

Double Chooz

on the θ_{13} quest

LBLN-Berkeley
September 2010

Anatael Cabrera
アナタエル カブレラ

CNRS / IN2P3
Double Chooz @ APC (Paris)

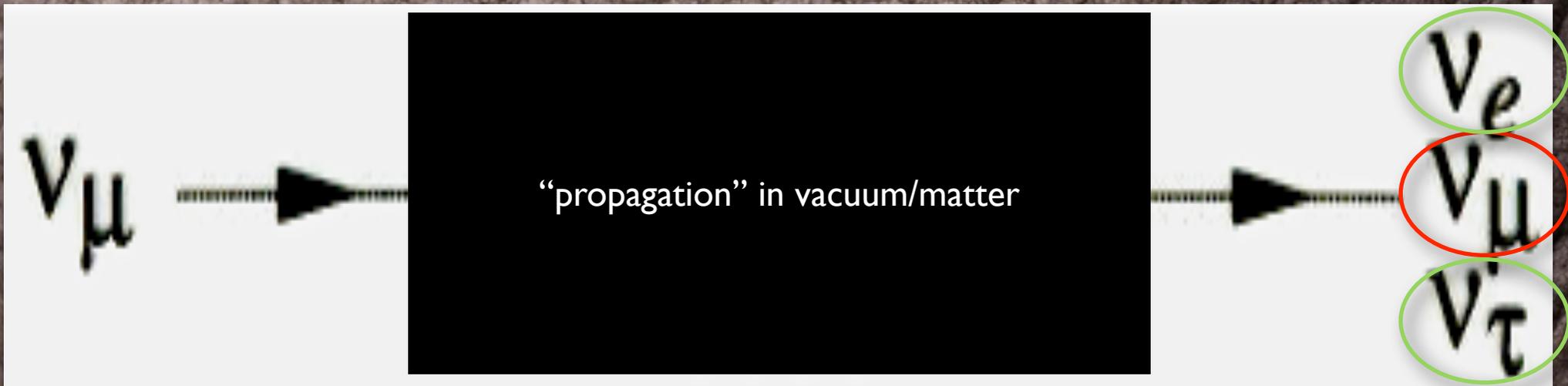
- **what?**
 - neutrino oscillation (a “micro-intro”)
 - experimental basics
 - status of leptonic mixing sector \Rightarrow the PMNS matrix
- **how?**
 - reactor neutrinos
- **where and when?** (not inclusive)
 - Double Chooz
- **concluding...**

neutrino oscillations...

Anatael Cabrera (CNRS-IN2P3 & APC)

Wednesday, 29 September 2010

Let's take ν_μ (a good example) to start with...



disappearance experiment goal
appearance experiment goal

how about “mixing”...?



$$V_{\alpha} = 0.5 \cdot V_1 + 0.5 \cdot V_2$$

assumption:
biology likes 50% mixing

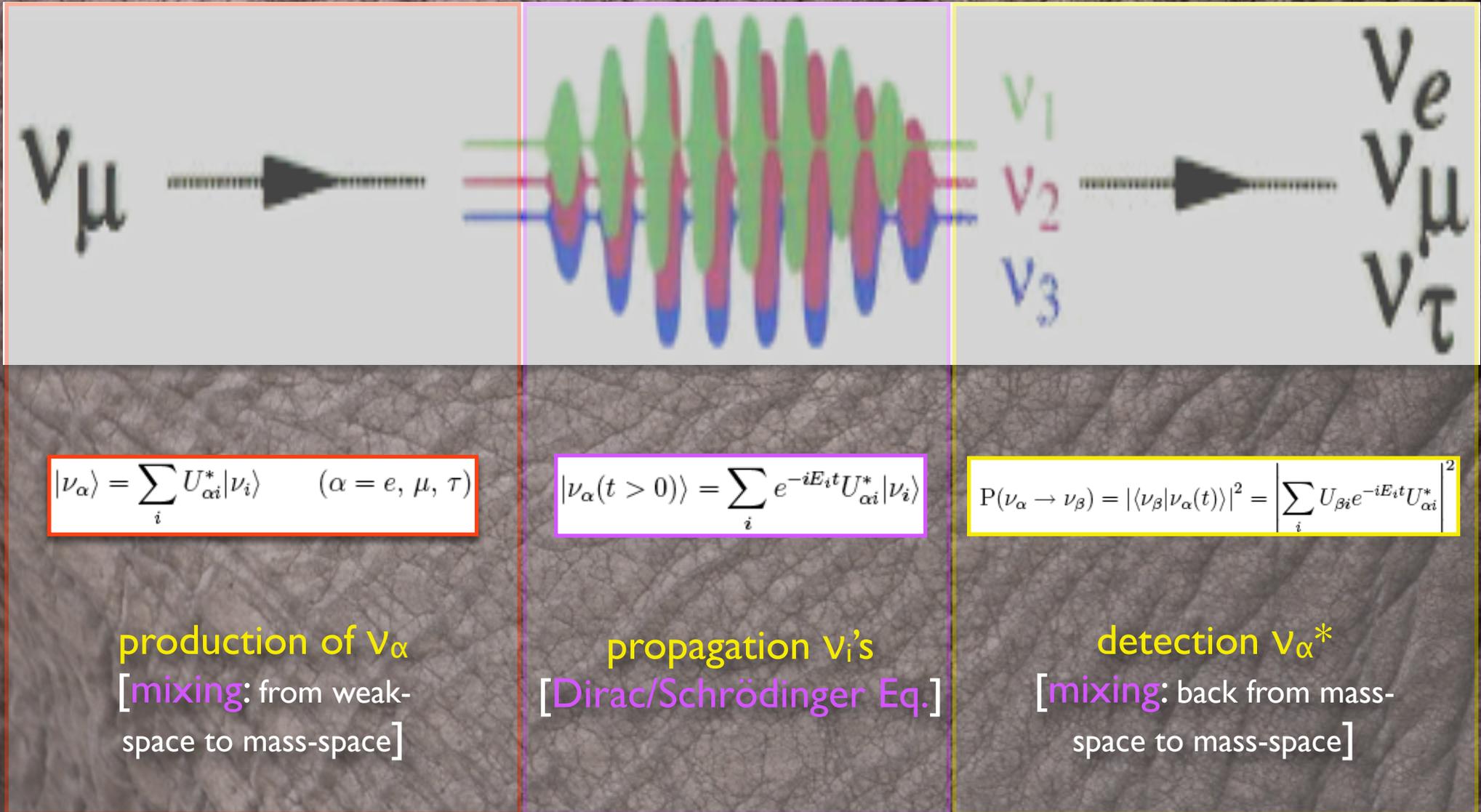
| Cabrera (CNRS-IN2P3 & APC)

$$(\nu_e, \nu_\mu, \nu_\tau)^T = \mathbf{U} (\nu_1, \nu_2, \nu_3)^T$$

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$|\nu_\alpha\rangle = \sum_i U_{\alpha i}^* |\nu_i\rangle \quad (\alpha = e, \mu, \tau)$$

if $\mathbf{U} = \mathbf{I} \Rightarrow$ trivial mixing (i.e. no mixing)



solution for 2x2 ν case...

$$P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2 2\theta \sin^2 \left(\frac{1.27 \Delta m^2 L}{E} \right)$$

Mixing in the leptonic sector (θ) \Rightarrow **PMNS** matrix (à la **CKM**)

Non-degenerate mass spectrum (Δm) \Rightarrow (macroscopic) *quantum interference*

L & **E** to be tuned (i.e. *experimental setup*) \Rightarrow measure **P(L₀, ΔE)**

- **flavour- ν s** (@ interaction) while **mass- ν s** (during propagation)
- experimental observation: **(dis)appearance** \Rightarrow free-Hamiltonian causing mutation
 - physics behind? a “mechanism” causing a non-diagonal free-Hamiltonian
- **high precision $O(<10\%)$** \Rightarrow experiments sensitive to 3x3 oscillation ν formalism
- oscillations explains most experimental evidence to date \Rightarrow only physics?
- **oscillation** means:
 - **mixing** in lepton sector: PMNS matrix (*à la CKM*, for quarks)
 - **prediction**: leptonic CP-violation (in-built on complexity of matrix)
 - **matrix must be unitary** \Rightarrow measure number ν s: 3?
 - **non-degenerate mass spectrum** of ν s
 - no lepton-flavour number conservation on SM, but total-lepton number
 - “mirroring” lepton-quark mixing \Rightarrow (appealing) \Rightarrow beyond SM?

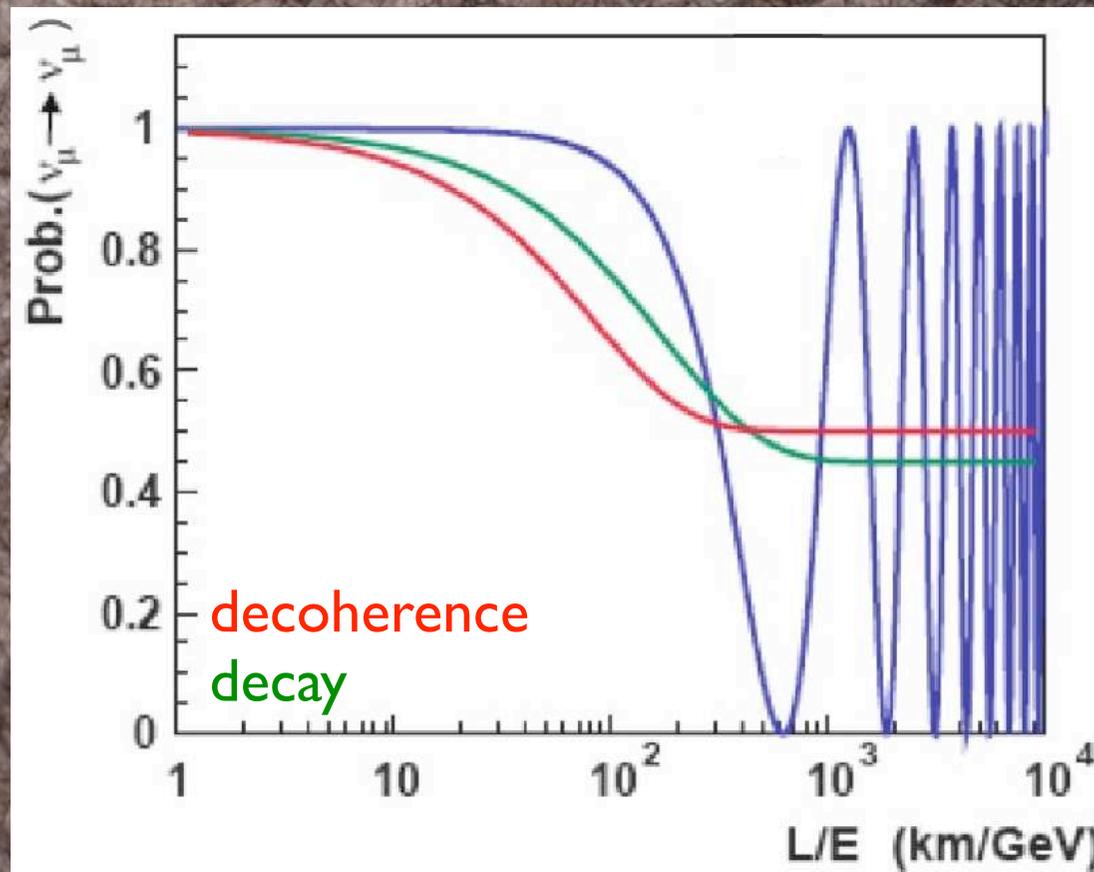


hunting ν -oscillations...

2x2 oscillations probability...

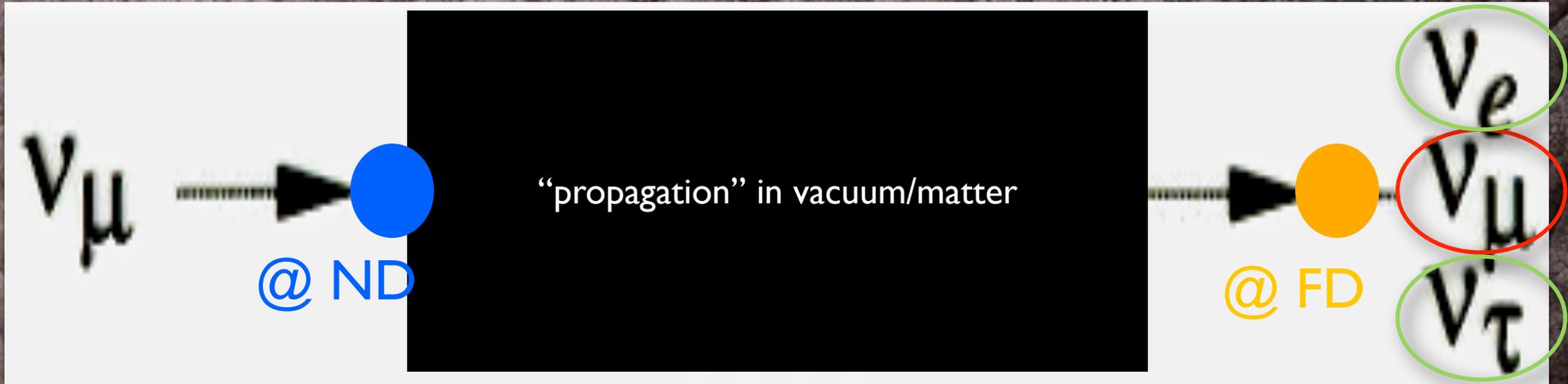
$$P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2 2\theta \sin^2 \left(\frac{1.27 \Delta m^2 L}{E} \right)$$

Disappearance: E/L modulation unique feature!



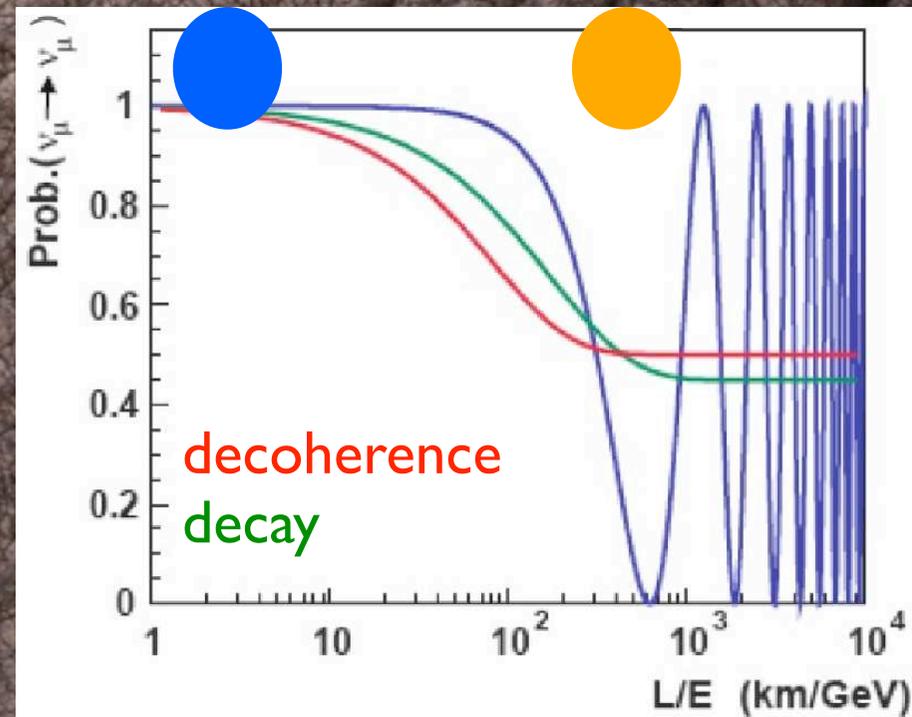
neutrino oscillations: a cartoon

disappearance experiment goal
appearance experiment goal



$L \Rightarrow$ one fixed value (if one source)

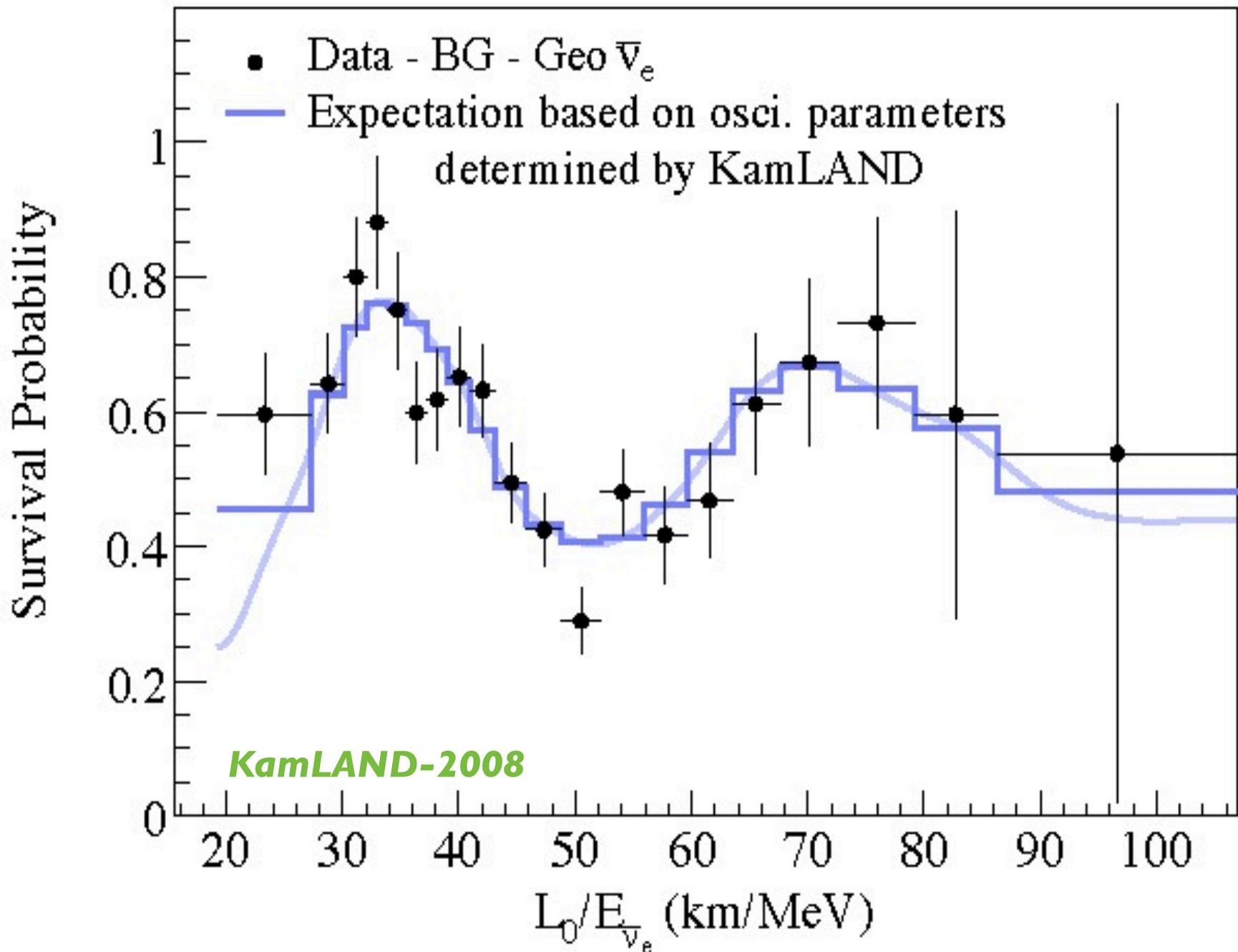
$E \Rightarrow$ several decades in energy (spectrum)

