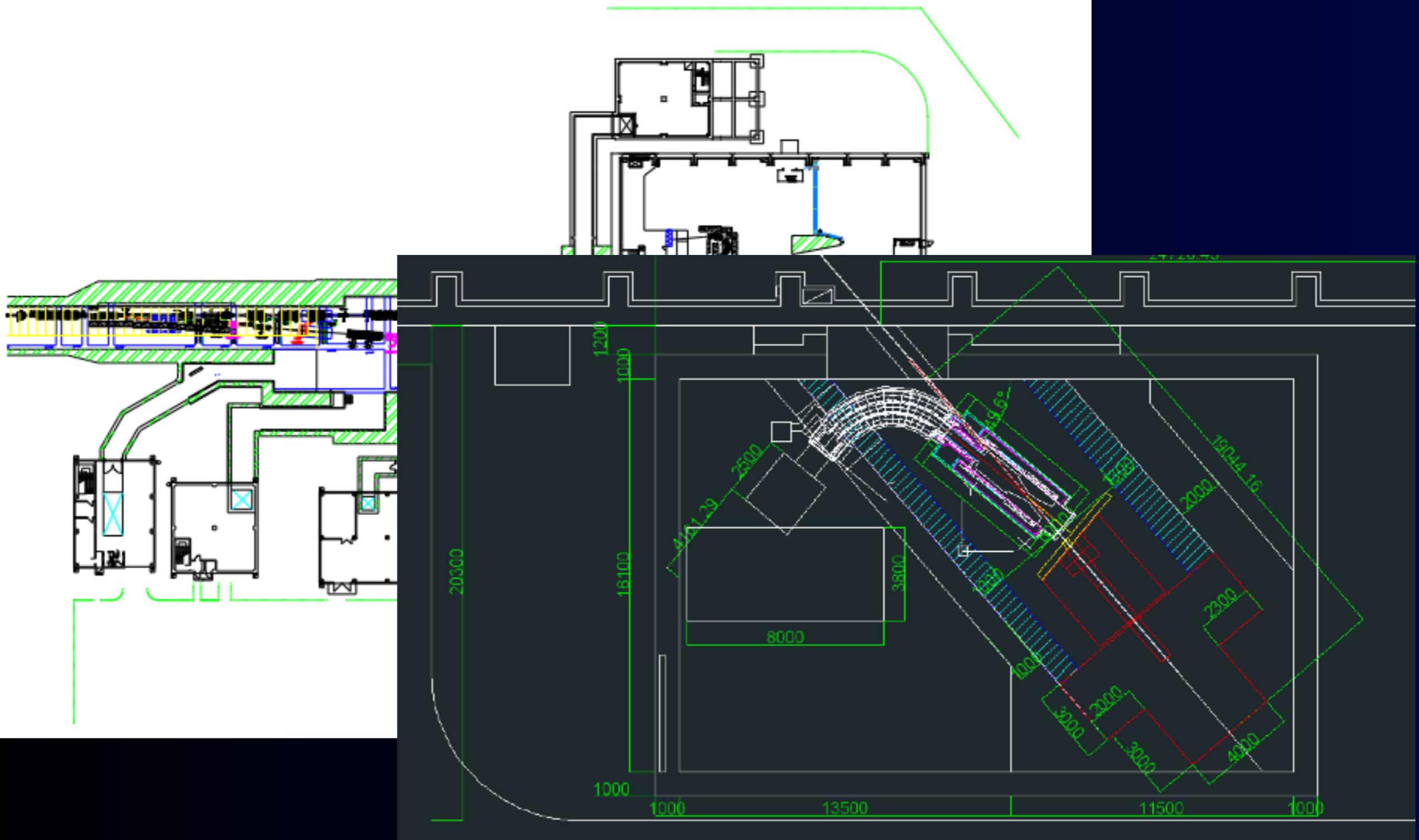
COMET Phase-I Experimental Layout



COMET muon beam-line :
 6×10^9 muon/sec with 3kW beam
produced. The world highest
intensity.