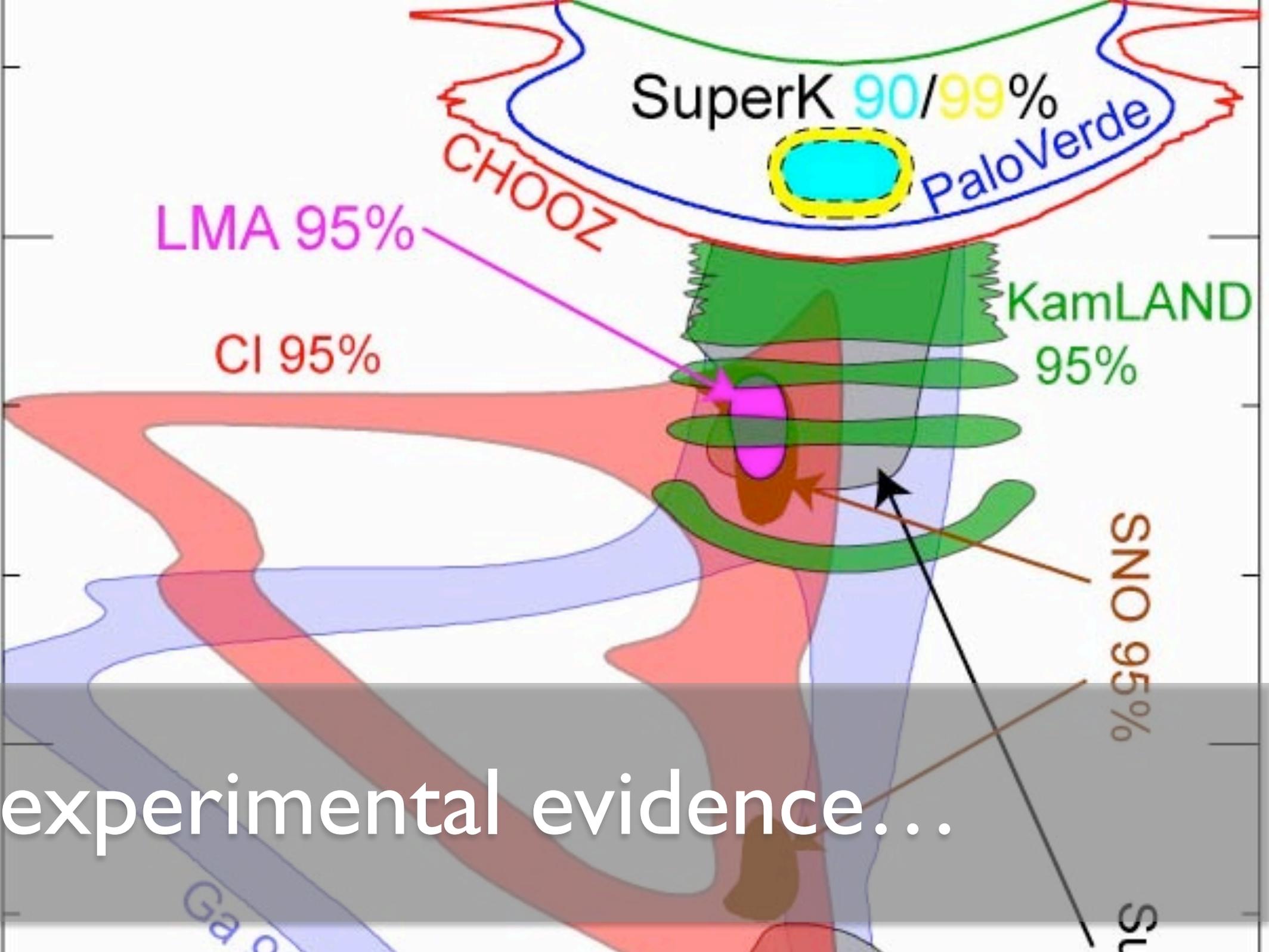


- **Experimentalists control L and E** \implies tune E/L ratio to maximise oscillation effect
- **Length (L):**
 - loss in flux ($1/L^2$)
 - longer allows to explore “Matter Effects” (or the MSW effect) or “Non-Standard-Interactions” (NSI) \implies not always desired (more observables)
- **Energy (E):**
 - determined by neutrino source (reactors, beams, sun, SNs, cosmic rays, etc)
 - **cross-section:** $\sigma(E) \propto E$ (linearly) [\implies statistics too]
 - **cross-section regime:** $\sigma(E)$ is very rich @ $O(1\text{ GeV})$
 - **lepton production threshold** @ low energy CC interactions
 - **detection efficiency:** typically tailored to **one lepton** or even **interaction type**
 - example: interactions type \implies high particle multiplicities

“solar” E/L @ KamLAND (reactor- ν_s)

The most beautiful E/L plot so far...





Solar & atmospheric “anomalies” appeared as *rather decoupled* problems
 \Rightarrow due to very different Δm^2 & the smallness of θ_{13} (sub-leading)

$$(\nu_e, \nu_\mu, \nu_\tau)^T = \mathbf{U} (\nu_1, \nu_2, \nu_3)^T$$

$$\begin{array}{c}
 \text{“atmospheric”} \Rightarrow \theta_{23} \\
 \text{“solar”} \Rightarrow \theta_{12}
 \end{array}
 \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ \text{sub-leading} & 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & 0 & c_{13} \\ \text{sub-leading} & & & \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$P(\nu_\mu \rightarrow \nu_\mu)$

$P(\nu_e \rightarrow \nu_e) \text{ \& } P(\nu_\mu \rightarrow \nu_e)$

$P(\nu_e \rightarrow \nu_x)$

Anatael Cabrera
 アナタエル カブレラ

PMNS (Unitary & 3x3) \Rightarrow 3 mixing angles & 1 complex phase \Rightarrow **leptonic CP violation**
 Double Chooz @ APC (Paris)

★ $\sin^2(2\theta_{13}) < 0.12-0.20$

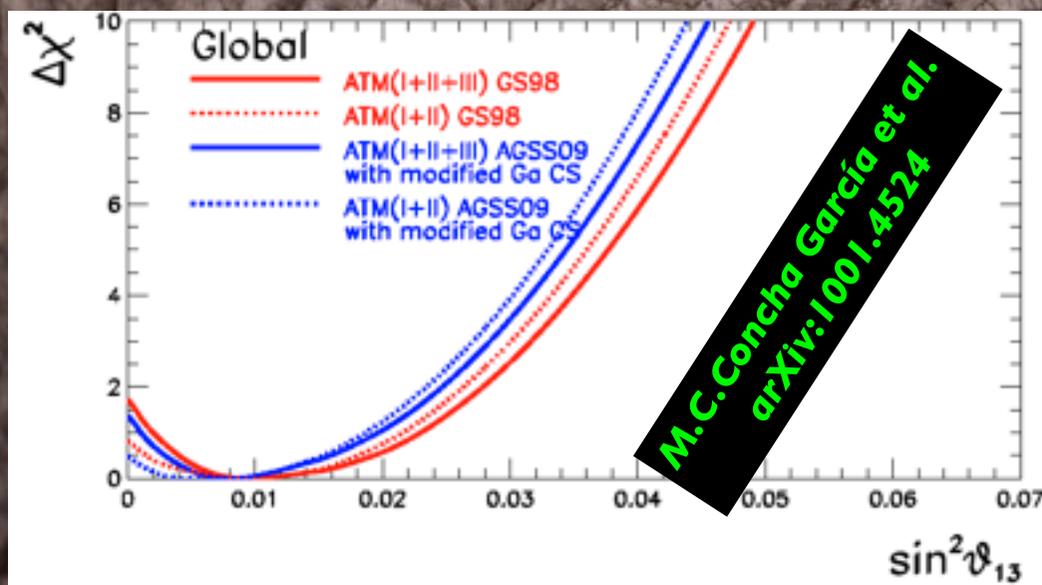
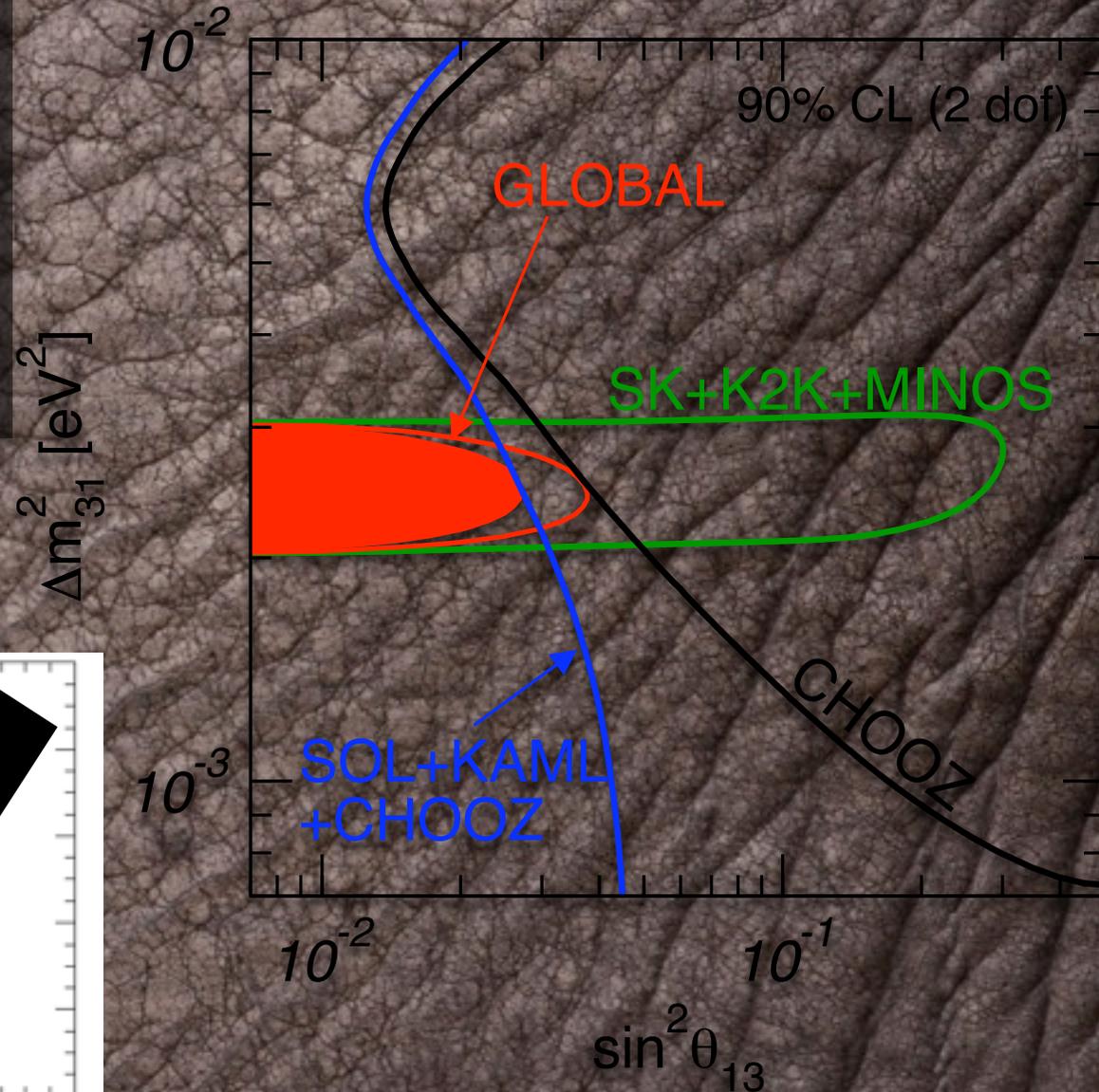
CHOOZ only @ 90%CL

★ $\sin^2(2\theta_{13}) < \sim 0.10$

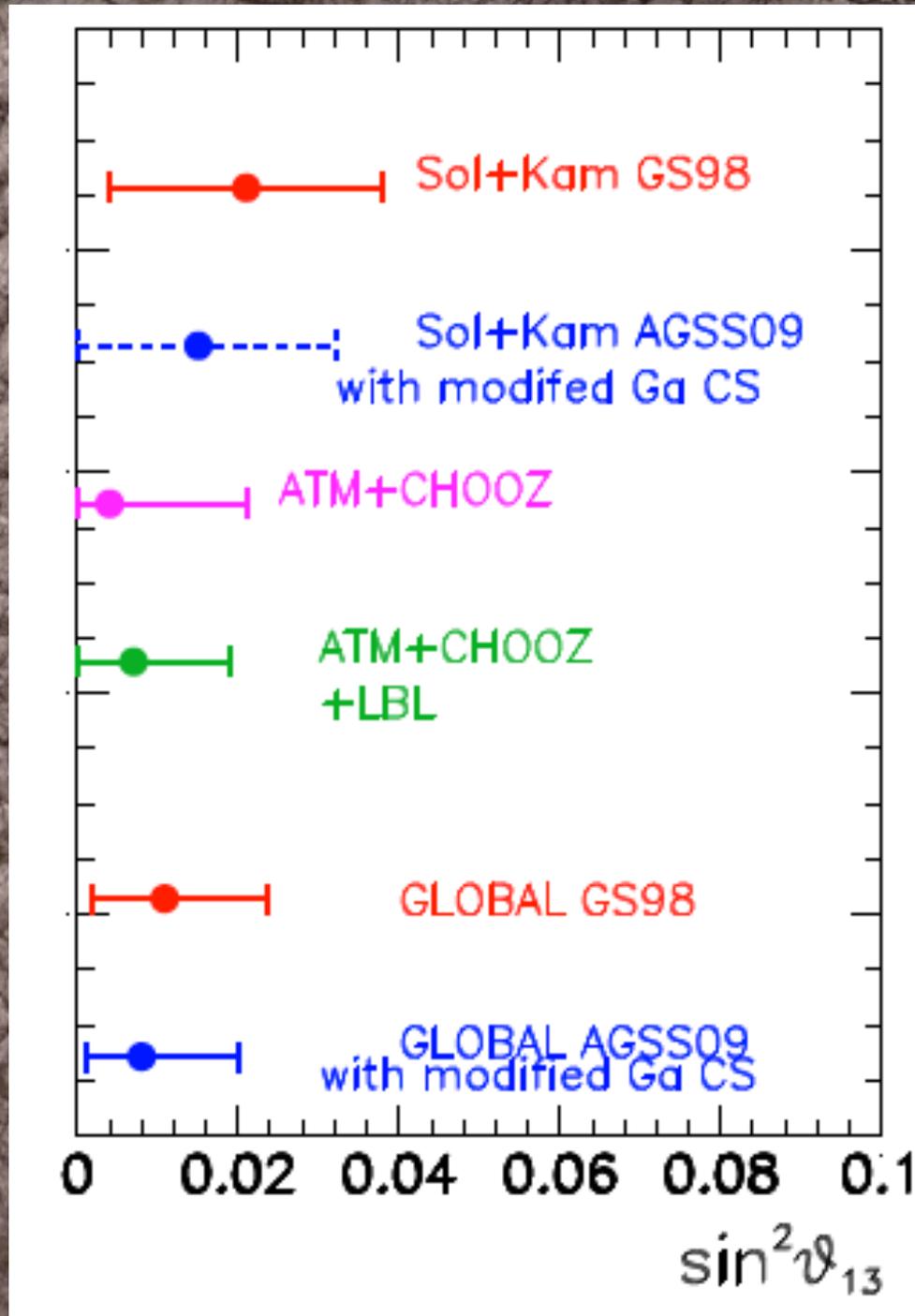
Global Analysis @ 90%CL

★ Global fit claims [Bari et al] \Rightarrow
“hint” for a non-zero θ_{13} @ $\sim 1.5\sigma$

M. Concha-García et al. arXiv:1001.4524
T. Schwetz et al. hep-ph/0606060



M.C. Concha García et al.
arXiv:1001.4524

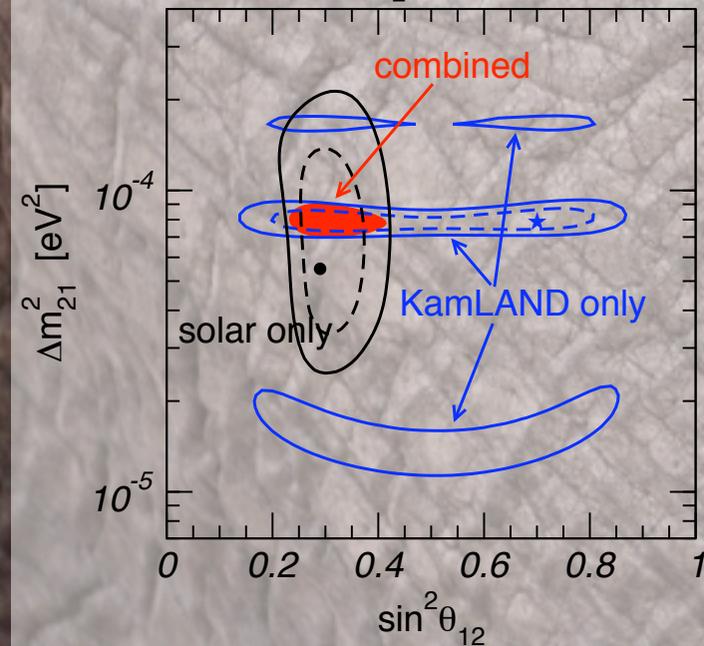


PMNS: large mixing (unlike CKM)...