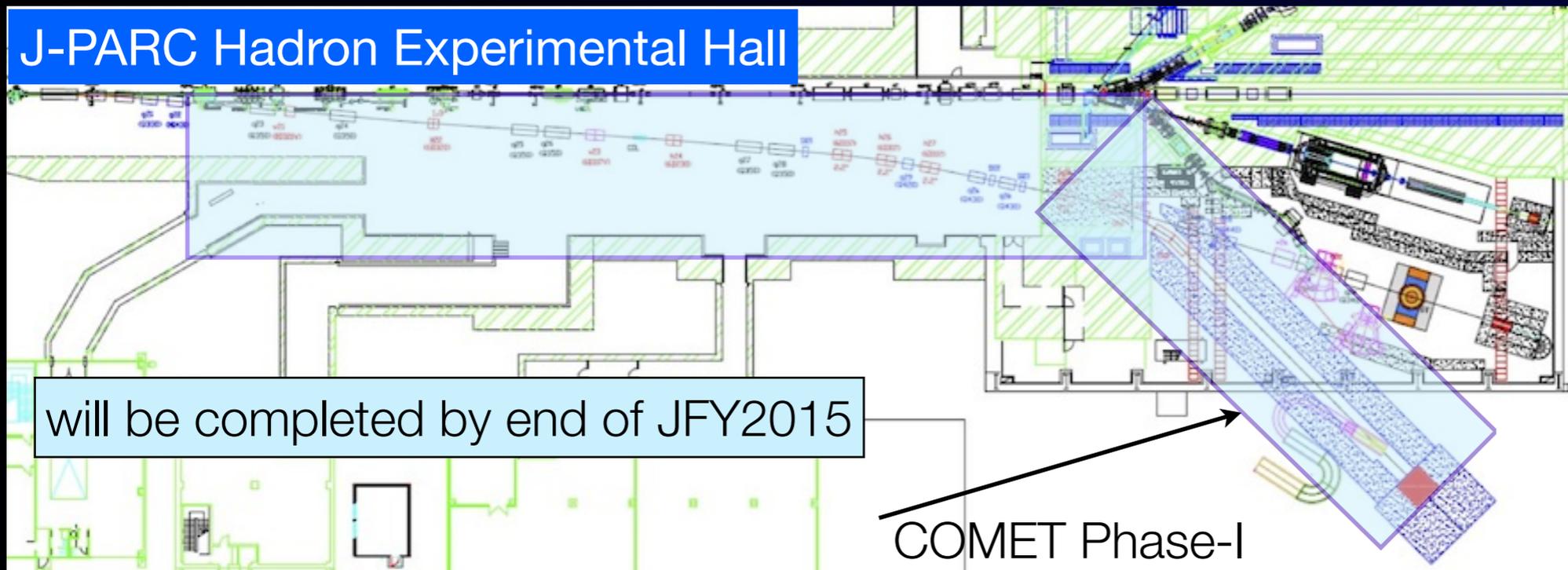
COMET Phase-I detector :
About 10^{16} muons are stopped in
the target. Electron from μ -e
conversion will be measured

COMET Beam line



Funds for Phase-I is secured.....

Budget for COMET Phase-I has been approved.



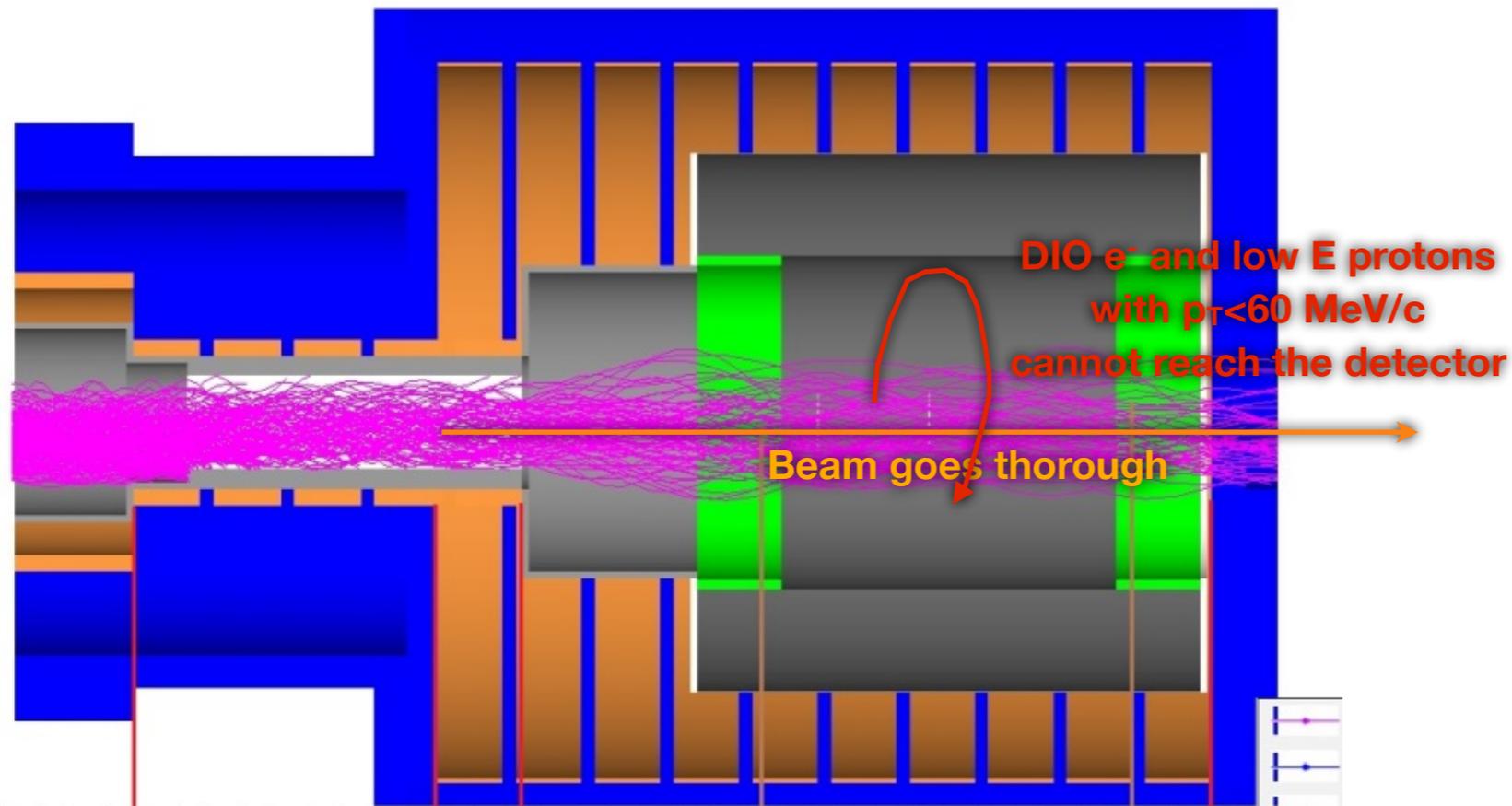
JFY2012
Supplemental
budget

High momentum
proton beam
line for nuclear
physics

Items		done by
proton beam-line	general use	KEK
muon beam-line		
COMET detector	exp. proper	COMET collaboration