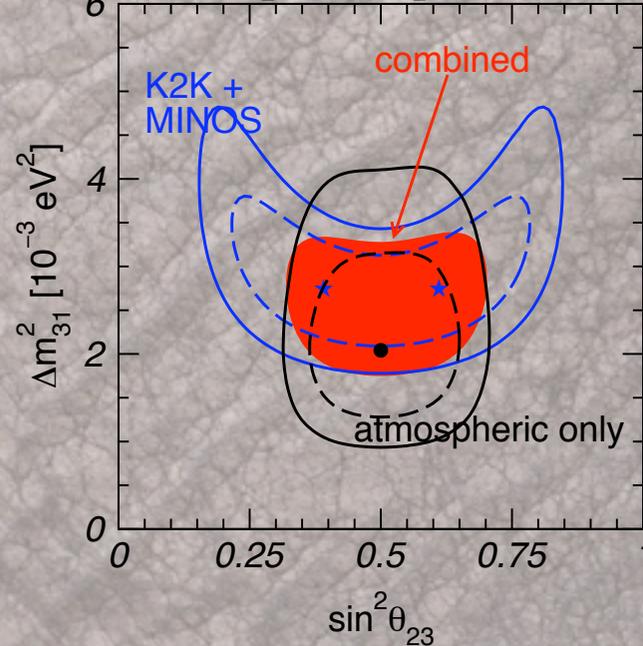
T. Schwetz et al. [hep-ph/0606060](http://arxiv.org/abs/hep-ph/0606060)

parameter	$bf \pm 1\sigma$	1σ acc.	2σ range	3σ range
Δm_{21}^2 [10^{-5}eV^2]	7.9 ± 0.3	4%	7.3 – 8.5	7.1 – 8.9
$ \Delta m_{31}^2 $ [10^{-3}eV^2]	$2.5^{+0.20}_{-0.25}$	10%	2.1 – 3.0	1.9 – 3.2
$\sin^2 \theta_{12}$	$0.30^{+0.02}_{-0.03}$	9%	0.26 – 0.36	0.24 – 0.40
$\sin^2 \theta_{23}$	$0.50^{+0.08}_{-0.07}$	16%	0.38 – 0.64	0.34 – 0.68
$\sin^2 \theta_{13}$	—	—	≤ 0.025	≤ 0.041

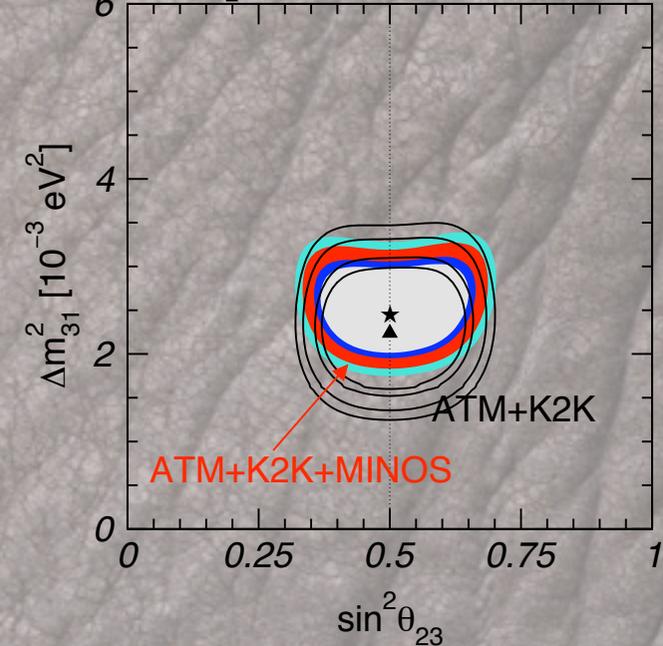
"solar" parameters



"atmospheric" parameters



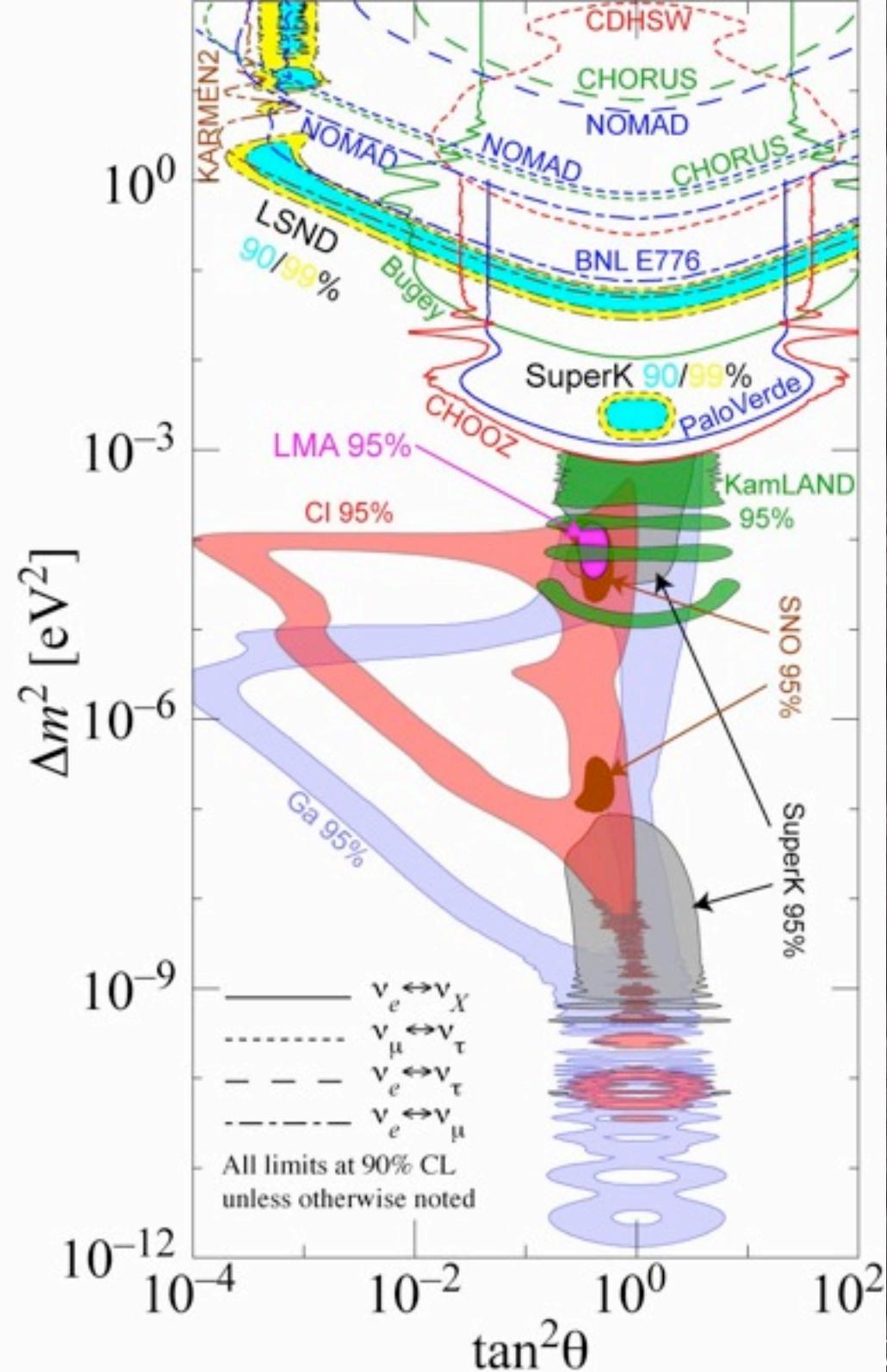
impact of MINOS data

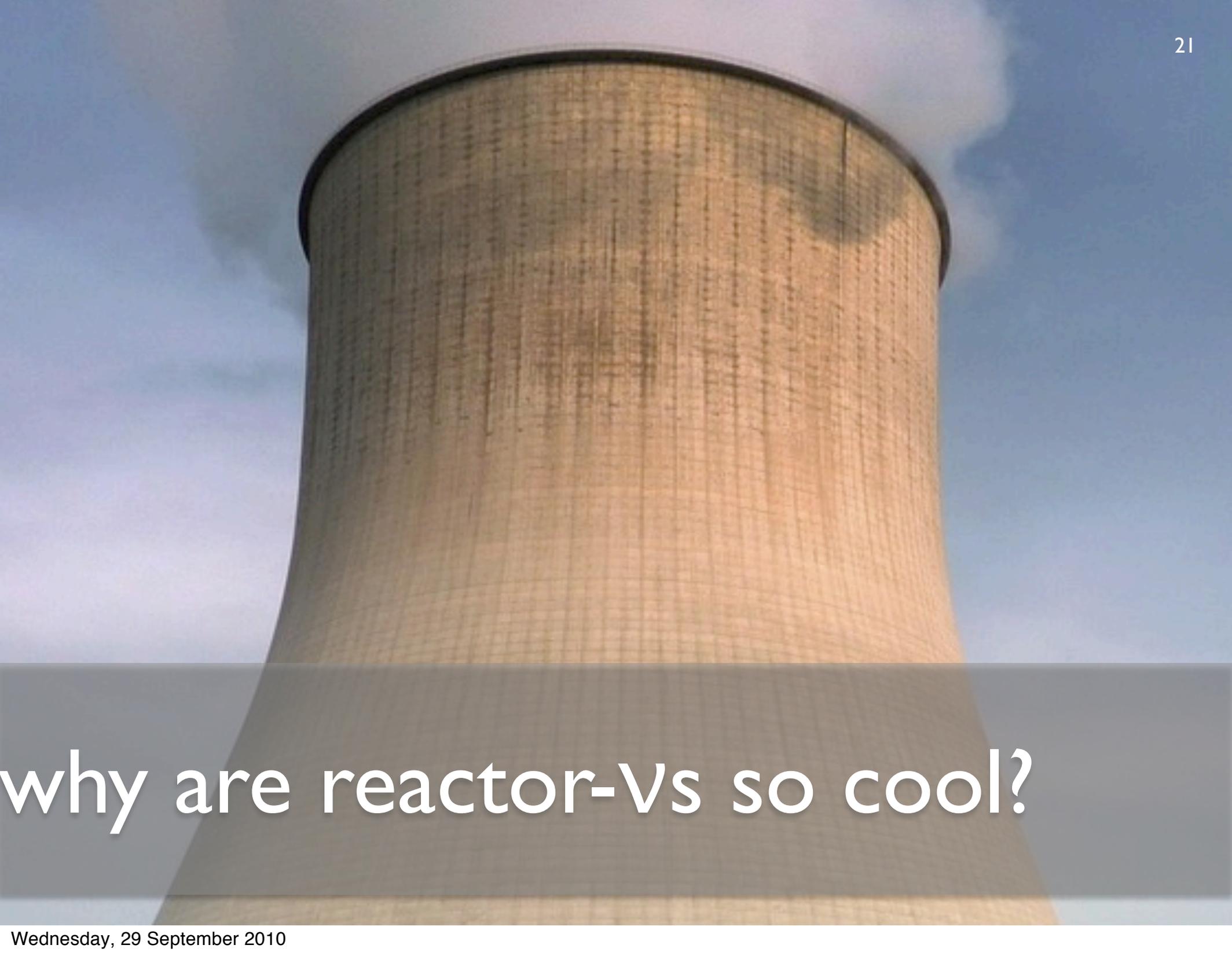


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visual summary of
experimental neutrino
oscillations phase-space
campaign...

$\theta_{12}, \theta_{23}, \Delta m^2_{12}, \Delta m^2_{23}$

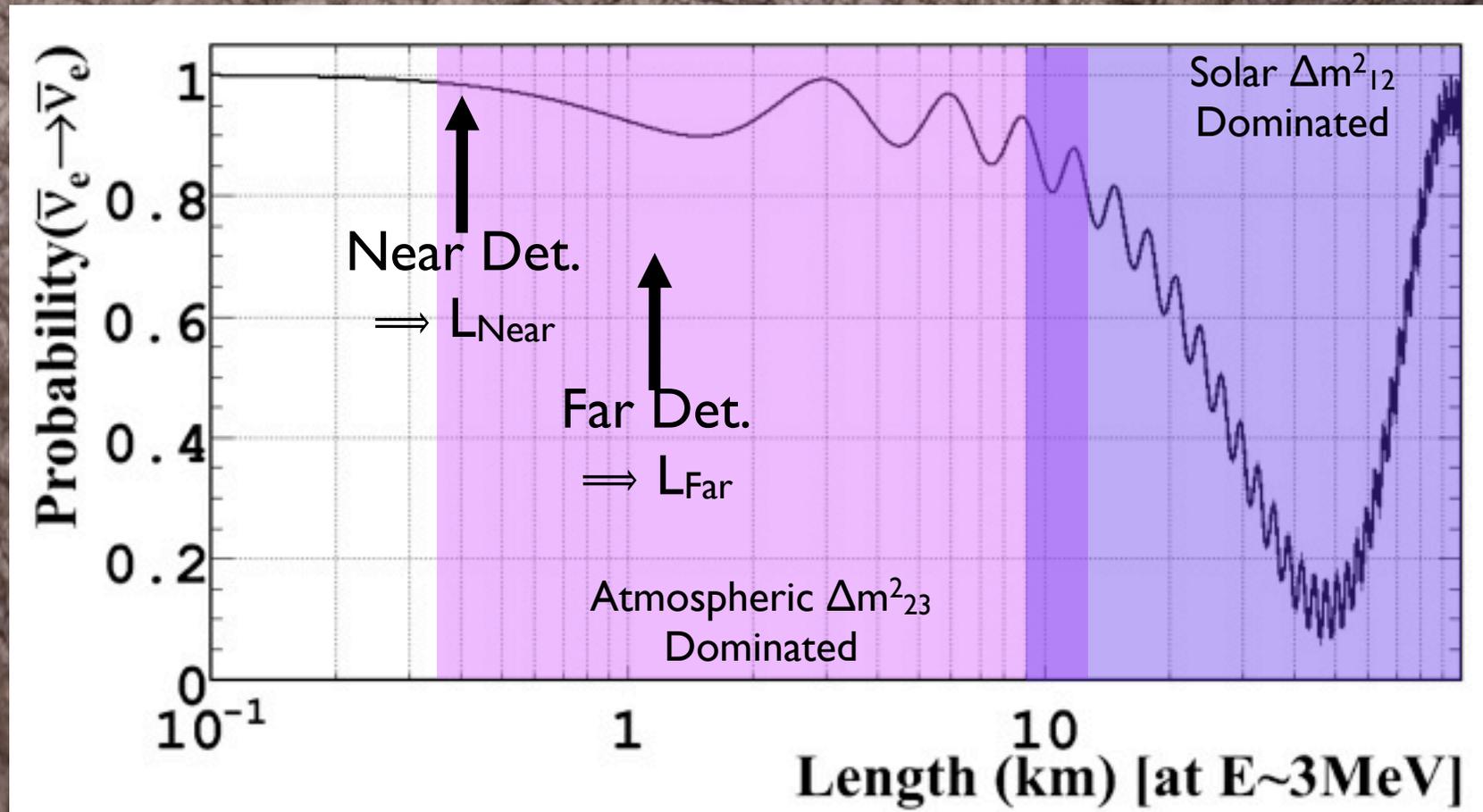




why are reactor-Vs so cool?

$$P(\nu_e \rightarrow \nu_e) \sim 1 - \sin^2(2\theta_{13}) \sin^2(\Delta m_{23}^2 L_0/E)$$

[plot: $E = 3\text{MeV}$, $\sin^2(2\theta_{13}) = 0.1$, $\Delta m_{23}^2 = 2.5 \times 10^{-3} \text{eV}^2$]



ND \Rightarrow **reduce systematic uncertainties** (mainly flux rate & shape) wrt FD

DC strongly involved in leading efforts to **improve reactor flux measurements**

- **copious** (high statistics) source and **free** vs
 - highly reliable beam \Rightarrow beam OFF cost $O(1M\text{€}/\text{day})$ (**strike-proof**)
- excellent **$\delta(E/L)$ resolution** \Rightarrow perfect for disappearance experiment
- [for θ_{13} searches] **short baselines** \Rightarrow
 - small detectors (less expensive)
 - negligible matter & “NSI” effects (useful for global analysis input)
- high & **well known cross-section** (no NC contamination)
- **BG** \Rightarrow overburden, shielding, radio-purity (possible “reactor OFF”)
- **trivial multi-detector extrapolation**: flux scales with $1/L^2$ (isotropic)
- **rich calibration energy scale**: many radioactive sources @ few MeV
- **1 unknown & 1 observable** \Rightarrow unambiguous θ_{13} signature
 - compelling *synergy* wrt to beam results (several unknown's) & global picture!