Detector budget
(CDC and
detector
solenoid) has
been secured.

Cylindrical Drift Chamber Detector (CyDet)

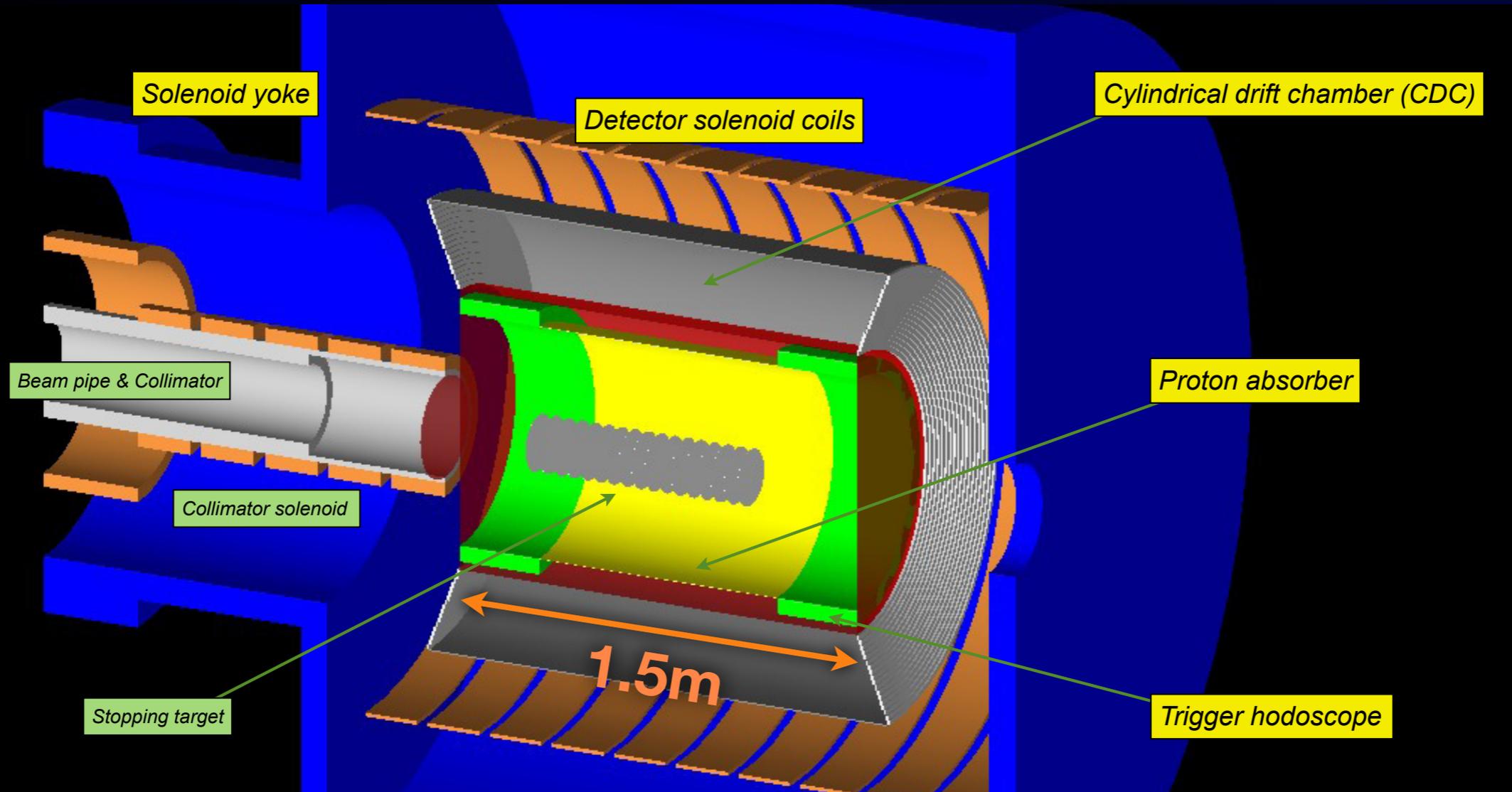


- to avoid hits of beam particles, DIO e^- , and low energy protons

Why CyDet ?

For Phase-I, no curved solenoid electron transport to momentum and charge select the particles is available. In the CyDet geometry, no beam particles hit the detector, and low momentum tracks do not reach the detector.

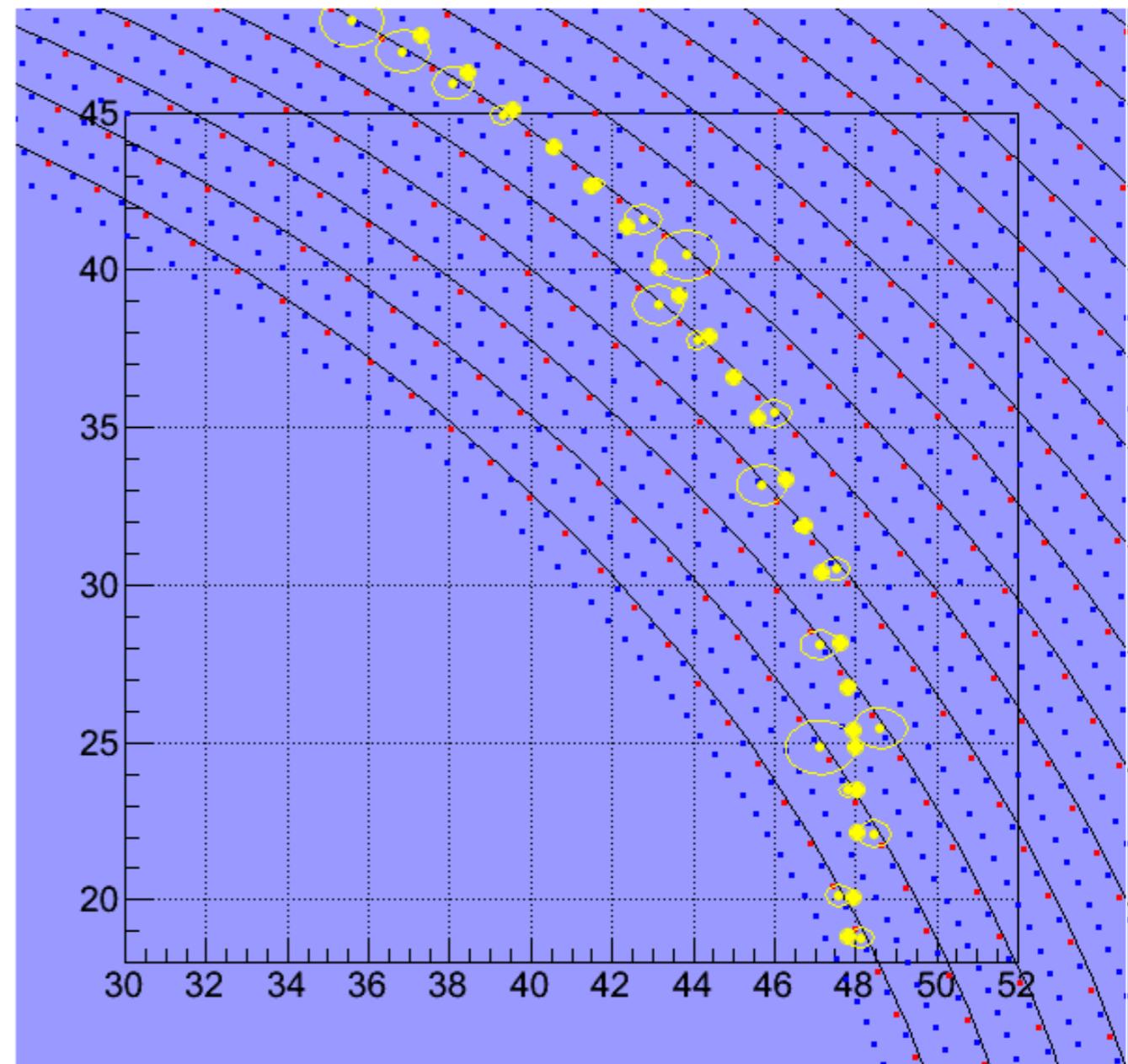
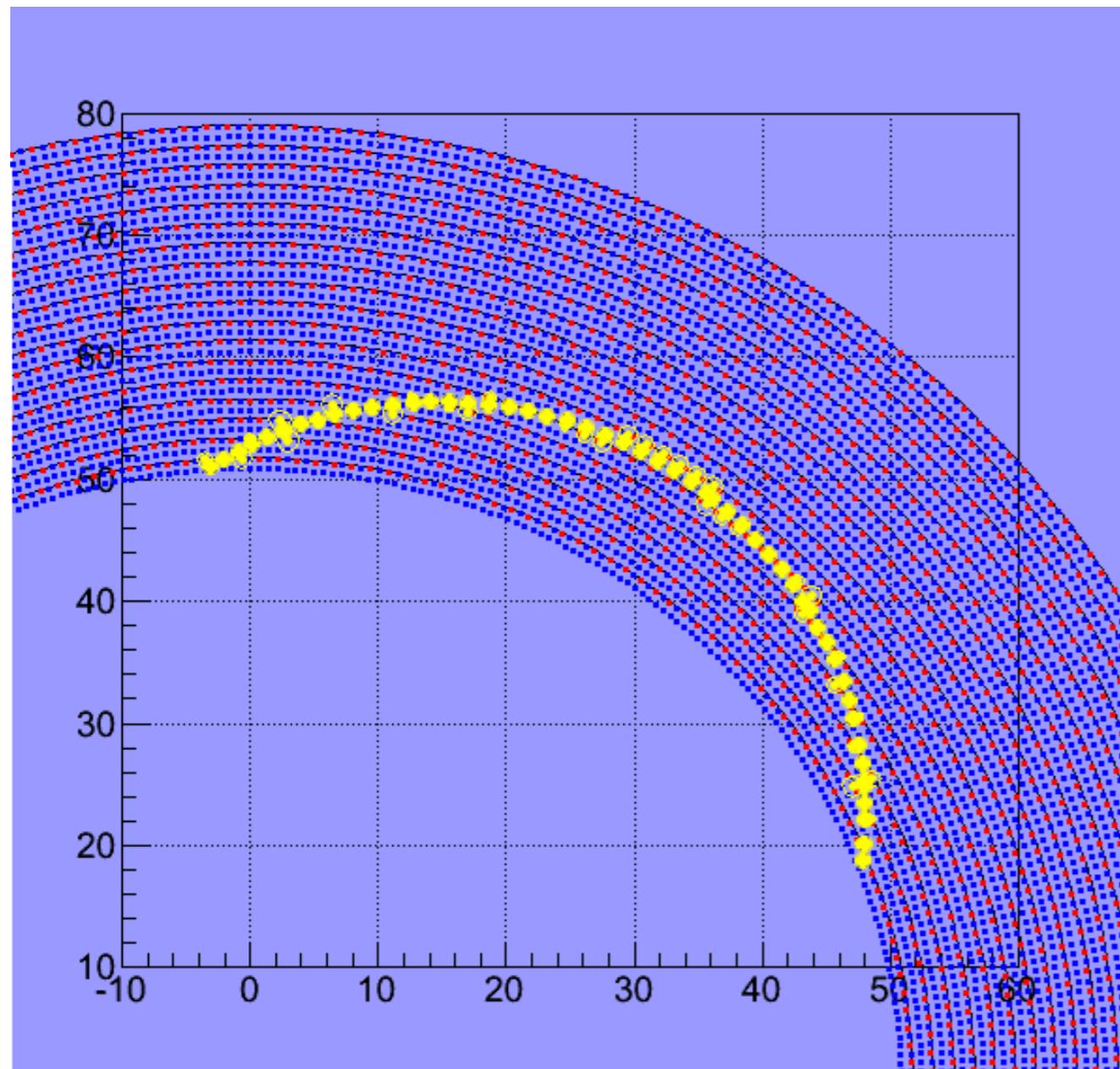
Cylindrical Drift Chamber Detector (CyDet)