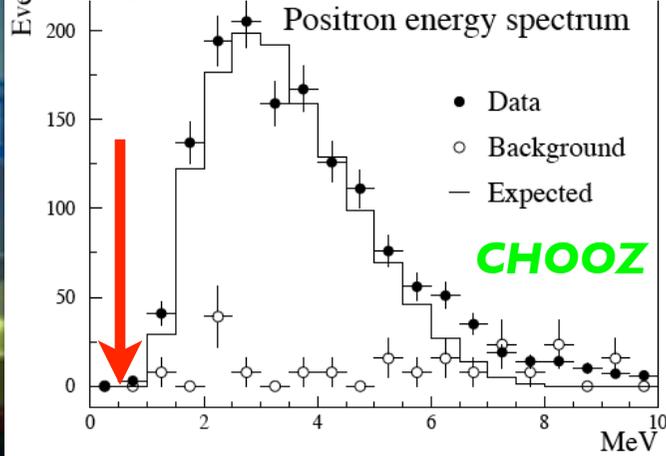
ON

OFF

Rue de Hayaumat

inverse- β reaction

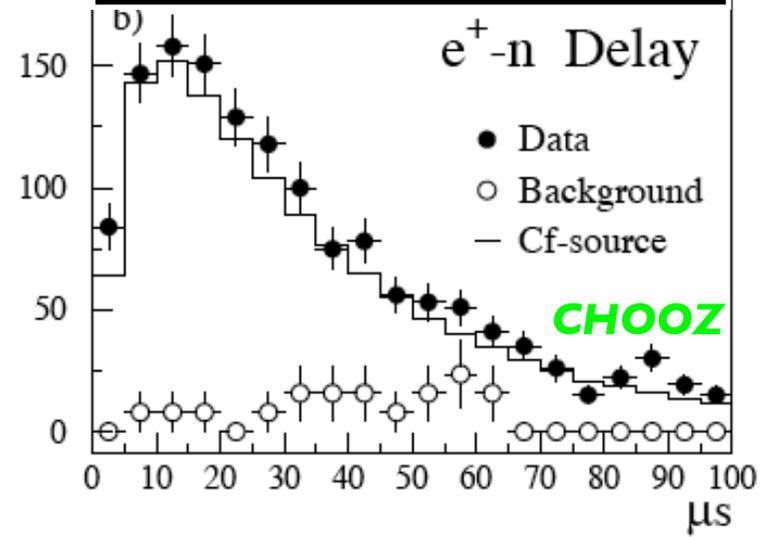
β^+ Energy \Rightarrow ν Energy
(spectral distortion possible)



anti- ν

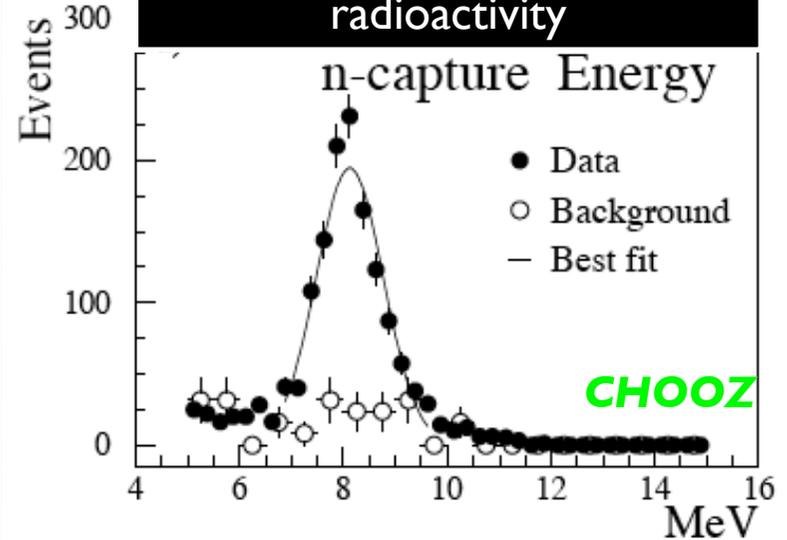
$$\sigma_{\text{threshold}}(E_\nu) = 1.8 \text{ MeV}$$

n is captured ΔT after β^+

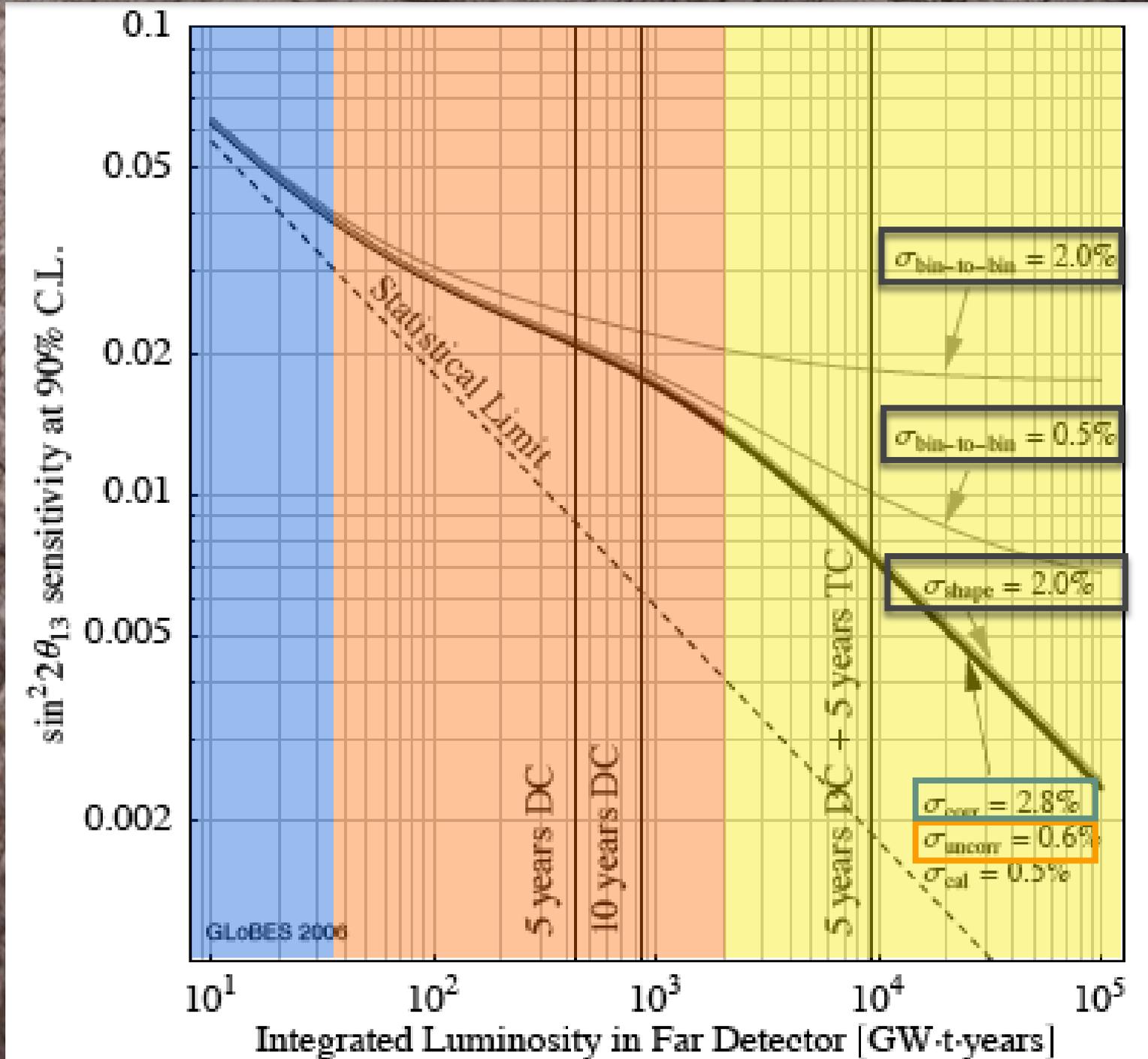


Coincidence ΔT depends on Gd concentration
 \Rightarrow excellent BG rejection mechanism

n-Gd capture signal: well above radioactivity



Apollonio et al (CHOOZ) hep-ex/0301017



Double Chooz Collaboration



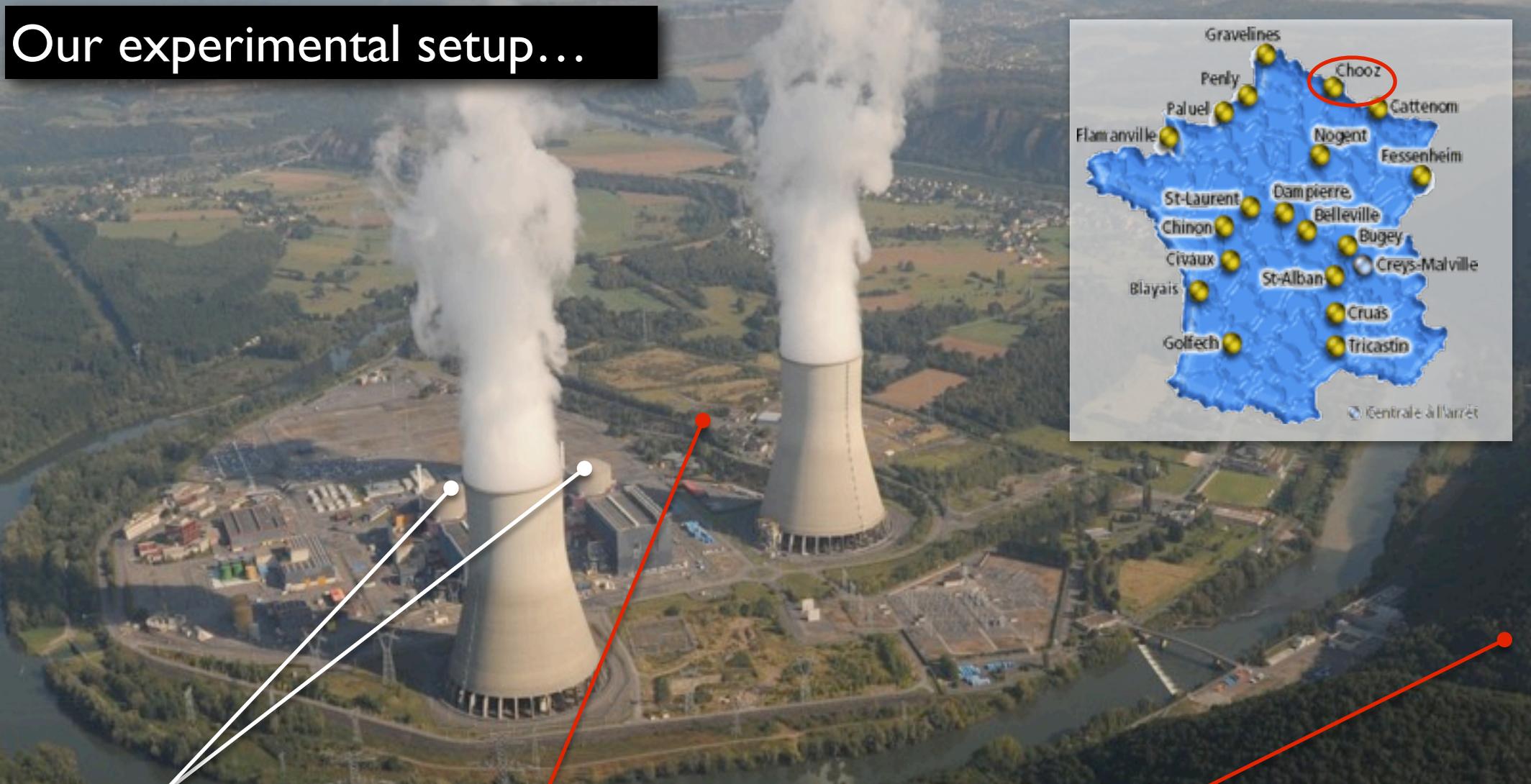
~150 physicists in 35 institutions over 9 countries in 3 continents

Hervé de Kerret (IN2P3-APC)
Spokesman



Wednesday, 29 September 2010

Our experimental setup...



Chooz Reactors
Power: 8.5GW_{th}
(N4s: most powerful)



Near
<L> 400m
400v/day
120mwe
Target: 8.2t
Early 2012



Far
<L> 1050m
50v/day
300mwe
Target: 8.2t
end of 2010

our θ_{13} knowledge versus time...

DC Proposal: hep-ex/0606025

Phases:

DC-I (FD only)

10x more statistics than CHOOZ

Limited by:

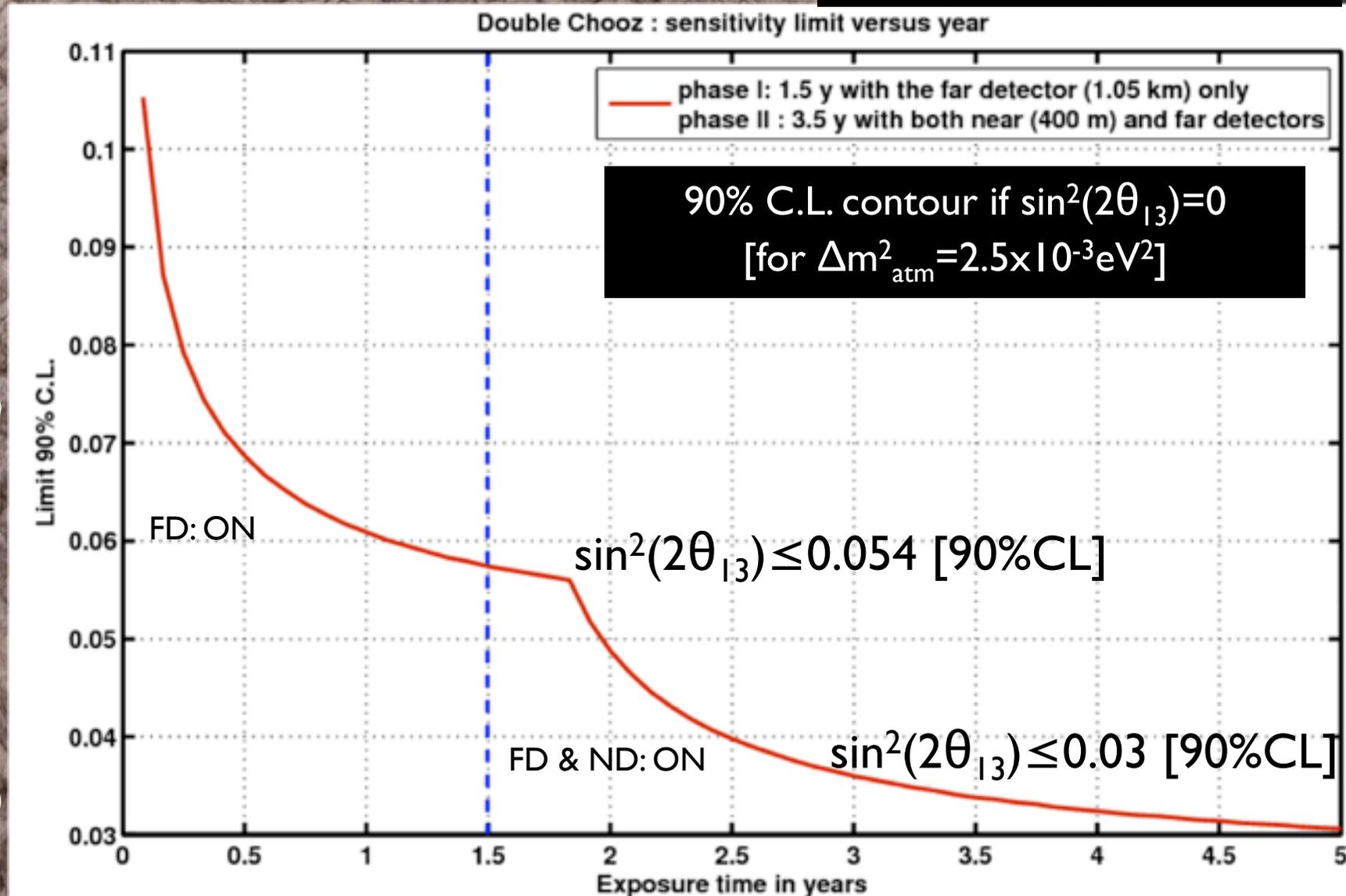
rate and shape reactor flux uncertainties (2.8%)

DC-II (FD+ND)

more robust

Limited by:

inter-detector normalisation systematic uncertainties (0.6%)

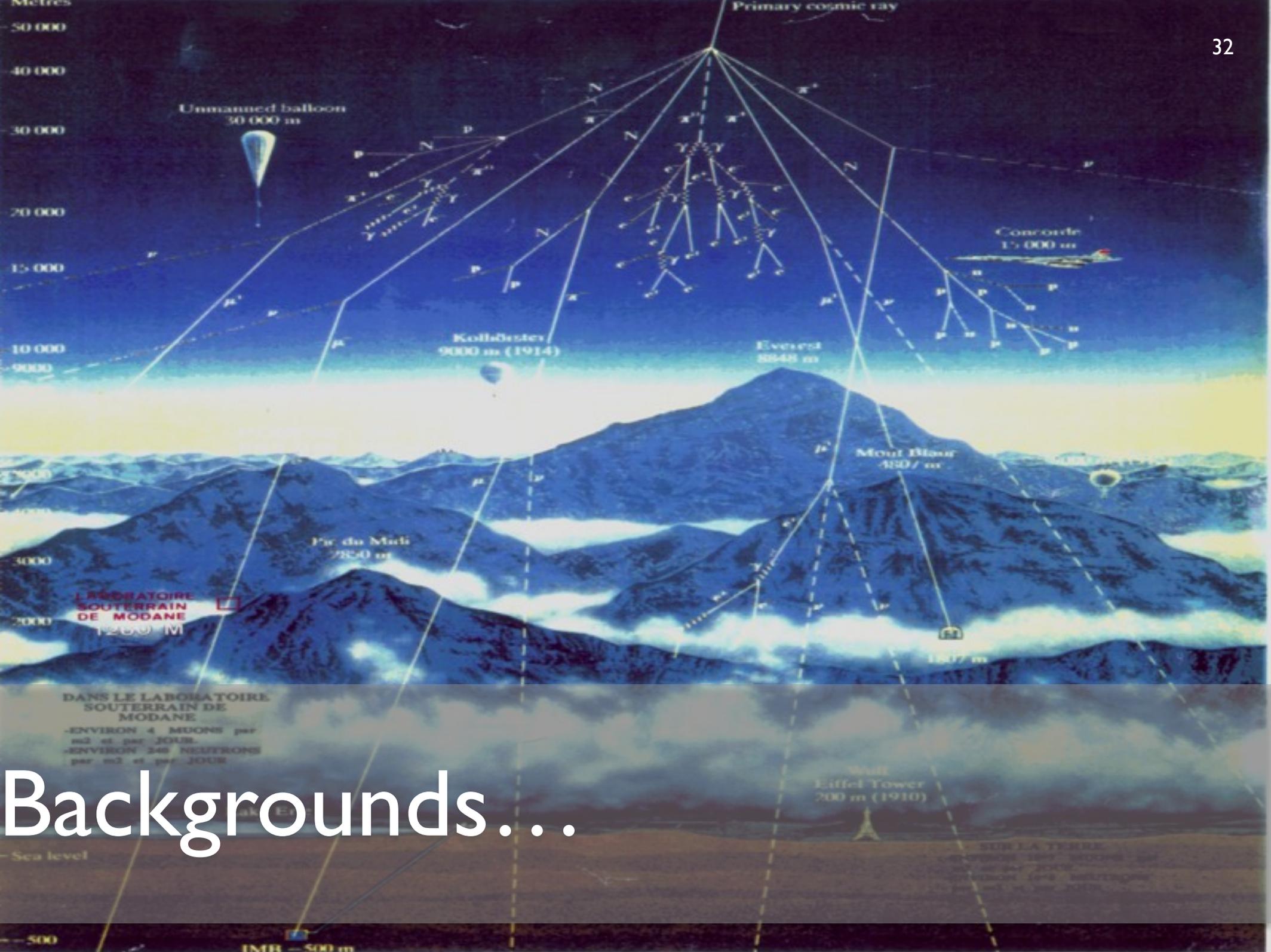
Discovery @ 3σ if $\sin^2(2\theta_{13}) > 0.05$

A close-up photograph of an elephant's head, focusing on its eye and the surrounding wrinkled, brownish-tan skin. The skin has a highly textured, cracked appearance. The eye is dark and partially obscured by the folds of the skin. A semi-transparent dark grey horizontal bar is overlaid at the bottom of the image, containing the text "how do we do that?".

how do we do that?