CDC
design

- z hit position by stereo layers (all stereo layers)
- reduction of multiple scattering by helium based gas mixture
- large inner radius to reduce DIO electron hits. (rate)
- proton absorber of CFRP to reduce protons from muon capture

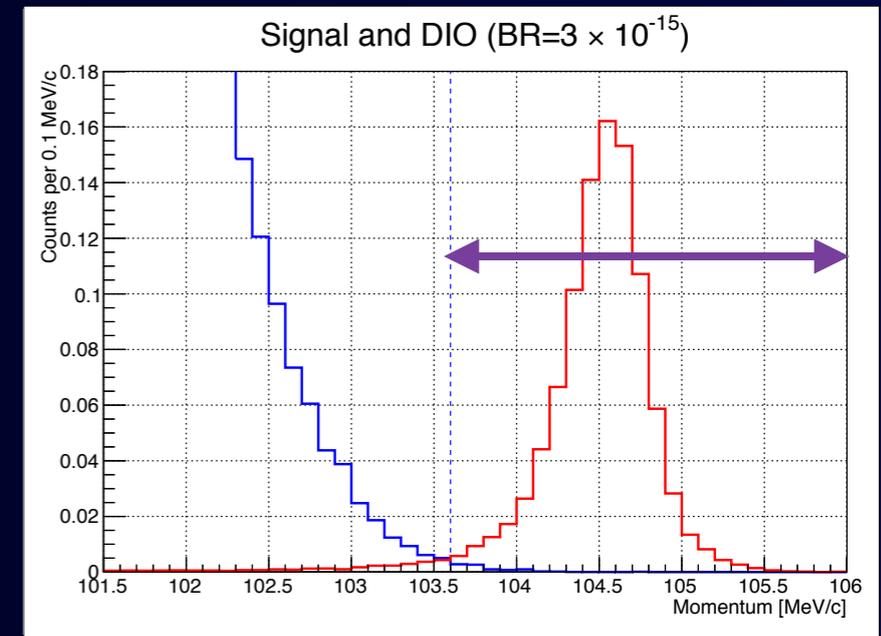
Cylindrical Drift Chamber Detector (CyDet)



Signal Sensitivity with CyDet

Signal Acceptance

Event selection	Value	Comments
Geometrical acceptance and tracking cuts	0.29	
Momentum selection	0.97	$103.6 \text{ MeV}/c < P_e < 106.0 \text{ MeV}/c$
Timing window	0.3	$700 \text{ ns} < t < 1100 \text{ ns}$
Trigger efficiency	0.8	
DAQ efficiency	0.8	
Track reconstruction efficiency	0.8	
Total	0.043	



Signal Sensitivity

$$B(\mu^- + Al \rightarrow e^- + Al) \sim \frac{1}{N_\mu \cdot f_{cap} \cdot A_e},$$

- $f_{cap} = 0.6$
- $A_e = 0.056$
- $N_\mu = 9.4 \times 10^{15}$ muons

$$B(\mu^- + Al \rightarrow e^- + Al) = 3.1 \times 10^{-17}$$

$$B(\mu^- + Al \rightarrow e^- + Al) < 7 \times 10^{-17} \quad (90\% C.L.)$$

Muon intensity

about 0.00064 muons stopped/proton

With 0.4 μA , a running time of about 90 days is needed.