Anatael Cabrera (CNRS-IN2P3 & APC)

- **flux uncertainty (~2%)** goes negligible \Rightarrow near Detector
- **S/BG > 50** \Rightarrow huge statistical power (two most reactors in the world)
 - fewer reactors \Rightarrow maybe “reactor-off”
 - large or many detectors? S/B often scales with radius (Volume/Surface)
- **reduce & understand backgrounds** in situ (even if no “reactor-off”)
 - optimal detector overburden, design, shielding, radio-purity, etc...
- **reduce & understand experimental systematics**: design & calibration
 - inter-detector normalisation: $\leq 0.6\%$
 - inter-detector energy calibration: $\leq 2\%$



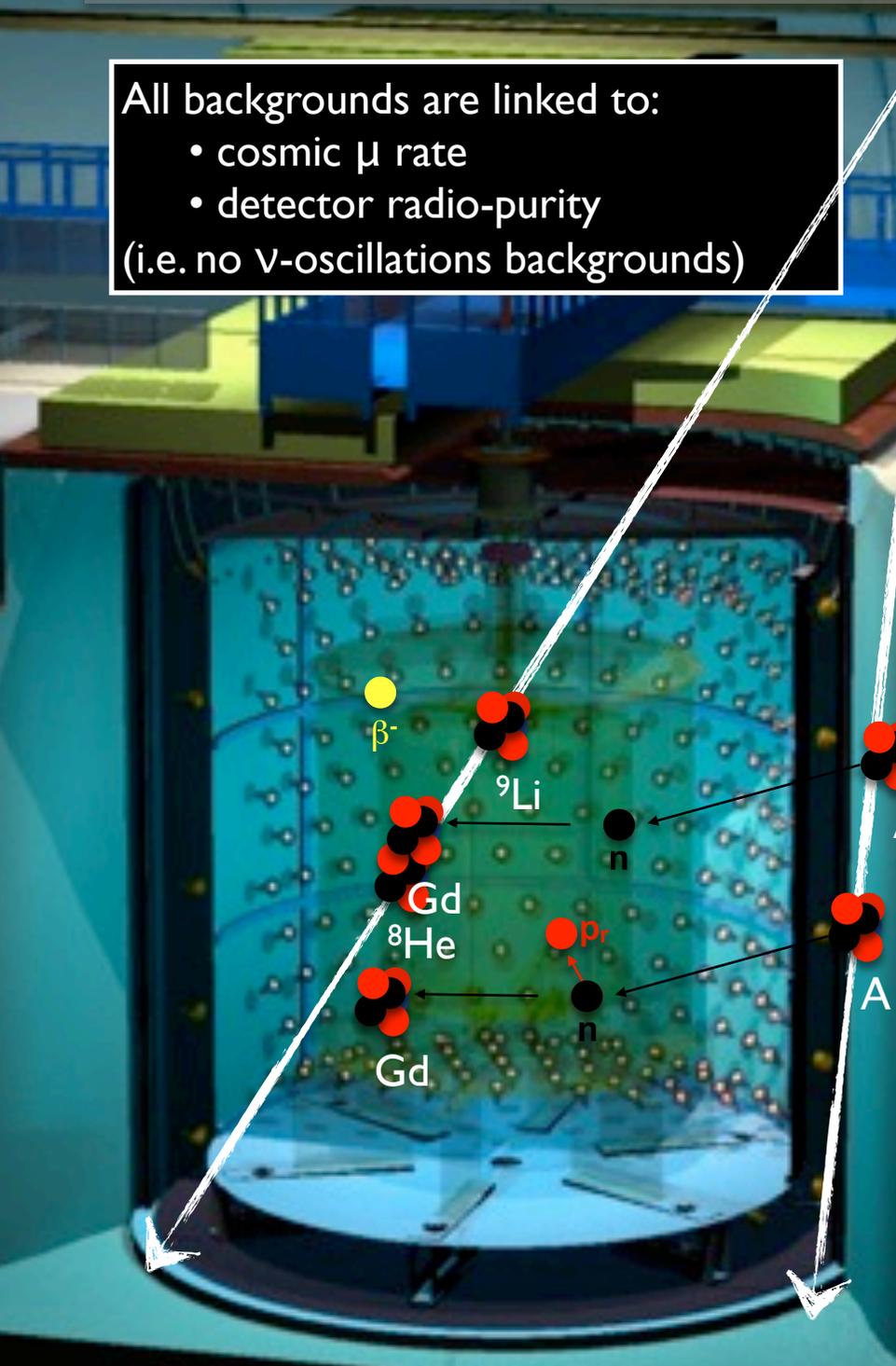
Backgrounds...

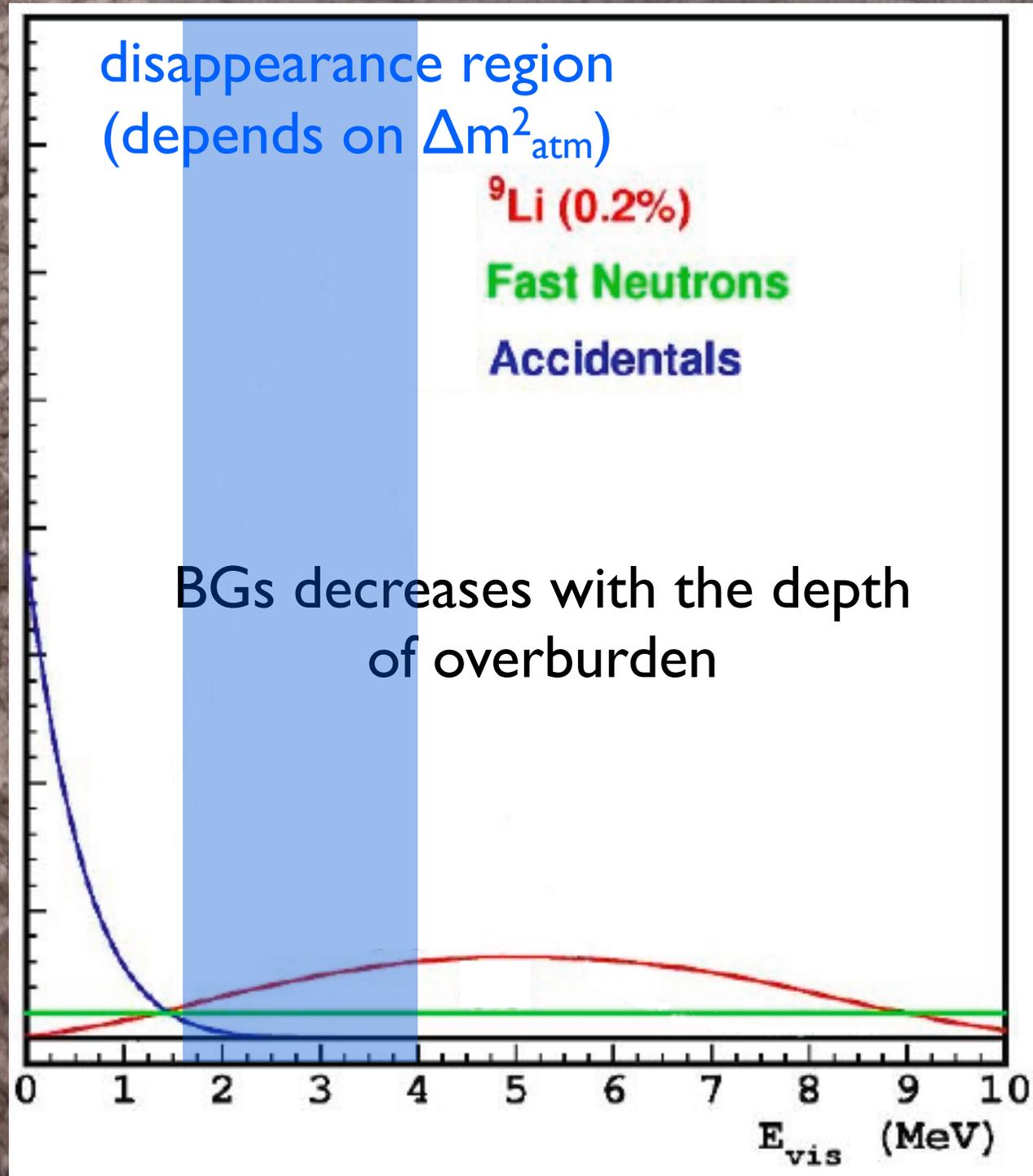
cosmic- μ

All backgrounds are linked to:

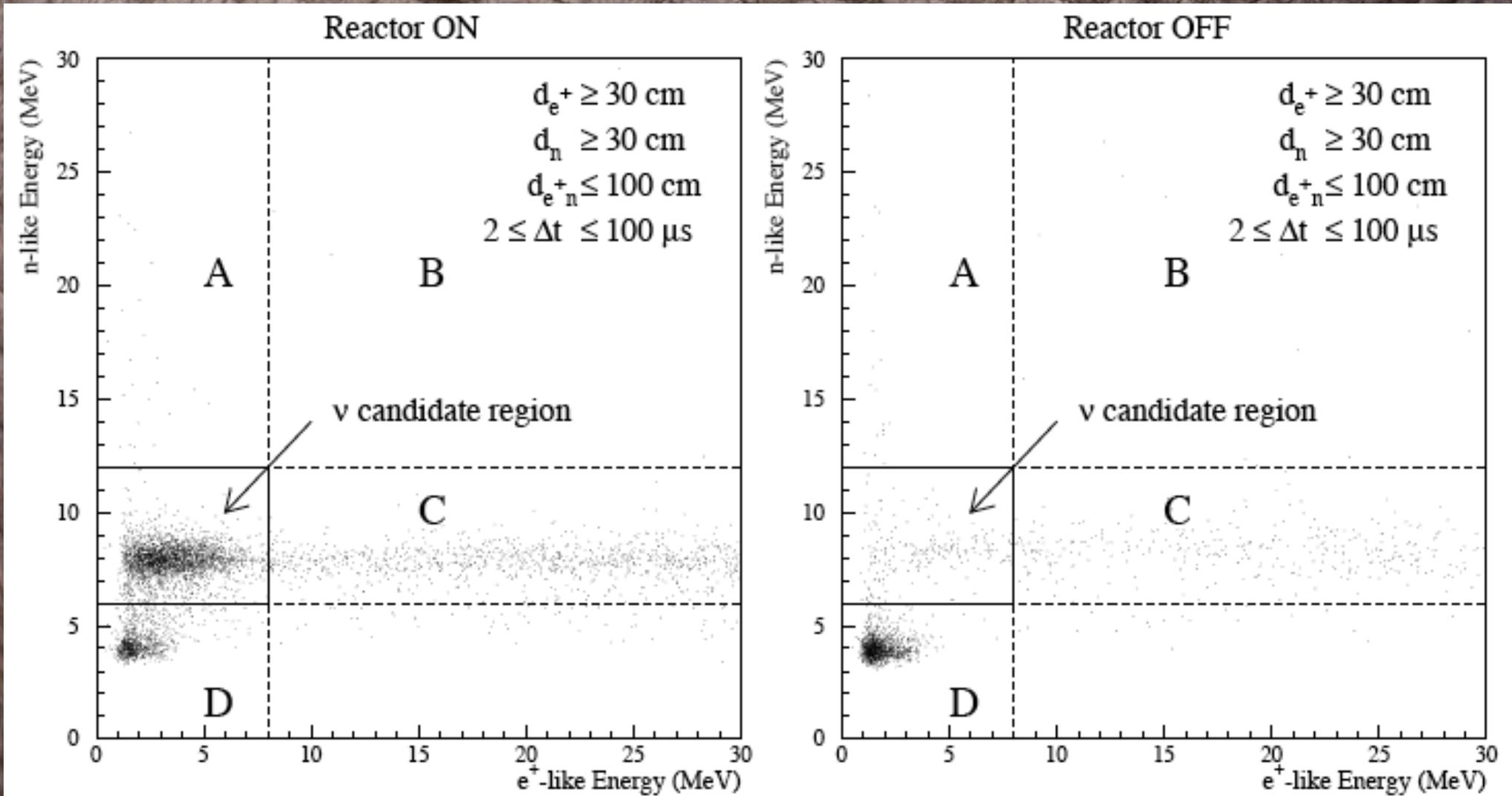
- cosmic μ rate
- detector radio-purity (i.e. no ν -oscillations backgrounds)

- BGs are reduced (wrt CHOOZ) \Rightarrow **detector design** & measure in situ
- μ tagging and detection \Rightarrow **IV, OV detectors** [see later] and **dedicated electronics**
- **Correlated BGs**
 - **β -n Isotopes** (half-life time ~ 100 ms) \Rightarrow impossible to veto (time/location)
 - **Fast-n's** p-recoil & Gd capture
- **Accidental BGs**
 - radio-active β^- (rate @ ID < 10 Bq)
 - radio-purity campaigns & all installation under cleaning environ.
 - fast-n thermalises and captures on Gd





switched off signal \Rightarrow measure background "naked"!



Apollonio et al (CHOOZ): hep-ex/0301017

more reactors \Rightarrow less likely!

FD \Rightarrow much input from CHOOZ experiment
 ND \Rightarrow extrapolation from FD/CHOOZ to new site

Table 12: Summary of the background subtraction error at the Far and Near detector (preliminary). Background rate and shape with their corresponding uncertainties are used for the calculation of the sensitivity. The systematics correspond to our best estimate of the error associated with each particular background (this can be used as a “background systematic error”).

Detector	Site		Background				
			Accidental Materials	PMTs	Fast n	Correlated μ -Capture	^9Li
CHOOZ (24 ν /d)	Far	Rate (d^{-1})	—	—	—	—	0.6 ± 0.4
		Rate (d^{-1})	0.42 ± 0.05		1.01 ± 0.04	± 0.1	$(stat) \pm 0.1$
		bkg/ ν	1.6%			4%	
		Systematics	0.2%			0.4%	
Double Chooz (69 ν /d)	Far	Rate (d^{-1})	1 ± 0.1	1 ± 0.1	0.15 ± 0.15	0.42 ± 0.2	1 ± 0.5
		bkg/ ν	1.4%	1.4%	0.2%	0.6%	1.4%
		Systematics	0.2%	0.2%	0.2%	0.3%	0.7%
Double Chooz (990 ν /d)	Near	Rate (d^{-1})	7.2 ± 1.0	7.2 ± 1.0	1.4 ± 0.14	2.6 ± 1.2	5.2 ± 3.2
		bkg/ ν	0.7%	0.7%	0.14%	0.26%	0.6%
		Systematics	0.1%	0.1%	0.2%	0.1%	0.3%

DC Proposal: hep-ex/0606025

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Error Source	Error Type	Error Description	CHOOZ Absolute	DC Absolute	DC Relative	
Reactor	Reactor					
		Production Cross Section	1.90%	1.90%		
		Core Powers	0.70%	0.70%		
		Energy per Fission	0.60%	0.60%		
Detector	Detector					
	Free H in TG	Detection Cross Section	0.30%	0.10%		
		Volume	0.30%	0.20%	0.20%	
		Fiducial Volume	0.20%	0.20%		
		Density		0.10%	0.01%	
	H/C (Chemical Composition)	0.80%	0.80%	0.10%		
Analysis	Electronics	Dead Time	0.25%		0.00%	
	Analysis					
	Positron	Particle Id				
		Escape		0.10%		
		Capture		0.00%		
	Neutron	Identification Cut		0.80%	0.10%	0.10%
		Escape		1.00%		
		Capture (% Gd)		0.85%	0.30%	0.30%
	Anti-neutrino	Identification Cut		0.40%	0.10%	0.10%
		Time Cut		0.40%	0.10%	0.10%
Distance Cut			0.30%			
Unicity (neutron multiplicity)			0.50%			
	Efficiency uncert due to bkg					
Total			2.90%	2.31%	0.46%	

Upgraded to expected ND location

N2P3 & APC)