Background List



Intrinsic physics backgrounds

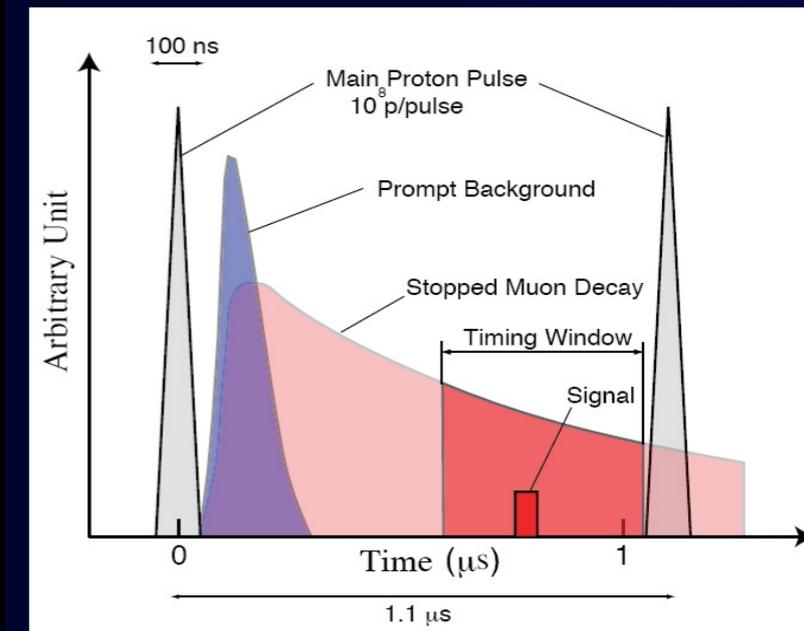
- | | | |
|---|---|--|
| 1 | Muon decay in orbit (DIO) | Bound muons decay in a muonic atom |
| 2 | Radiative muon capture (external) | $\mu^- + A \rightarrow \nu_\mu + A' + \gamma$,
followed by $\gamma \rightarrow e^- + e^+$ |
| 3 | Radiative muon capture (internal) | $\mu^- + A \rightarrow \nu_\mu + e^+ + e^- + A'$, |
| 4 | Neutron emission after
after muon capture | $\mu^- + A \rightarrow \nu_\mu + A' + n$,
and neutrons produce e^- |
| 5 | Charged particle emission
after muon capture | $\mu^- + A \rightarrow \nu_\mu + A' + p$ (or d or α),
followed by charged particles produce e^- |

Beam related prompt/delayed backgrounds

- | | | |
|----|-----------------------------------|--|
| 6 | Radiative pion capture (external) | $\pi^- + A \rightarrow \gamma + A'$, $\gamma \rightarrow e^- + e^+$ |
| 7 | Radiative pion capture (internal) | $\pi^- + A \rightarrow e^+ + e^- + A'$ |
| 8 | Beam electrons | e^- scattering off a muon stopping target |
| 9 | Muon decay in flight | μ^- decays in flight to produce e^- |
| 10 | Pion decay in flight | π^- decays in flight to produce e^- |
| 11 | Neutron induced backgrounds | neutrons hit material to produce e^- |
| 12 | \bar{p} induced backgrounds | \bar{p} hits material to produce e^- |

Other backgrounds

- | | |
|----|----------------------------------|
| 14 | Cosmic-ray induced backgrounds |
| 15 | Room neutron induced backgrounds |
| 16 | False tracking |



prompt and delayed
backgrounds

Background Estimate with CyDet

Table 26: Summary of the estimated background events for a single-event sensitivity of 3.1×10^{-15} with a proton extinction factor of 3×10^{-11} .

Type	Background	Estimated events
Physics	Muon decay in orbit	0.01
Physics	Radiative muon capture	5.6×10^{-4}
Physics	Neutron emission after muon capture	< 0.001
Physics	Charged particle emission after muon capture	< 0.001
Prompt Beam	Beam electrons (prompt)	7.1×10^{-4}
Prompt Beam	Muon decay in flight (prompt)	$\leq 1.7 \times 10^{-4}$
Prompt Beam	Pion decay in flight (prompt)	$\leq 2.0 \times 10^{-3}$
Prompt Beam	Other beam particles	$\leq 2.4 \times 10^{-6}$
Prompt Beam	Radiative pion capture(prompt)	4.24×10^{-4}
Delayed Beam	Beam electrons (delayed)	~ 0
Delayed Beam	Muon decay in flight (delayed)	~ 0
Delayed Beam	Pion decay in flight (delayed)	~ 0
Delayed Beam	Radiative pion capture (delayed)	~ 0
Delayed Beam	Anti-proton induced backgrounds	0.007
Others	Electrons from cosmic ray muons	< 0.0001
Total		0.019

Schedule (Facility)

2013

- Design of the building & beam line
- Bid tendering and start construction
- Design of superconducting solenoid magnets and start of construction
- Production of SC wires as well
- Design of the pion production target