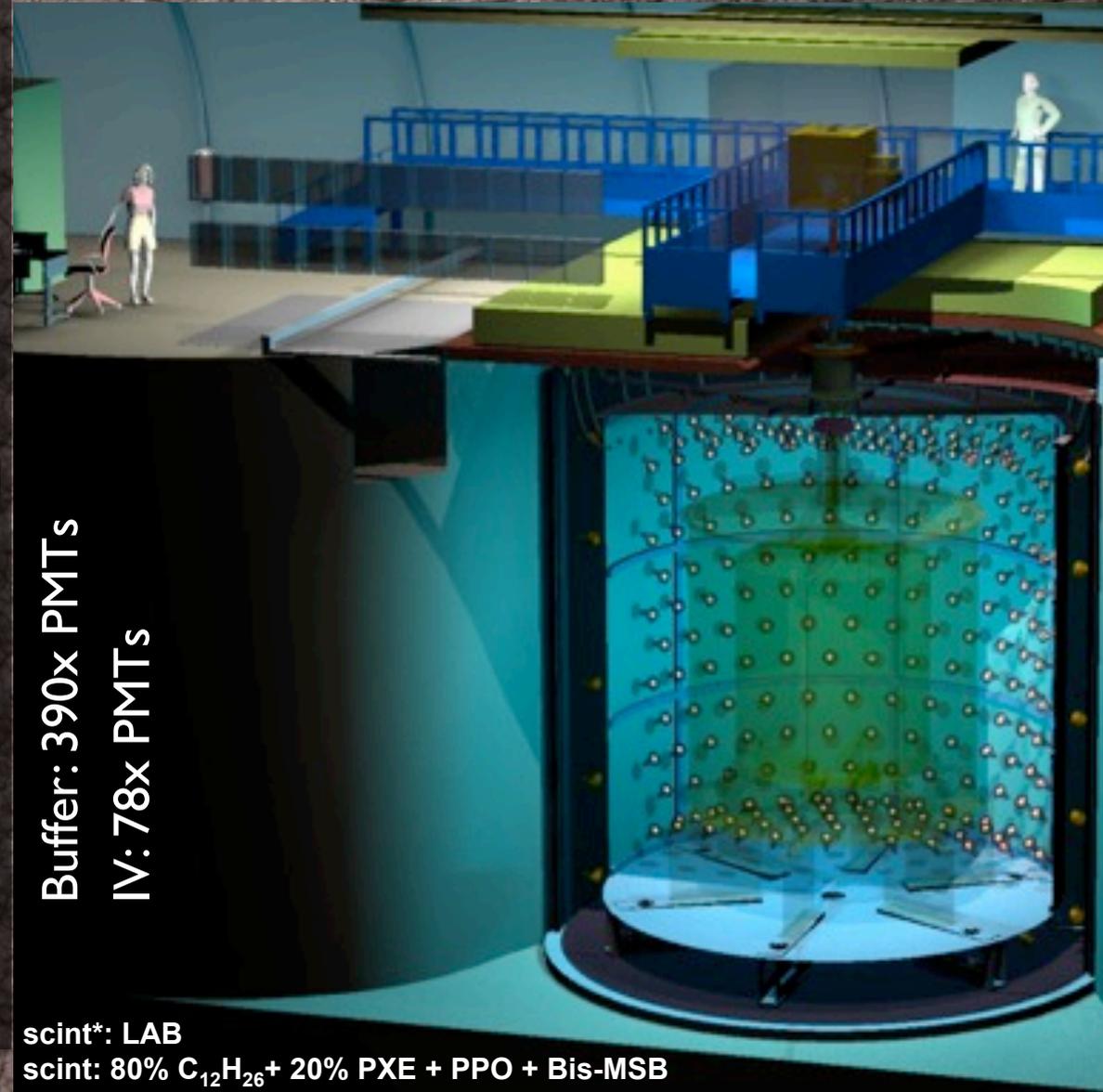
A person wearing a white protective suit and a white hood is working on a large array of circular detector modules. The modules are arranged in a grid pattern on a metal structure. The person is pointing at one of the modules. The background is a yellowish wall with more modules.

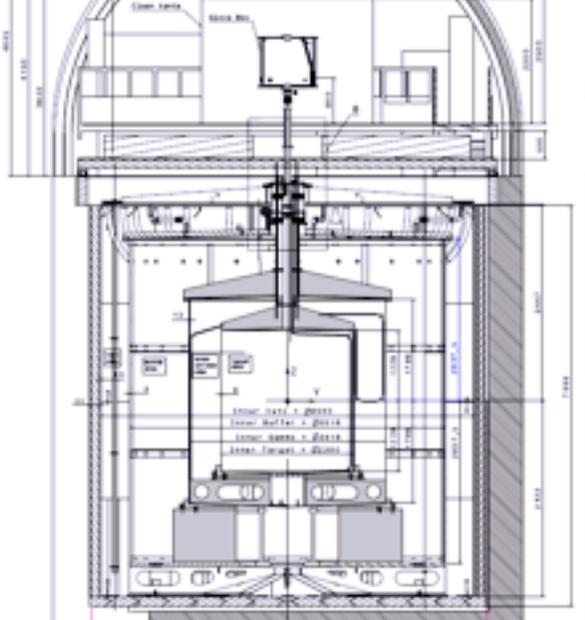
the detector: a θ_{13} -LAND

- **Pit:** 7m x 7m (FD: CHOOZ lab) \Rightarrow max. fiducial volume cylinder
- **Inner-Detector**
 - **Target:** acrylics + scint & 0.1%Gd \Rightarrow n-Gd interaction region
 - **γ -Catcher:** acrylics + scint \Rightarrow extra calorimetry containment
 - **Buffer:** oil no scint \Rightarrow isolation
- **Inner-Veto:** scint* \Rightarrow tagged μ s and fast-n
- **Outer-Veto:** scint-strips (a la MINOS) \Rightarrow tagged near-by μ s
- **γ -Shield:** 15cm steel \Rightarrow reduce rock- γ s (singles)
- **Glove-Box** \Rightarrow calibration apparatus contamination-less

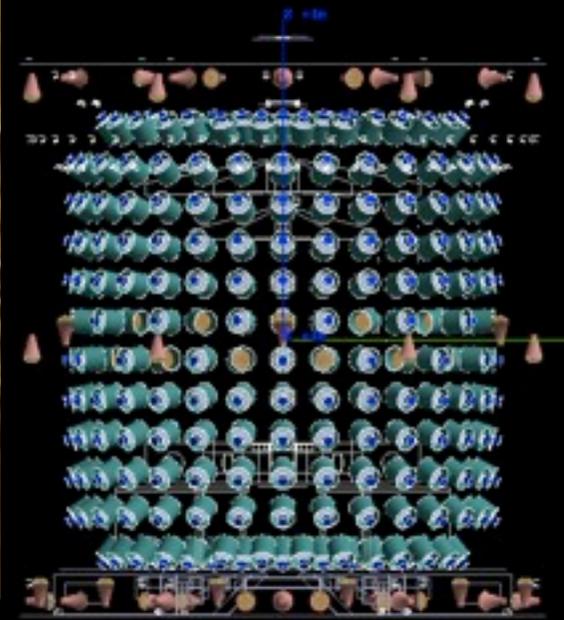
θ_{13} -LAND must...

- * inter-detector comparison systematic $< 1\%$
- * radio-purity & material compatibility

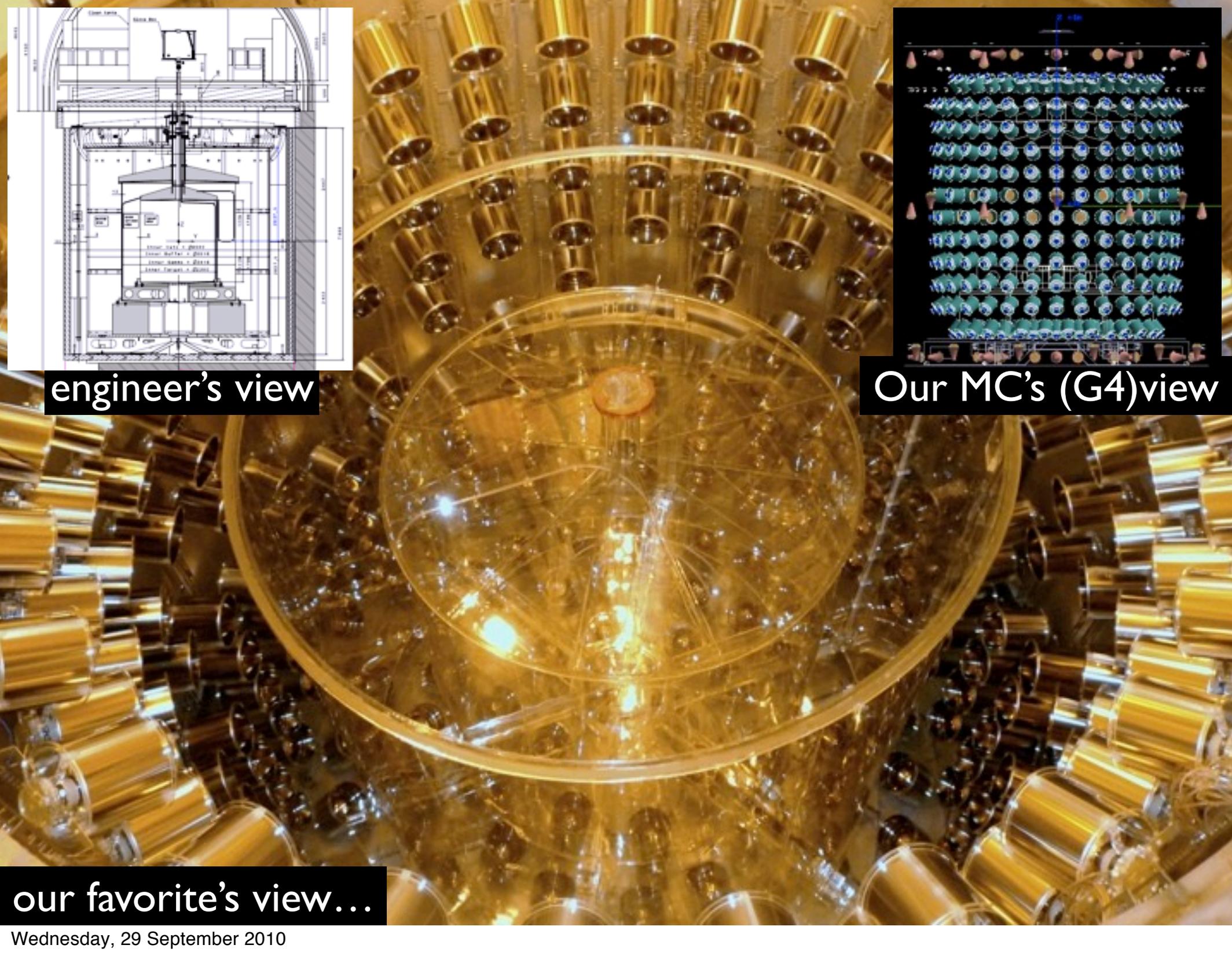




engineer's view



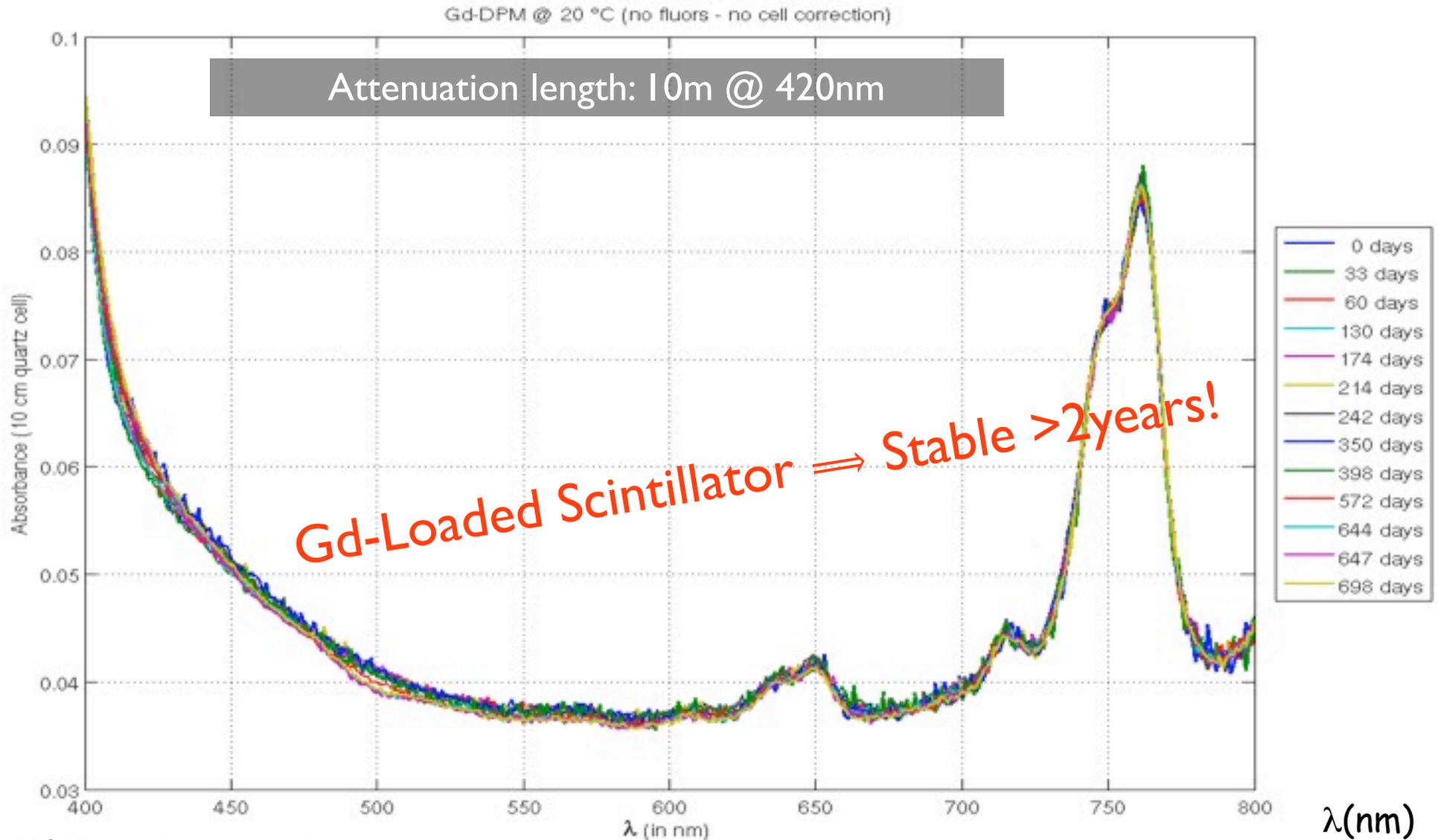
Our MC's (G4) view



our favorite's view...

Gd doped liquid scintillator

Liquid Scintillator: 80% Dodecane + PXE 20% + 0.1%Gd



UV-VIS-IR scintillator transmission



readout & online

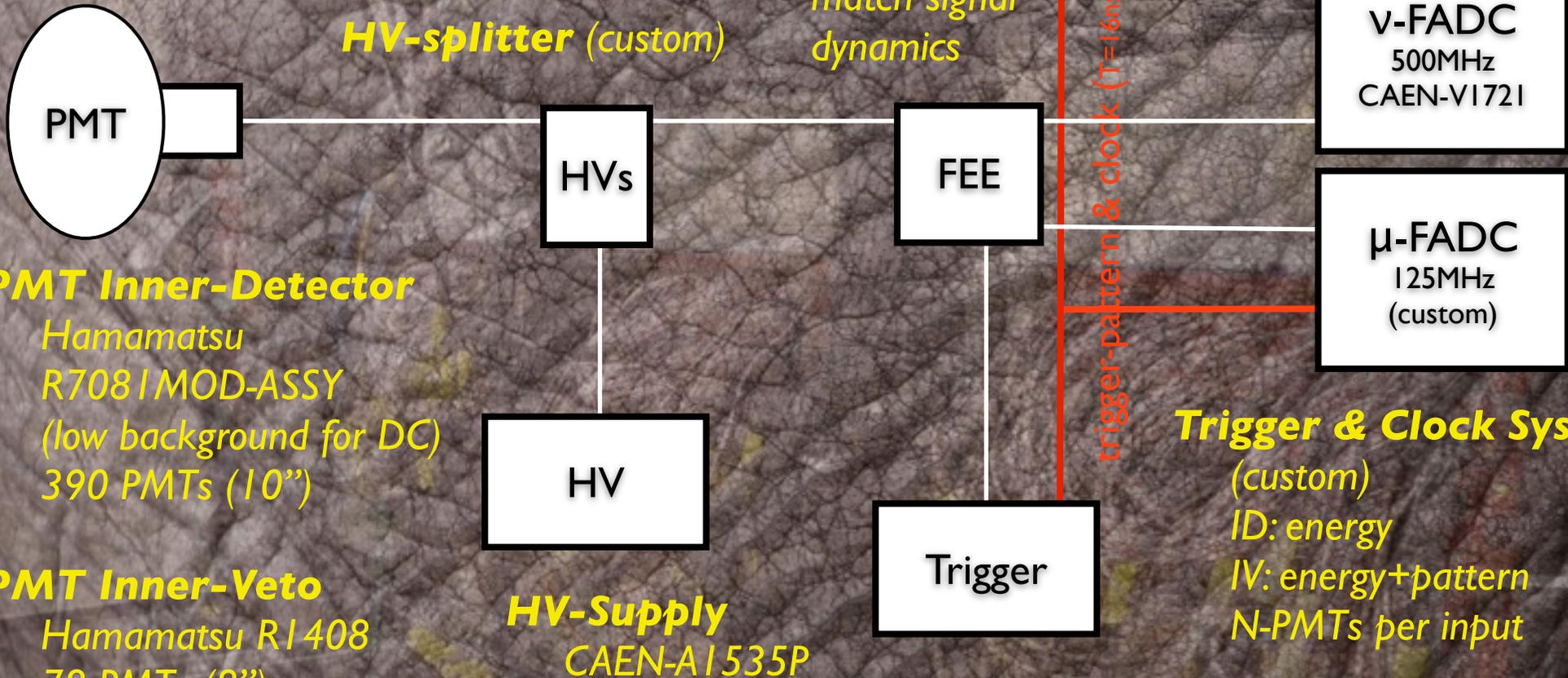
Trigger Rate <300Hz (cosmic- μ s @ ND)

VME-Bus

DAQ software in Ada

HV-splitter (custom)

FEE (custom)
match signal
dynamics



PMT Inner-Detector

Hamamatsu
R7081MOD-ASSY
(low background for DC)
390 PMTs (10")

PMT Inner-Veto

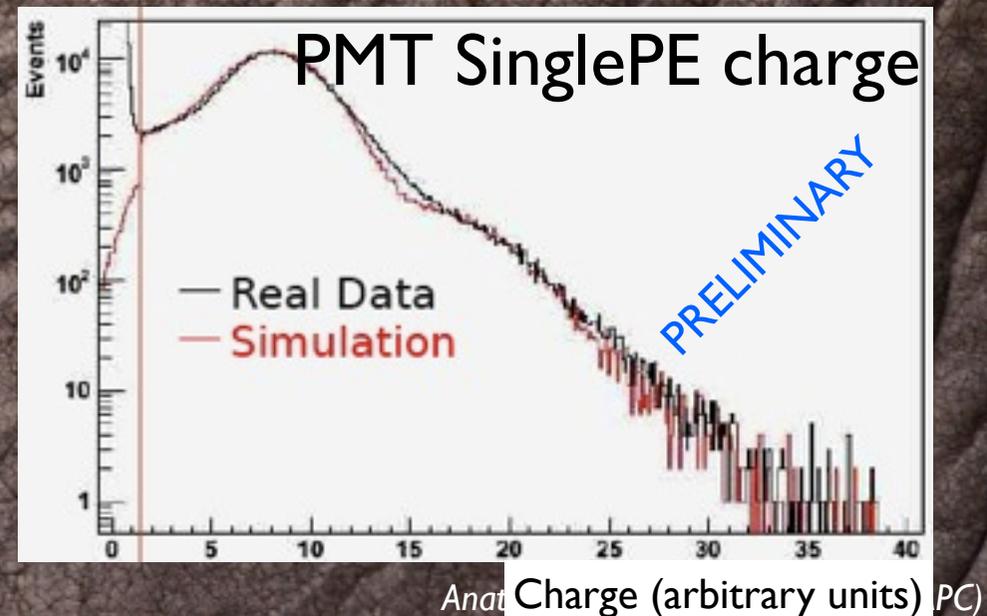
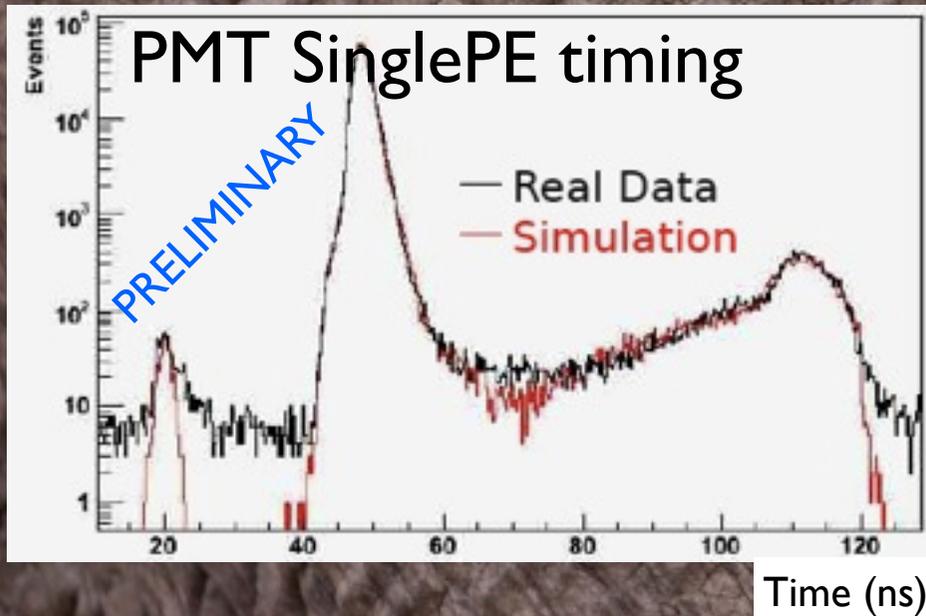
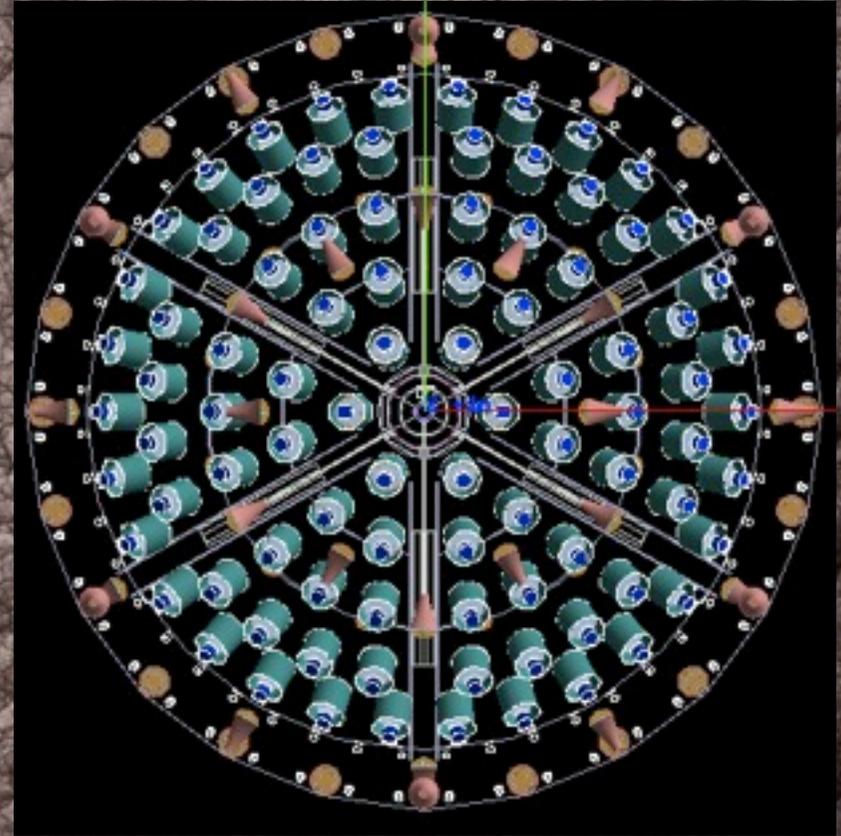
Hamamatsu RI408
78 PMTs (8")
(from IMB)

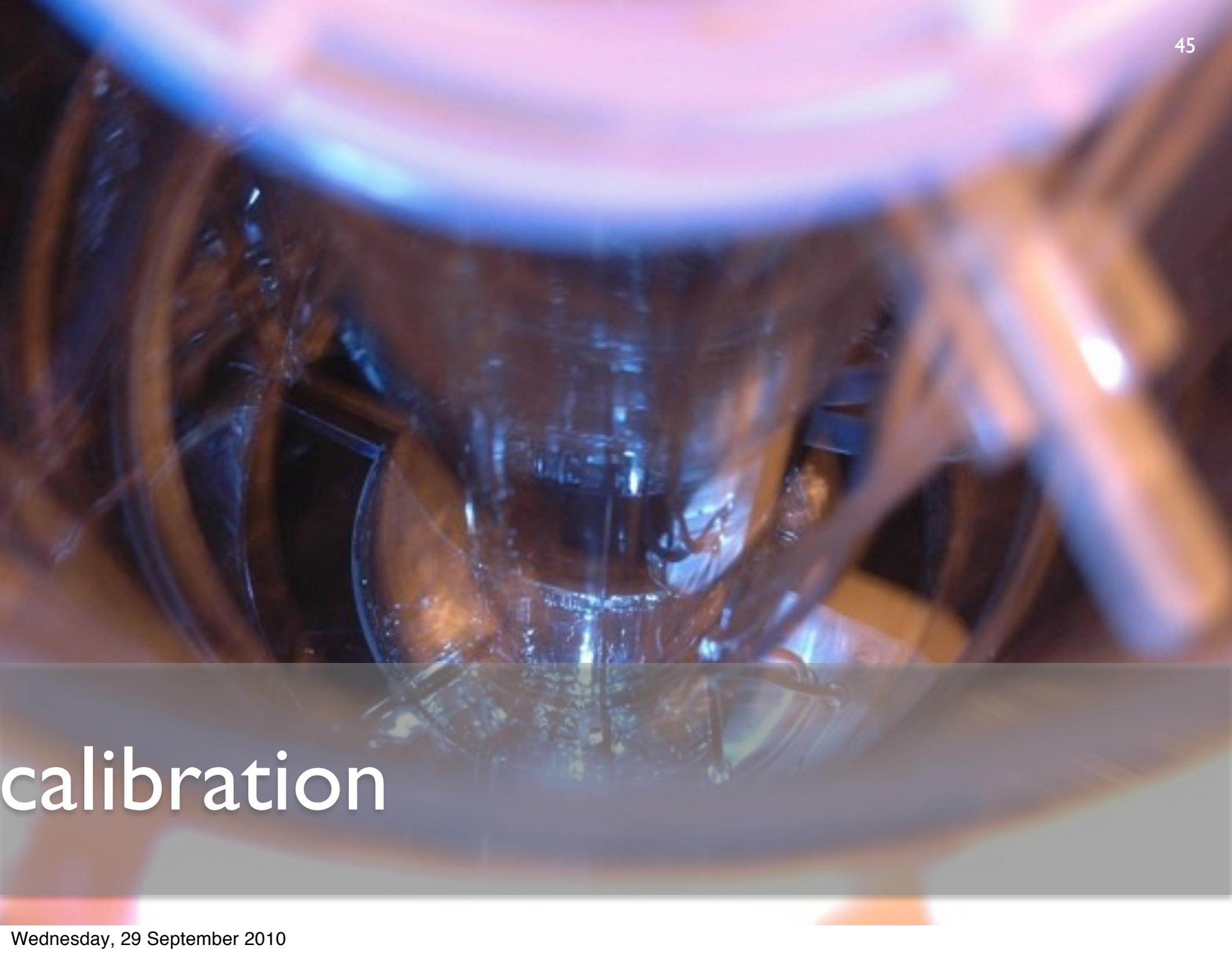
online/electronics integration benches...
"Vertical Slice" (1PMT) @ APC
"Super-Vertical Slice" (30 PMTs) @ MPIK

OV readout a la OPERA (Hamamatsu M64 + Maroc2-chip)

Anatael Cabrera (CNRS-IN2P3 & APC)

- Complete detector simulation...
- Physics: Geant4
- Optical Interface: Geant4
 - based on “GLG4sim”
- dedicated detector simulation (optical interface, PMT, electronics, trigger) tuned to test-bench data ⇒ **understand/evaluate systematics**





calibration

- **Redundancy** \Rightarrow cross-check & understand inter-detector systematics
- **light sources (embedded, i.e. non-intrusive):** LEDs in ID and IV
 - monitor stability of readout (timing, gain) and scintillator
- **light sources (deployed):** LED, red-laser & UV-laser
 - PM gain, timing, scintillator stability & attenuation
- **radioactive source:** across most energy scale
 - H-capture, Cs¹³⁷, Na²², K⁴⁰, Co⁶⁰, etc
- **n-sources:** n capture on Gd (study efficiencies)
 - Cf²⁵²(untagged) & AmBe (light tagged) \Rightarrow 3D deployable
- **3D calibration strategy:** map full detector response
 - along z-axis, articulated arm (off z-axis), GC & Buffer tubes
- 2 detectors \Rightarrow calibration source absolute knowledge less important
 - same source response comparison ND and FD (cancel some systematics)



A person wearing a white protective suit and gloves is working on a large piece of equipment. The person is leaning over the equipment, and their hands are visible as they work. The equipment has several vertical metal rods. The background is dark, and the lighting is focused on the person and the equipment.

far detector status:
constructed \Rightarrow now to be filled!



the neutrino source: 2 of the most powerful reactors in the world



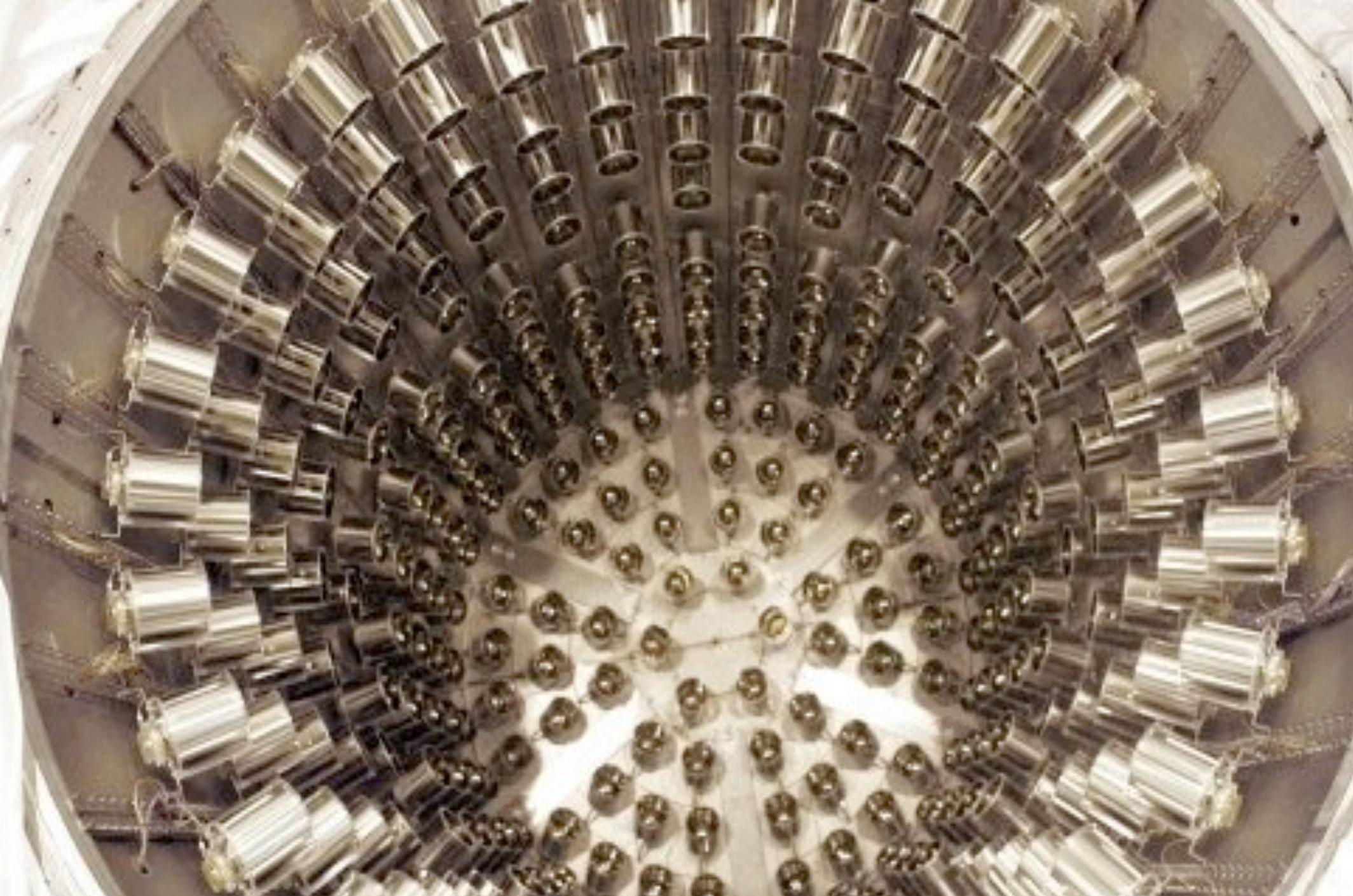
FD underground lab access: 300m walk @ a small angle



Inner-Veto detector (scintillator) \Rightarrow tag μ s & possible fast-neutrons



Inner-Detector PMT installation



Inner-Detector PMT system (15% photocathode coverage & μ -metal)



Inner-Detector Acrylics (installation) \Rightarrow our fiducial volumes!



final mechanical tuning of the chimney (hyper-delicate interface)

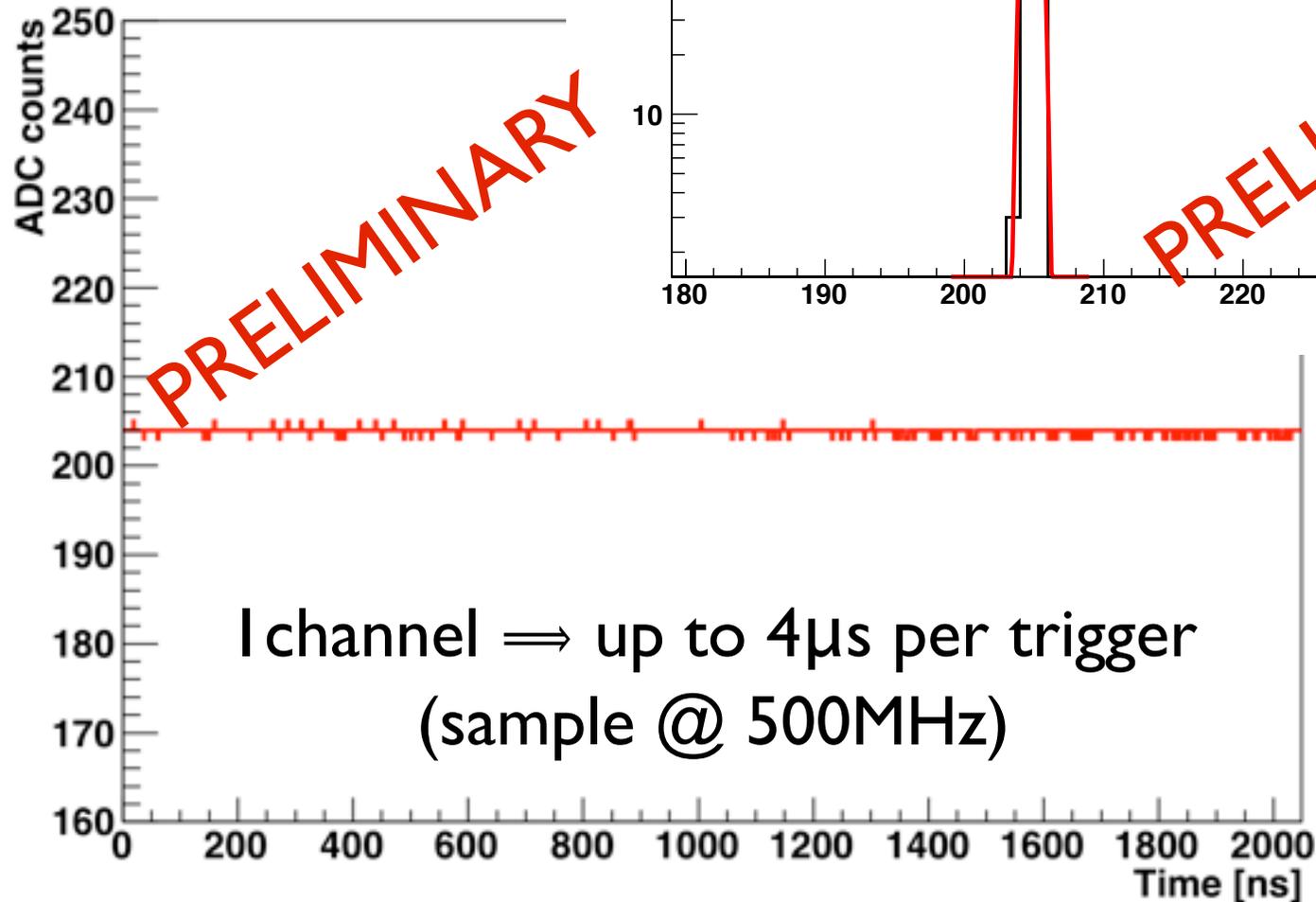


FD “pre-commissioning” ...