2014

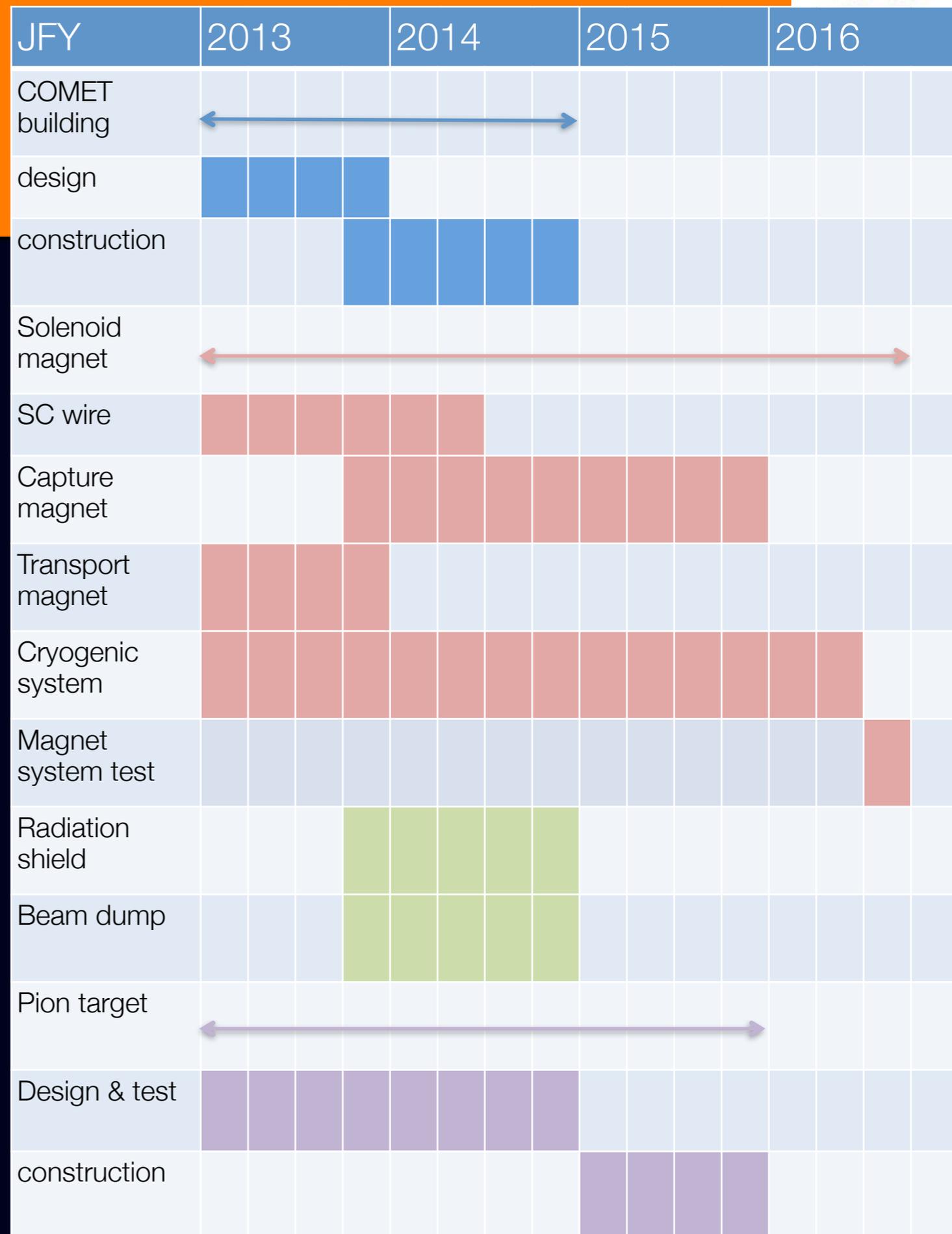
- Completion of the building
- Construction of superconducting solenoid magnets
- Start magnet and radiation shielding (and beam dump) installation
- Transport solenoid
- Start preparation of cryogenic system
- Tests of the target production target

2015

- Construction of superconducting solenoid magnets
- Preparation of cryogenic system
- Construction of the pion production target

2016

- Installation of the capture solenoid
- Completion of the cryogenic system
- Tests of the magnet system
- Installation of the target
- Ready to accept the 8GeV beam



Schedule (Detector)



Detector construction

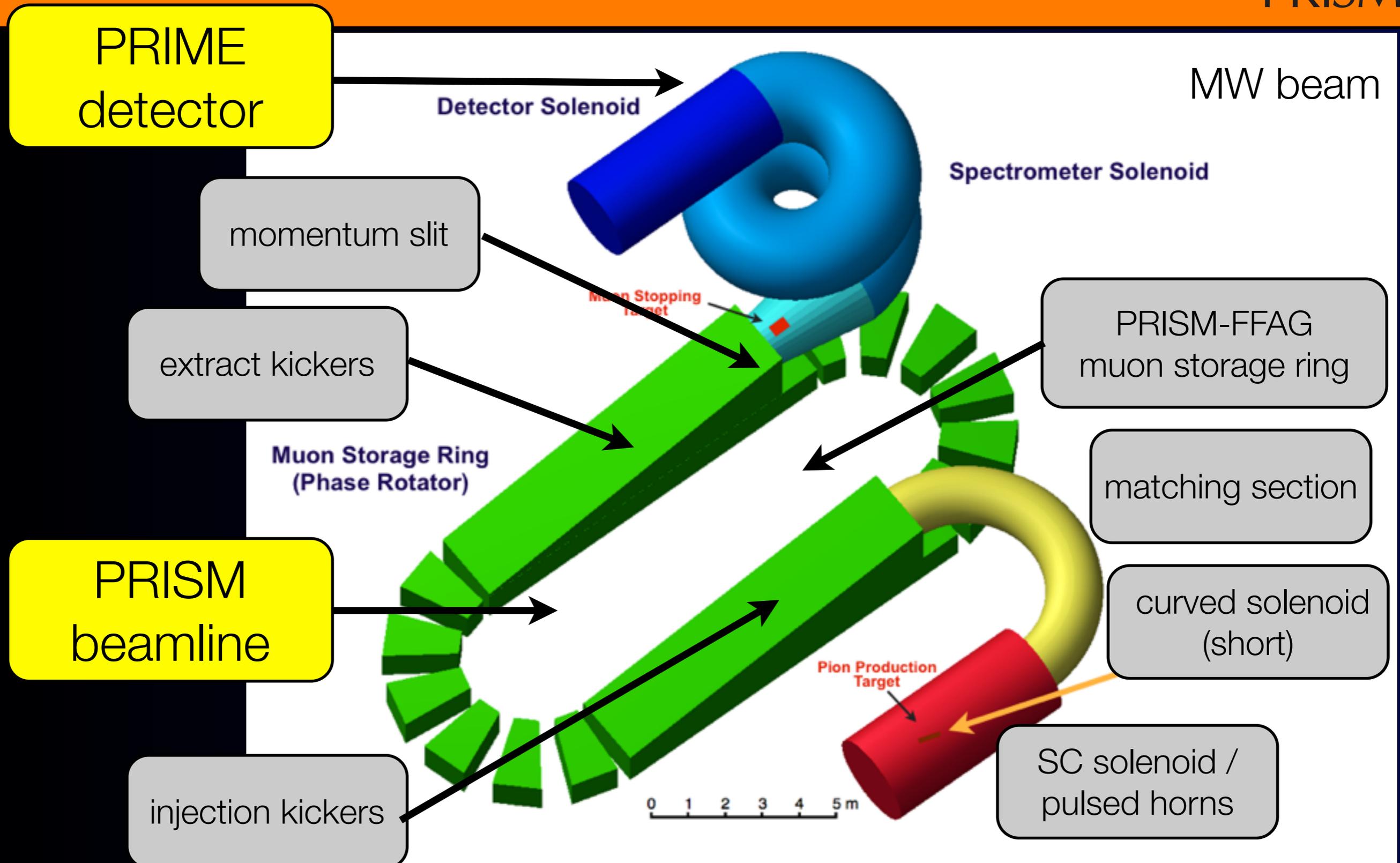
System	Works	2013	2014	2015	2016
CDC	Prototype construction	■			
	Endplate construction	■			
	Inner & outer wall construction	■			
	wire stringing & assembly		■		
	readout electronics	■			
	gas system construction		■		
	cosmic test			■	
CDC trigger counter	counter construction		■		
	readout electronics		■		
Detector Solenoid	design	■			
	superconductor fabrication	■			
	coil winding		■		
	cryostat construction			■	
	iron yoke production		■		
	power supply			■	
Collimator Solenoid	design	■			
	superconductor fabrication	■			
	coil winding		■		
	cryostat construction			■	
	iron yoke construction			■	
	power supply			■	
Muon Target	construction			■	

System	Works	2013	2014	2015	2016
Straw Chamber	straw construction		■		
	ROESTI prototype	■			
	ROESTI production		■		
	vacuum chamber			■	
	gas system construction			■	
	ECAL	crystal construction		■	
ROESTI prototype		■			
ROESTI production			■		
Cosmic Veto	counter construction		■		
	SiPM production		■		
	installation			■	
Muon Target	construction			■	
DAQ	trigger prototype		■		
	trigger production			■	
	DAQ development	■			
Muon collimator	beam measurement				■
	construction				■
X-ray detector	procurement	■			
Experiment	Transportation				■
	Installation				■
	Engineering run				■
	Physics run				■

Future Future Prospects
of μ -e conversion of 3×10^{-19}



μ -e conversion at S.E. sensitivity of 3×10^{-19} PRISM/PRIME (with muon storage ring)



R&D on the PRISM-FFAG Muon Storage Ring at Osaka University



PRISM-FFAG (6 sectors) in RCNP, Osaka



demonstration of phase rotation has been done.

Summary



- CLFV would give the best opportunity to search for BSM. (So far, no BSM signals at the LHC.)
- Muon to electron conversion could be one of the important CLFV processes in terms of theoretical and experimental points of view.
- COMET (Phase-II) at J-PARC is aiming at S.E. sensitivity of 3×10^{-17} .
- After the staged approach, COMET Phase-I is aiming at S.E. sensitivity of 3×10^{-15} . The beam line construction has been funded at KEK and the construction will start in 2013. Hope to do a measurement in 2016.

IKU (go ahead)