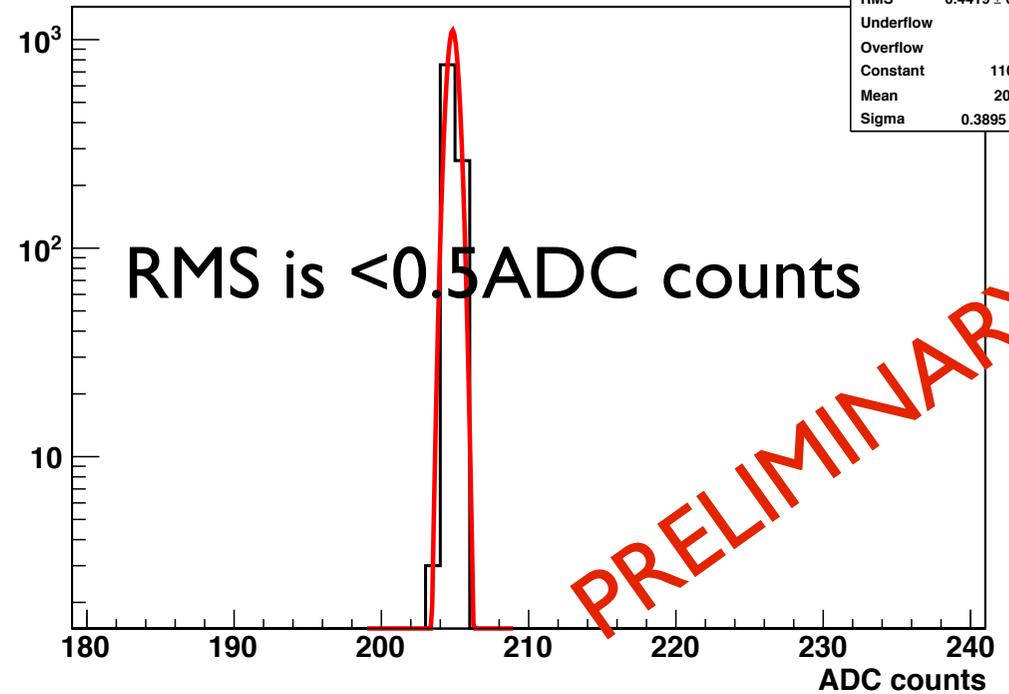
checking electronics noise with FADC...

Measured @
Chooz

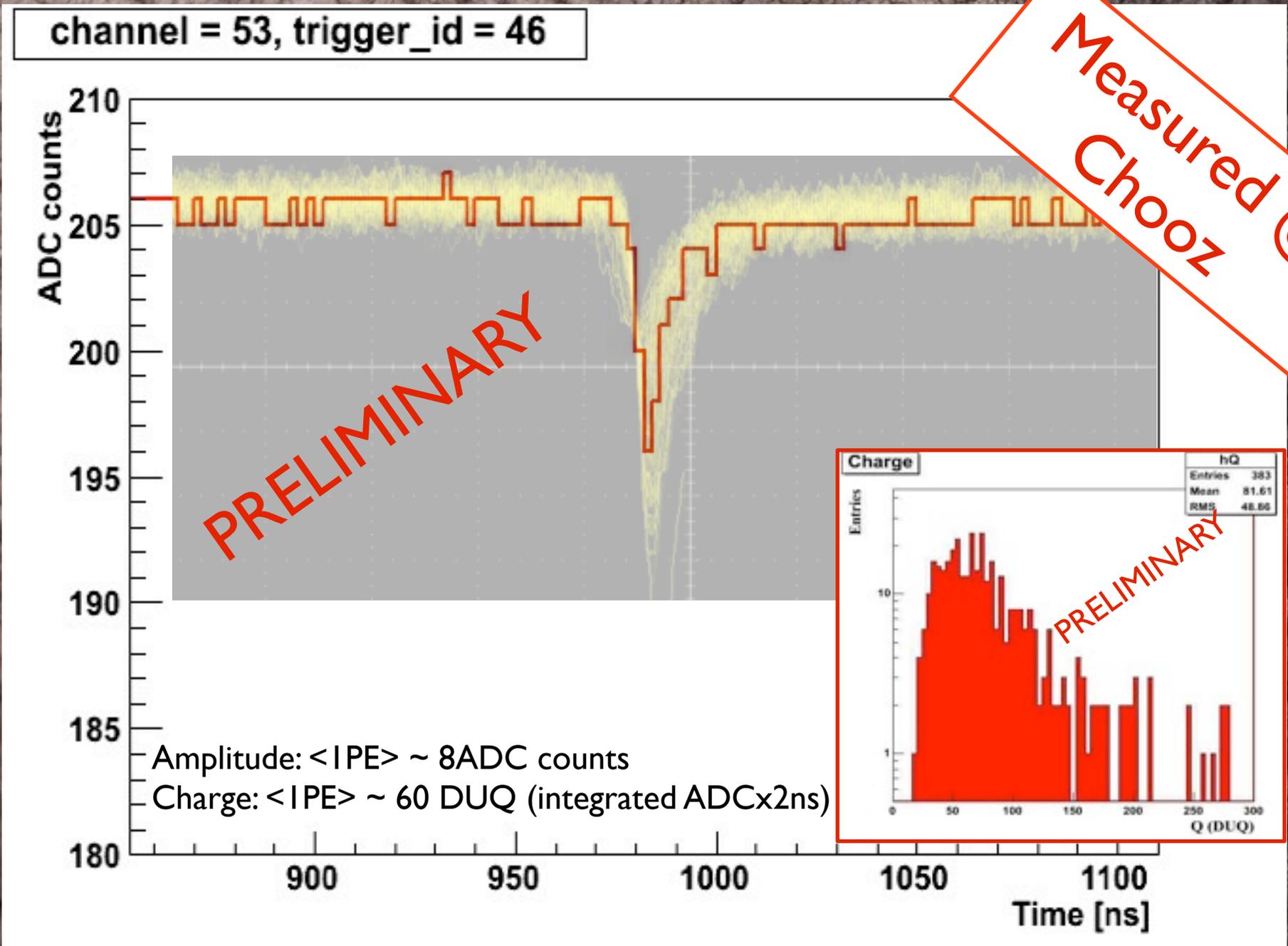
Pedestal



Pedestal Ch 0, TriggerID 1



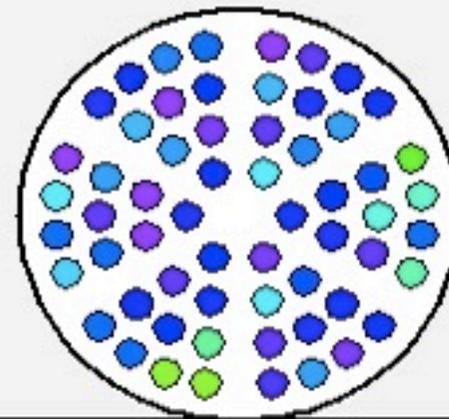
the first single-PEs detected @ FD...



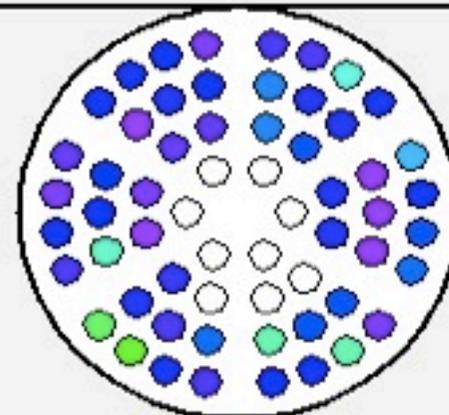
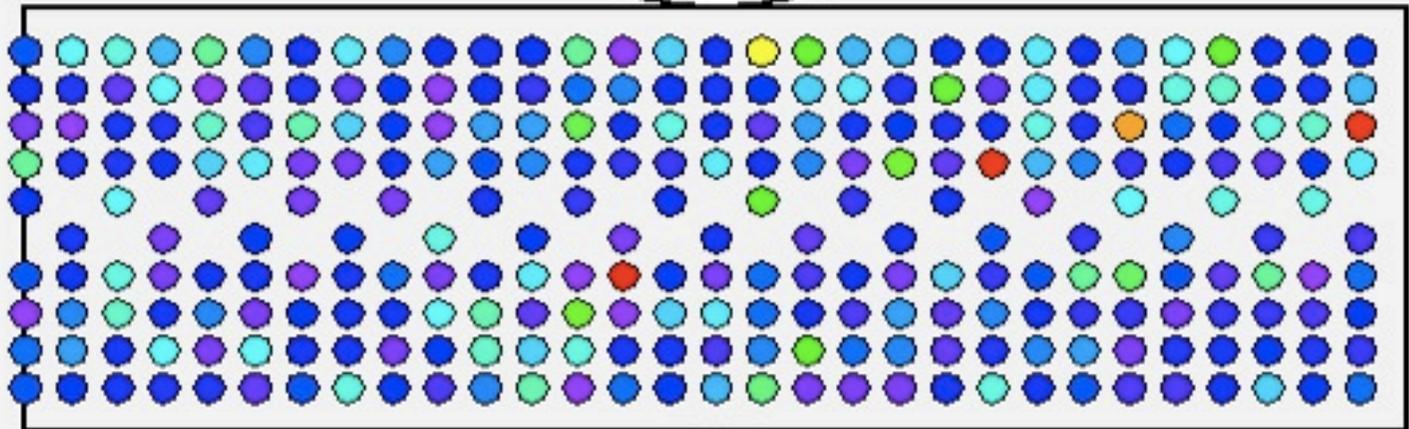
first internal triggers...

First Energy-Trigger
⇒ large energy deposition
over ~70ns coincidence
without scintillator

No Mapping
⇒ no spatial
reconstruction then



PRELIMINARY



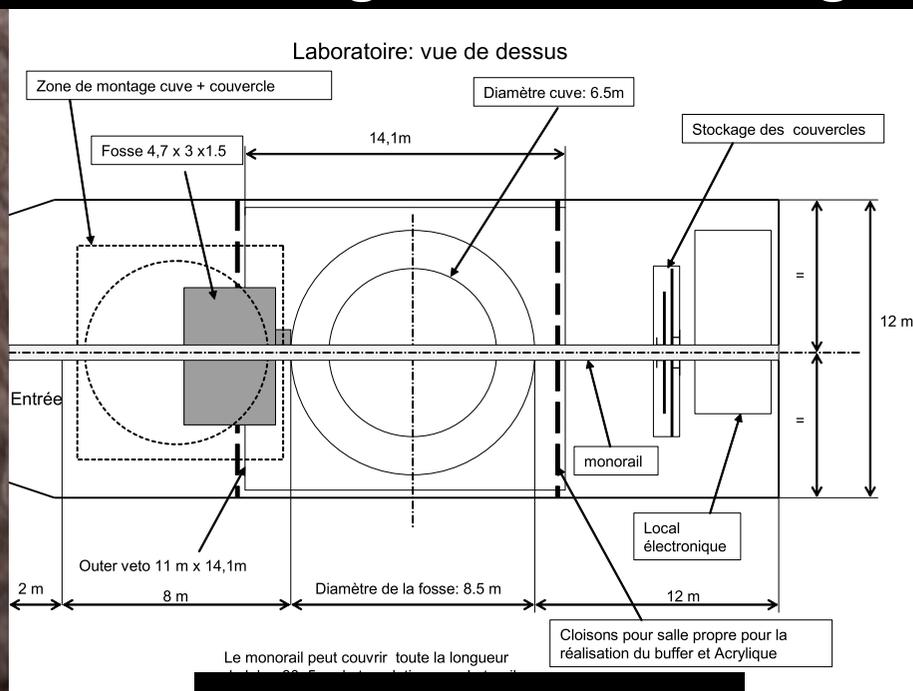
Most electronics up & running! (IFADC card missing)



near detector status...

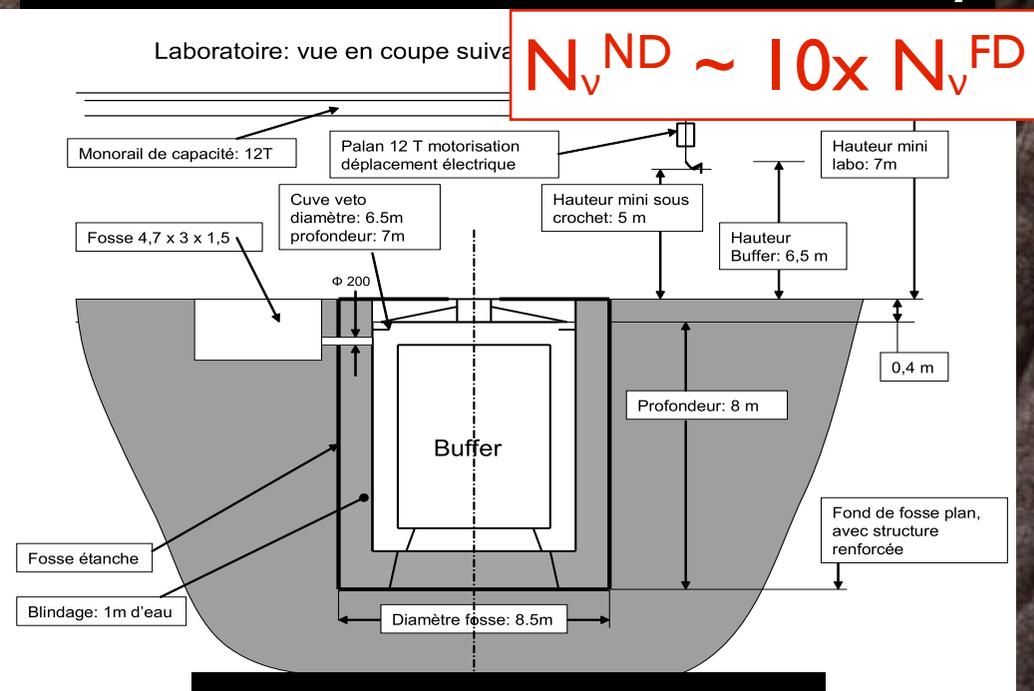
- **all authorisations & funding secured...**
 - ground studies completed since end of 2009
 - involvement of EDF on the construction
- tendering over summer \Rightarrow **digging from end of 2010**
- **lab available by end of 2011**
 - ND construction \Rightarrow **30 weeks** (many lessons learned from FD)
 - goal: **data taking by towards 2012**

ND Lab designed for building ND



3x zones concept

1m water shield to radioactivity



easier filling concept (IN2P3 & APC)

A close-up photograph of an elephant's head, focusing on its eye and the surrounding wrinkled, textured skin. The skin is a mix of brown and tan tones, with deep creases and ridges. The eye is dark and partially obscured by the folds of the skin.

what to remember?

Anatael Cabrera (CNRS-IN2P3 & APC)

- **neutrino oscillations framework is not exhausted from physics** (but harder)...
 - the “gate” to understand leptonic flavour sector $\Rightarrow \theta_{13}$ critical
 - compelling research programme: Double Chooz, Daya Bay, RENO
- **we must measure θ_{13} ...**
 - SM leptonic mixing parameter \Rightarrow it must be measured
 - even a zero-value is interesting (less fortunate for experimentalist)
- **Double Chooz...**
 - DC has **led much of the θ_{13} -LAND business** “*on the shoulders of giants*” (CHOOZ, Borexino, KamLAND, SNO, etc). Its legacy is a reality...
 - DC input on θ_{13} (measurement/limit) will deepen our insight on a global perspective of the (light) neutrino flavour sector \Rightarrow **critical synergies with beams (MINOS, T2K, NOvA, etc) and Solar/KamLAND results**

- **Double Chooz FD is about to start data taking...**
 - FD construction \Rightarrow **finished!**
 - FD first lights \Rightarrow **since June** (“dry” detector)
 - FD filling \Rightarrow **ready to start...**
 - FD first scintillation lights \Rightarrow during filling data-taking (October)
 - FD commissioning \Rightarrow **hoping to start by late December!**
 - FD publication on θ_{13} \Rightarrow soon next year (our goal)!
- DC can obtain **CHOOZ** worth of **signal** data \sim 2 months of running
 - **$\sin^2(2\theta_{13}) \leq 0.054$ @ 90%CL** with FD only (about 1.5years of data)
 - **$\sin^2(2\theta_{13}) \leq 0.030$ @ 90%CL** with FD & ND (about 3years of data)
- Near Detector digging end of 2010 \Rightarrow **running by end of 2012!**

Anatael Cabrera (CNRS/IN2P3)

APC (Paris)

anatael@in2p3.fr

[else, just google “anatael”]

thank you...

Anatael Cabrera (CNRS-IN2P3 & APC)

Wednesday, 29 September 2010